



Michael Buballa

TU Darmstadt

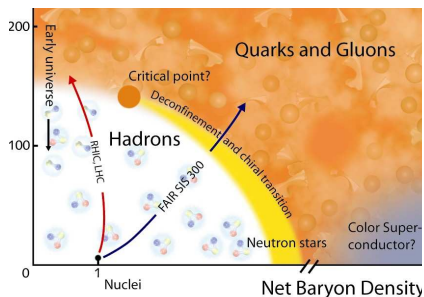
HFHF Theory Retreat

Castiglione della Pescaia, Italy, September 12 - 16, 2022



**Focus:** Moderate temperature and (not asymptotically) high density

- ▶ **theoretically hard:**
  - ▶ non-perturbative
  - ▶ sign problem on the lattice
- ▶ **phenomenologically interesting:**
  - ▶ neutron stars and neutron-star mergers
  - ▶ CBM physics at FAIR
- ▶ **regions of special interest:**
  - ▶ critical point
  - ▶ color superconducting phases

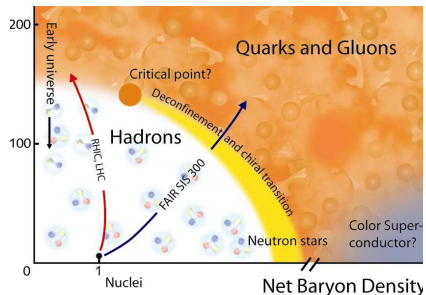


**Aim of this talk:**

Pedagogical introduction, laying the ground for Lennart's and Hosein's talks.

**Focus:** Moderate temperature and (not asymptotically) high density

- ▶ **theoretically hard:**
  - ▶ non-perturbative
  - ▶ sign problem on the lattice
- ▶ **phenomenologically interesting:**
  - ▶ neutron stars and neutron-star mergers
  - ▶ CBM physics at FAIR
- ▶ **regions of special interest:**
  - ▶ inhomogeneous chiral phases
  - ▶ color superconducting phases



**Aim of this talk:**

Pedagogical introduction, laying the ground for Lennart's and Hosein's talks.



# COLOR SUPERCONDUCTIVITY

---

# Why (color) superconductivity? - Cooper instabilities

---

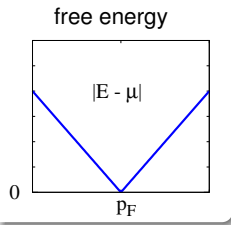


TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

# Why (color) superconductivity?

## - Cooper instabilities

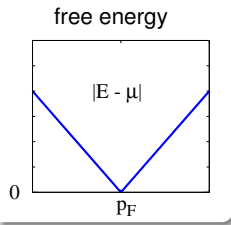
- ▶ Noninteracting fermions at  $T = 0$ :
  - ▶ Particles at the Fermi surface can be created at the Fermi surface with no free-energy cost.



# Why (color) superconductivity?

## - Cooper instabilities

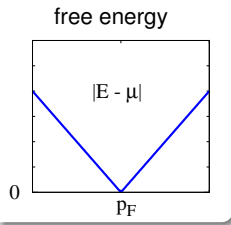
- ▶ Noninteracting fermions at  $T = 0$ :
  - ▶ Particles at the Fermi surface can be created at the Fermi surface with no free-energy cost.
- ▶ Add (arbitrarily small) attraction:
  - ▶ Fermi surface unstable against pair creation (“Cooper pairs”)



# Why (color) superconductivity?

## - Cooper instabilities

- ▶ Noninteracting fermions at  $T = 0$ :
  - ▶ Particles at the Fermi surface can be created at the Fermi surface with no free-energy cost.
- ▶ Add (arbitrarily small) attraction:
  - ▶ Fermi surface unstable against pair creation (“Cooper pairs”)
  - ▶ Bose condensation of the Cooper pairs

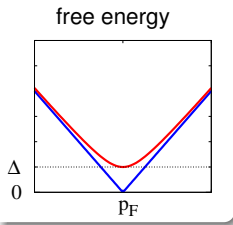




# Why (color) superconductivity?

## - Cooper instabilities

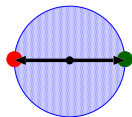
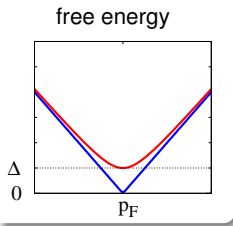
- ▶ Noninteracting fermions at  $T = 0$ :
  - ▶ Particles at the Fermi surface can be created at the Fermi surface with no free-energy cost.
- ▶ Add (arbitrarily small) attraction:
  - ▶ Fermi surface unstable against pair creation ("Cooper pairs")
  - ▶ Bose condensation of the Cooper pairs
    - rearrangement of the Fermi surface
    - gaps



# Why (color) superconductivity?

## - Cooper instabilities

- ▶ Noninteracting fermions at  $T = 0$ :
  - ▶ Particles at the Fermi surface can be created at the Fermi surface with no free-energy cost.
- ▶ Add (arbitrarily small) attraction:
  - ▶ Fermi surface unstable against pair creation (“Cooper pairs”)
  - ▶ Bose condensation of the Cooper pairs
    - rearrangement of the Fermi surface
    - **gaps**
- ▶ BCS pairing:
  - ▶ pairs with vanishing total momentum:  $\vec{p}^{(1)} = -\vec{p}^{(2)}$
  - ▶ each partner close to the Fermi surface
    - works only if  $p_F^{(1)} \approx p_F^{(2)}$



- ▶ QCD: attractive quark-quark interaction
  - diquark condensates:  $\langle q_i \mathcal{O}_{ij} q_j \rangle$

- ▶ QCD: attractive quark-quark interaction

→ diquark condensates:  $\langle q_i \mathcal{O}_{ij} q_j \rangle$

- ▶ Pauli principle:

$$\mathcal{O} = \mathcal{O}_{spin} \otimes \mathcal{O}_{color} \otimes \mathcal{O}_{flavor} = \text{totally antisymmetric}$$

- ▶ QCD: attractive quark-quark interaction

→ **diquark condensates:**  $\langle q_i \mathcal{O}_{ij} q_j \rangle$

- ▶ Pauli principle:

$$\mathcal{O} = \mathcal{O}_{spin} \otimes \mathcal{O}_{color} \otimes \mathcal{O}_{flavor} = \text{totally antisymmetric}$$

- ▶ most attractive channel:

- ▶ spin 0 (= antisymmetric)
- ▶ color  $\bar{3}$  (= antisymmetric)

→ antisymmetric in flavor

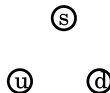
→ pairing between **different flavors**

- ▶ QCD: attractive quark-quark interaction
  - **diquark condensates:**  $\langle q_i \mathcal{O}_{ij} q_j \rangle$
- ▶ Pauli principle:
  - $\mathcal{O} = \mathcal{O}_{spin} \otimes \mathcal{O}_{color} \otimes \mathcal{O}_{flavor} = \text{totally antisymmetric}$
- ▶ most attractive channel:
  - ▶ spin 0 (= antisymmetric)
  - ▶ color  $\bar{3}$  (= antisymmetric)
  - antisymmetric in flavor
  - pairing between **different flavors**
- ▶ example:  $(\uparrow\downarrow - \downarrow\uparrow) \otimes (rg - gr) \otimes (ud - du)$

# Three-flavor systems

► Pairing patterns in flavor space:

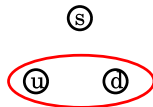
no pairing: “normal quark matter” (NQ)



► Pairing patterns in flavor space:

two-flavor superconducting (2SC) phase

(+ two analogous phases with  $us$  or  $ds$  pairing)

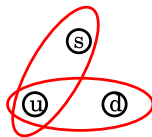




► Pairing patterns in flavor space:

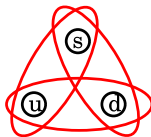
uSC phase

(similar: dSC phase, sSC)



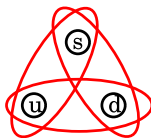
# Three-flavor systems

- ▶ Pairing patterns in flavor space:  
color-flavor locked (CFL) phase



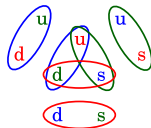
# Three-flavor systems

- ▶ Pairing patterns in flavor space:  
color-flavor locked (CFL) phase



- ▶ CFL pairing (more explicitly):

$$(\uparrow\downarrow - \downarrow\uparrow) \otimes \left( \begin{aligned} &(ud - du) \otimes (rg - gr) \\ &+ (ds - sd) \otimes (gb - bg) \\ &+ (su - us) \otimes (br - rb) \end{aligned} \right)$$



## (More) formal definition of the phases

### ► Diquark condensates:

$$(\uparrow\downarrow - \downarrow\uparrow) \otimes (ud - du) \otimes (rg - gr) \leftrightarrow \langle q^T C \gamma_5 \tau_2 \lambda_2 q \rangle \sim: \Delta_2$$

$$(\uparrow\downarrow - \downarrow\uparrow) \otimes (ds - sd) \otimes (gb - bg) \leftrightarrow \langle q^T C \gamma_5 \tau_5 \lambda_5 q \rangle \sim: \Delta_5$$

$$(\uparrow\downarrow - \downarrow\uparrow) \otimes (su - us) \otimes (br - rb) \leftrightarrow \langle q^T C \gamma_5 \tau_7 \lambda_7 q \rangle \sim: \Delta_7$$

$C = i\gamma^2\gamma^0$  charge conjugation matrix,  $C\gamma_5 \rightarrow J^P = 0^+$

$\tau_A$ : antisymmetric Gell-Mann matrices in flavor space

$\lambda_A$ : antisymmetric Gell-Mann matrices in color space

### ► Phases:

- NQ:  $\Delta_2 = \Delta_5 = \Delta_7 = 0$
- 2SC:  $\Delta_2 \neq 0, \Delta_5 = \Delta_7 = 0$
- CFL:  $\Delta_2 = \Delta_5 = \Delta_7 \neq 0$  (ideal case; realistic:  $\Delta_2 \approx \Delta_5 \approx \Delta_7 \neq 0$ )
- ...



$$\Delta_2 = \langle q^T C \gamma_5 \tau_2 \lambda_2 q \rangle$$

## ▶ gauge symmetries:

- ▶ **color:**  $q \rightarrow e^{i\theta_a \frac{\lambda_a}{2}} q$  blue quarks unpaired  $\Rightarrow SU(3)_c \rightarrow SU(2)_c$   
→ 5 of the 8 gluons get a nonzero **Meissner mass**.

- ▶ **electromagnetism:**  $q \rightarrow e^{i\alpha Q} q$ ,  $Q = \text{diag}_f(\frac{2}{3}, -\frac{1}{3})$  **broken**

But there is an **unbroken**  $U(1)$  gauge symmetry with charge  $\tilde{Q} = Q - \frac{1}{2\sqrt{3}} \lambda_8$ .

→ **color superconductor but not electromagnetic superconductor**

## ▶ global symmetries:

- ▶ **baryon number:**  $q \rightarrow e^{i\alpha} q \Rightarrow \Delta_2 \rightarrow e^{2i\alpha} \Delta_2$  **broken**

But there is an **unbroken** “modified baryon number”  $q \rightarrow e^{i\alpha(1-\sqrt{3}\lambda_8)} q$

- ▶  $SU(2)_L \times SU(2)_R$  **chiral symmetry:** **conserved**

→ **same global symmetries as 2-flavor restored phase, no Goldstone bosons**

# Symmetries of the (ideal) CFL phase

$$\langle q^T C \gamma_5 \tau_2 \lambda_2 q \rangle = \langle q^T C \gamma_5 \tau_2 \lambda_2 q \rangle = \langle q^T C \gamma_5 \tau_2 \lambda_2 q \rangle = \Delta$$

- ▶ **color:**  $SU(3)_c$  broken completely
- ▶ **chiral symmetry:**  $SU(3)_L \times SU(3)_R$  broken completely

but:

**residual  $SU(3)$**  under **combined color-flavor** rotations:  $q \rightarrow e^{i\theta_a(\tau_a - \lambda_a^T)} q$

→ “color-flavor locking”:  $SU(3)_c \times SU(3)_L \times SU(3)_R \rightarrow SU(3)_{V+c}$

→ 8 massive gluons + 8 pseudoscalar Goldstone bosons (chiral limit)

- ▶ **baryon number:**  $U(1)$  broken → 1 scalar Goldstone boson

- ▶ **electromagnetism:**

**unbroken  $U(1)$**  gauge symmetry with charge  $\tilde{Q} = Q - \frac{1}{2}\lambda_3 - \frac{1}{2\sqrt{3}}\lambda_8$

→ color but not electromagnetic superconductor, baryon number superfluid

---

# Which phase is favored?

## - Realistic systems

---



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

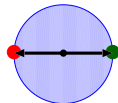
# Which phase is favored?

## - Realistic systems

### ► Reminder:

- Cooper instability: each partner close to the Fermi surface
- BCS pairing:  $\vec{p}^{(1)} = -\vec{p}^{(2)}$

→ works only if  $p_F^{(1)} \approx p_F^{(2)}$





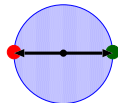
# Which phase is favored?

## - Realistic systems

### ▶ Reminder:

- ▶ Cooper instability: each partner close to the Fermi surface
- ▶ BCS pairing:  $\vec{p}^{(1)} = -\vec{p}^{(2)}$

→ works only if  $p_F^{(1)} \approx p_F^{(2)}$



- ▶ Quark masses:  $M_s \gg M_{u,d} \Rightarrow$  **unequal Fermi momenta**  $p_F^{(f)} = \sqrt{\mu^2 - M_f^2}$

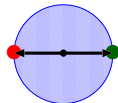
# Which phase is favored?

## - Realistic systems

### ▶ Reminder:

- ▶ Cooper instability: each partner close to the Fermi surface
- ▶ BCS pairing:  $\vec{p}^{(1)} = -\vec{p}^{(2)}$

→ works only if  $p_F^{(1)} \approx p_F^{(2)}$



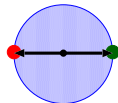
- ▶ Quark masses:  $M_s \gg M_{u,d} \Rightarrow$  **unequal Fermi momenta**  $p_F^{(f)} = \sqrt{\mu^2 - M_f^2}$
- ▶ Chandrasekhar, Clogston (1962): pairing still favored if  $\delta p_F \lesssim \frac{\Delta}{\sqrt{2}}$

# Which phase is favored?

## - Realistic systems

### ▶ Reminder:

- ▶ Cooper instability: each partner close to the Fermi surface
- ▶ BCS pairing:  $\vec{p}^{(1)} = -\vec{p}^{(2)}$



→ works only if  $p_F^{(1)} \approx p_F^{(2)}$

▶ Quark masses:  $M_s \gg M_{u,d} \Rightarrow$  **unequal Fermi momenta**  $p_F^{(f)} = \sqrt{\mu^2 - M_f^2}$

▶ Chandrasekhar, Clogston (1962): pairing still favored if  $\delta p_F \lesssim \frac{\Delta}{\sqrt{2}}$

### ▶ Expected phase structure:

- ▶  $\mu \gg M_s \Rightarrow p_F^{(s)} \approx p_F^{(u,d)} \rightarrow$  CFL
- ▶  $\mu \lesssim M_s \Rightarrow p_F^{(s)} \ll p_F^{(u,d)} \rightarrow$  2SC

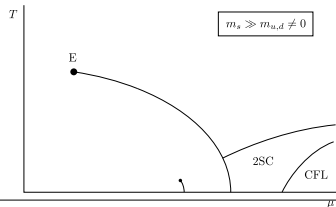


Figure: educated guess [Rajagopal (1999)]

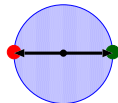
# Which phase is favored?

## - Realistic systems

### ▶ Reminder:

- ▶ Cooper instability: each partner close to the Fermi surface
- ▶ BCS pairing:  $\vec{p}^{(1)} = -\vec{p}^{(2)}$

→ works only if  $p_F^{(1)} \approx p_F^{(2)}$



▶ Quark masses:  $M_s \gg M_{u,d} \Rightarrow$  **unequal Fermi momenta**  $p_F^{(f)} = \sqrt{\mu^2 - M_f^2}$

▶ Chandrasekhar, Clogston (1962): pairing still favored if  $\delta p_F \lesssim \frac{\Delta}{\sqrt{2}}$

### ▶ Expected phase structure:

- ▶  $\mu \gg M_s \Rightarrow p_F^{(s)} \approx p_F^{(u,d)} \rightarrow$  CFL
- ▶  $\mu \lesssim M_s \Rightarrow p_F^{(s)} \ll p_F^{(u,d)} \rightarrow$  2SC

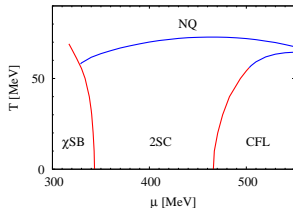


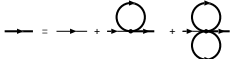
Figure: NJL [M. Oertel, MB (2002); MB (2005)]

# Role of the strange quark mass

- ▶ **NJL model:** treatment of (dynamical) masses and gaps on an equal footing
  - $T$  and  $\mu$  dependent quantities

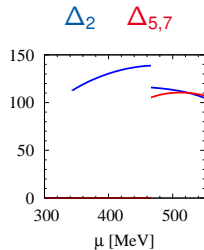
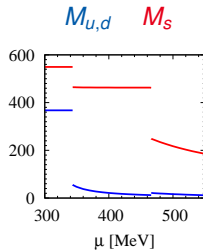
# Role of the strange quark mass

- ▶ **NJL model:** treatment of (dynamical) masses and gaps on an equal footing
  - $T$  and  $\mu$  dependent quantities

▶ Masses: 

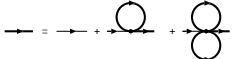
$$M_s = m_s - 4G\langle\bar{s}s\rangle + 2K\langle\bar{u}u\rangle\langle\bar{d}d\rangle$$

- $M_s$  large in the 2SC phase
- stabilizes the 2SC phase



# Role of the strange quark mass

- ▶ **NJL model:** treatment of (dynamical) masses and gaps on an equal footing  
→  $T$  and  $\mu$  dependent quantities

- ▶ **Masses:**   
$$M_s = m_s - 4G\langle\bar{s}s\rangle + 2K\langle\bar{u}u\rangle\langle\bar{d}d\rangle$$

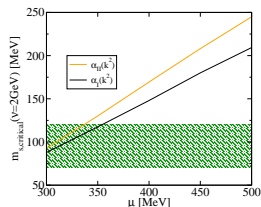
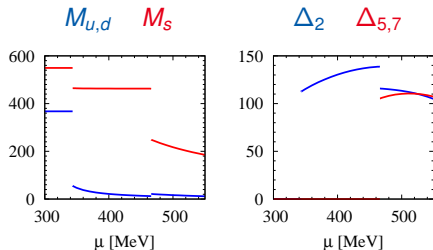
- $M_s$  large in the 2SC phase
- stabilizes the 2SC phase

- ▶ **Dyson-Schwinger QCD studies**

[Nickel, Alkofer, Wambach (2006)]

$$\text{quark} \text{---}^{-1} = \text{quark} \text{---}^{-1} + \text{quark} \text{---}^{-1} \text{---} \text{ghost} \text{---}^{-1}$$

- gluons screened by light quarks
- $M_s$  smaller in the 2SC phase
- CFL phase favored much earlier





## ► constraints in compact stars:

- color neutrality:  $n_r = n_g = n_b$
- electric neutrality:  $n_Q = \frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s - n_e = 0$
- $\beta$  equilibrium:  $\mu_e = \mu_d - \mu_u \Rightarrow n_e \ll n_{u,d}$



## ► constraints in compact stars:

- color neutrality: *(minor effect)*
  - electric neutrality:
  - $\beta$  equilibrium:
- $$\left. \begin{array}{l} \text{electric neutrality:} \\ \beta \text{ equilibrium:} \end{array} \right\} \frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s \approx 0$$

► constraints in compact stars:

- color neutrality: *(minor effect)*
  - electric neutrality:
  - $\beta$  equilibrium:
- $$\left. \begin{array}{l} \text{electric neutrality:} \\ \beta \text{ equilibrium:} \end{array} \right\} \frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s \approx 0$$

- $M_S > M_{u,d} \Rightarrow$  All flavors have different Fermi momenta.

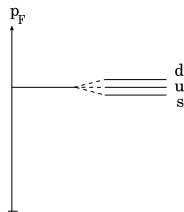
► constraints in compact stars:

- color neutrality: *(minor effect)*
  - electric neutrality:
  - $\beta$  equilibrium:
- $$\left. \begin{array}{l} \\ \\ \end{array} \right\} \frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s \approx 0$$

►  $M_S > M_{u,d} \Rightarrow$  All flavors have different Fermi momenta.

► Expansion in small  $M_S$  [Alford, Rajagopal (2002)]

- equidistant splitting
- no 2SC phase in compact stars



► constraints in compact stars:

- color neutrality: (minor effect)
- electric neutrality:  $\left. \begin{array}{l} \\ \\ \end{array} \right\} \frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s \approx 0$
- $\beta$  equilibrium:

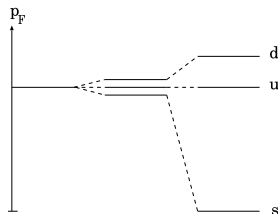
►  $M_S > M_{u,d} \Rightarrow$  All flavors have different Fermi momenta.

► Expansion in small  $M_S$  [Alford, Rajagopal (2002)]

- equidistant splitting
- no 2SC phase in compact stars

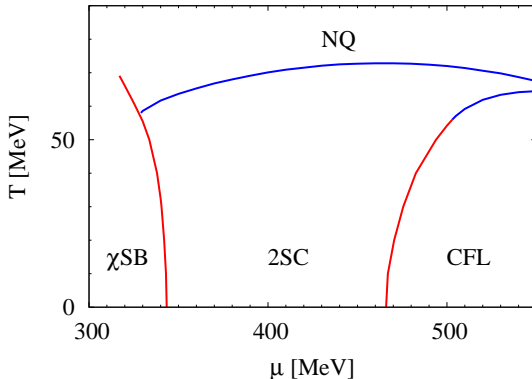
► Large  $M_S$

- $n_s \approx 0, n_d \approx 2n_u \Rightarrow p_F^{(d)} \approx 2^{1/3} p_F^{(u)} \approx 1.26 p_F^{(u)}$
- 2SC pairing possible for strong couplings



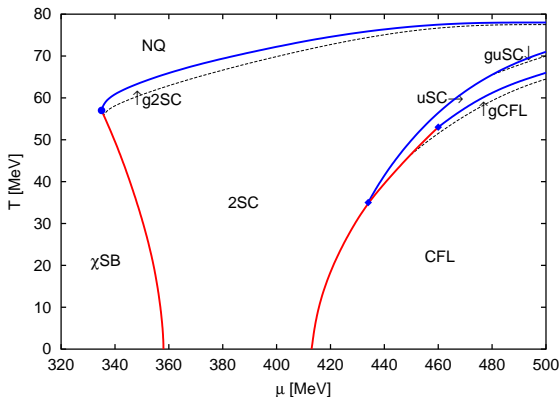
## Phase diagram **without** neutrality constraints

[M. Oertel, MB (2002)]



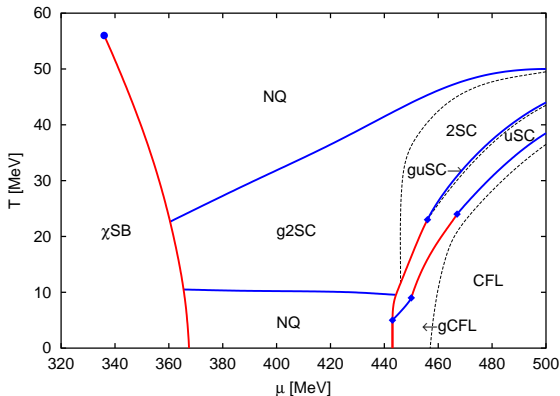
## Phase diagram with neutrality constraints: “strong” $qq$ coupling ( $H = G$ )

[Rüster, Werth, MB, Shovkovy, Rischke, (2005)]



Phase diagram with neutrality constraints: “intermediate”  $qq$  coupling ( $H = 0.75 G$ )

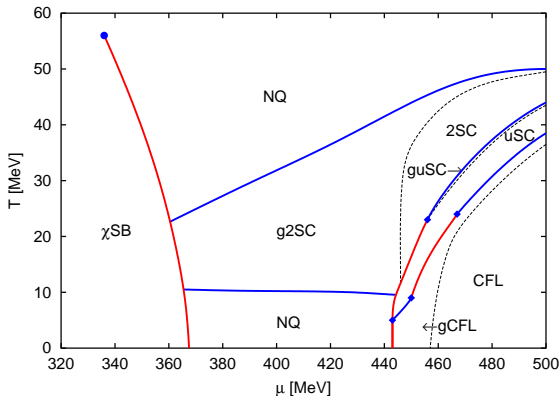
[Rüster, Werth, MB, Shovkovy, Rischke, (2005)]



# NJL model results

Phase diagram with neutrality constraints: “intermediate”  $qq$  coupling ( $H = 0.75 G$ )

[Rüster, Werth, MB, Shovkovy, Rischke, (2005)]

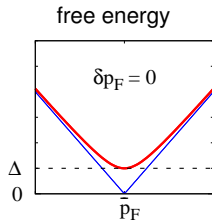


→ strong parameter dependence



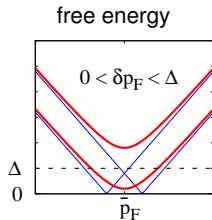
# Gapless color superconductors

- ▶ unequal Fermi momenta:  $p_F^{a,b} = \bar{p}_F \pm \delta p_F$



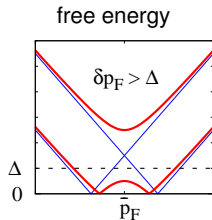
# Gapless color superconductors

- ▶ unequal Fermi momenta:  $p_F^{a,b} = \bar{p}_F \pm \delta p_F$ 
  - ▶ splitting of the quasiparticle modes:  $\omega(\vec{p}) \rightarrow |\omega(\vec{p}) \pm \delta p_F|$



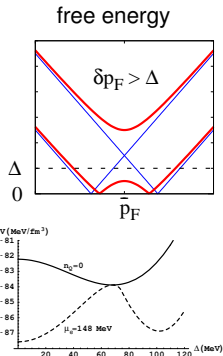
# Gapless color superconductors

- ▶ unequal Fermi momenta:  $p_F^{a,b} = \bar{p}_F \pm \delta p_F$ 
  - ▶ splitting of the quasiparticle modes:  $\omega(\vec{p}) \rightarrow |\omega(\vec{p}) \pm \delta p_F|$
  - ▶  $\delta p_F > \Delta \rightarrow$  gapless modes



# Gapless color superconductors

- ▶ unequal Fermi momenta:  $p_F^{a,b} = \bar{p}_F \pm \delta p_F$ 
  - ▶ splitting of the quasiparticle modes:  $\omega(\vec{p}) \rightarrow |\omega(\vec{p}) \pm \delta p_F|$
  - ▶  $\delta p_F > \Delta \rightarrow$  gapless modes
- ▶ gapless CSC phases
  - ▶ unstable solution (maximum) at fixed  $\mu_e$
  - ▶ can be most favored neutral homogeneous solution



[Shovkovy, Huang (2003)]

# Gapless color superconductors

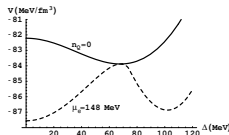
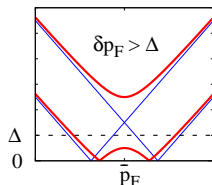
- ▶ unequal Fermi momenta:  $\rho_F^{a,b} = \bar{\rho}_F \pm \delta\rho_F$ 
  - ▶ splitting of the quasiparticle modes:  $\omega(\vec{p}) \rightarrow |\omega(\vec{p}) \pm \delta\rho_F|$
  - ▶  $\delta\rho_F > \Delta \rightarrow$  gapless modes
- ▶ gapless CSC phases
  - ▶ unstable solution (maximum) at fixed  $\mu_e$
  - ▶ can be most favored neutral homogeneous solution

- ▶ Meissner effect:



$$m_{M,a}^2 = -\frac{1}{2} \lim_{\vec{p} \rightarrow 0} \left( g_{ij} + \frac{p_i p_j}{p^2} \right) \Pi_{aa}^{ij}(0, \vec{p})$$

free energy



[Shovkovy, Huang (2003)]

# Gapless color superconductors

- ▶ unequal Fermi momenta:  $p_F^{a,b} = \bar{p}_F \pm \delta p_F$ 
  - ▶ splitting of the quasiparticle modes:  $\omega(\vec{p}) \rightarrow |\omega(\vec{p}) \pm \delta p_F|$
  - ▶  $\delta p_F > \Delta \rightarrow$  gapless modes
- ▶ gapless CSC phases
  - ▶ unstable solution (maximum) at fixed  $\mu_e$
  - ▶ can be most favored neutral homogeneous solution

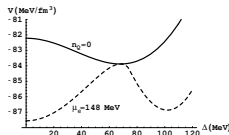
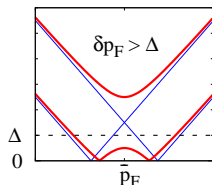
- ▶ Meissner effect:



$$m_{M,a}^2 = -\frac{1}{2} \lim_{\vec{p} \rightarrow 0} \left( g_{ij} + \frac{p_i p_j}{p^2} \right) \Pi_{aa}^{ij}(0, \vec{p})$$

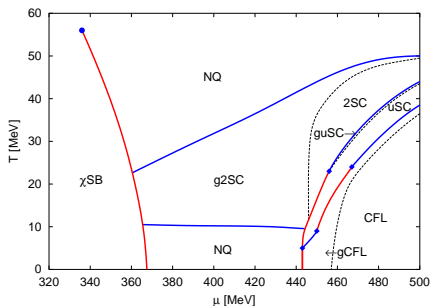
- ▶ chromomagnetic instability:  $m_{M,a}^2 < 0$  for  $\delta p_F > \begin{cases} \frac{\Delta}{\sqrt{2}} & a = 4, \dots, 7 \\ \Delta & a = 8 \end{cases}$

free energy



[Shovkovy, Huang (2003)]

# Main issues



- ▶ strong parameter dependence
- ▶ unstable phases

---

# Getting rid of the parameter dependence

---



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

---



# Getting rid of the parameter dependence



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

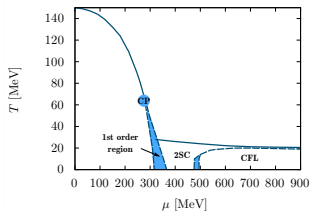
## 1. Theoretical approaches: starting from QCD

## 1. Theoretical approaches: starting from QCD

### ► Dyson-Schwinger equations:

[Nickel, Alkofer, Wambach (2006, 2008),  
Müller, MB, Wambach (2013, 2016)]

- phase diagram without neutrality constraints
- no cutoff artifacts at large  $\mu$
- determination of the pressure difficult
- still strong dependence on truncations and renormalization conditions



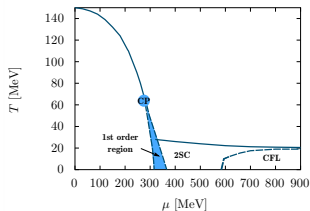
[Müller et al. (2013)]

## 1. Theoretical approaches: starting from QCD

### ► Dyson-Schwinger equations:

[Nickel, Alkofer, Wambach (2006, 2008),  
Müller, MB, Wambach (2013, 2016)]

- phase diagram without neutrality constraints
- no cutoff artifacts at large  $\mu$
- determination of the pressure difficult
- still strong dependence on truncations and renormalization conditions



[Müller et al. (2013)]

# Getting rid of the parameter dependence



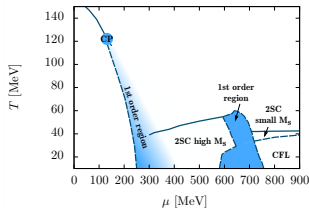
TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

## 1. Theoretical approaches: starting from QCD

### ► Dyson-Schwinger equations:

[Nickel, Alkofer, Wambach (2006, 2008),  
Müller, MB, Wambach (2013, 2016)]

- phase diagram without neutrality constraints
- no cutoff artifacts at large  $\mu$
- determination of the pressure difficult
- still strong dependence on truncations and renormalization conditions



[Müller et al. (2016)]

# Getting rid of the parameter dependence

## 1. Theoretical approaches: starting from QCD

### ► Dyson-Schwinger equations:

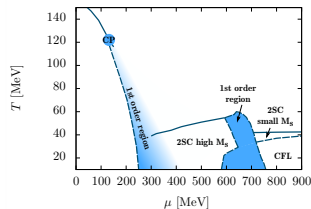
[Nickel, Alkofer, Wambach (2006, 2008),  
Müller, MB, Wambach (2013, 2016)]

- phase diagram without neutrality constraints
- no cutoff artifacts at large  $\mu$
- determination of the pressure difficult
- still strong dependence on truncations and renormalization conditions

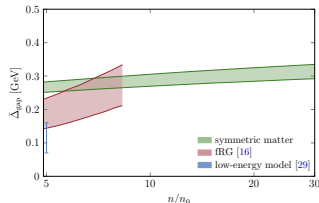
### ► Functional renormalization group:

[Braun, Schallmo (2022)]

- study 2SC pairing at  $T = 0$  by solving QCD flow equations at large  $\mu$  → **very large gaps!**



[Müller et al. (2016)]



[Braun, Schallmo (2022)]

# Getting rid of the parameter dependence



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

## 2. Using empirical information

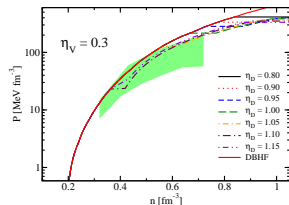
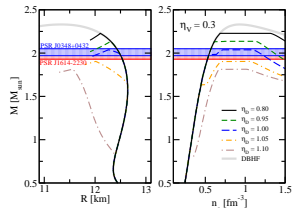
# Getting rid of the parameter dependence

## 2. Using empirical information

- ▶ Fitting NJL parameters to **astrophysical constraints** and **heavy-ion data**:

[Klähn, Blaschke, ... (2006, 2007, 2013, ... )]

- ▶ purely hadronic matter inconsistent (see also [Annala et al. (2020)])
- ▶ vector repulsion to be stiff enough
- ▶ strong  $qq$  interaction



[Klähn, Łastowiecki, Blaschke (2013)]

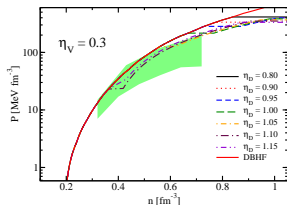
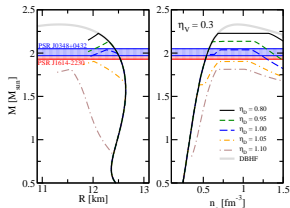
# Getting rid of the parameter dependence

## 2. Using empirical information

- ▶ Fitting NJL parameters to **astrophysical constraints** and **heavy-ion data**:

[Klähn, Blaschke, ... (2006, 2007, 2013, ... )]

- ▶ purely hadronic matter inconsistent (see also [Annala et al. (2020)])
- ▶ vector repulsion to be stiff enough
- ▶ strong  $qq$  interaction
- ▶ Signals of CSC in the gravitational-wave spectrum from **neutron-star mergers**?
  - ▶ part of project B09 in the CRC-TR 211  
↔ Hosein's thesis project  
(see his talk for preparatory work)



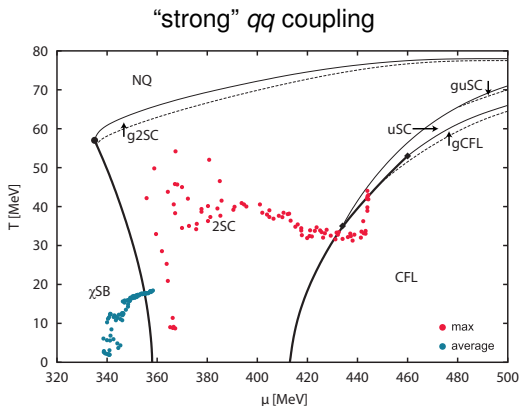
[Klähn, Łastowiecki, Blaschke (2013)]



# CSC phases in neutron-star mergers (propaganda plots)

## Overlay of unrelated calculations:

Points from merger simulations with purely hadronic EoS [E. Most, L. Rezzolla, priv. comm.]  
and NJL phase diagrams [Rüster et al. (2005, 2006)]

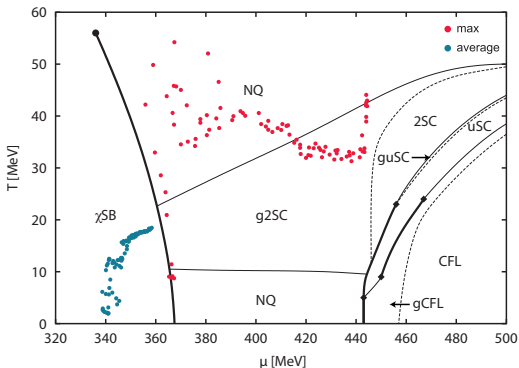


# CSC phases in neutron-star mergers (propaganda plots)

## Overlay of unrelated calculations:

Points from merger simulations with purely hadronic EoS [E. Most, L. Rezzolla, priv. comm.]  
and NJL phase diagrams [Rüster et al. (2005, 2006)]

“intermediate”  $qq$  coupling

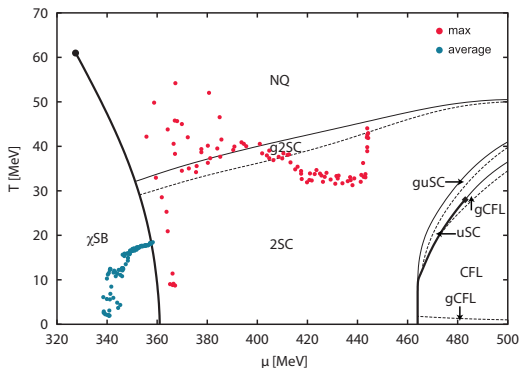


# CSC phases in neutron-star mergers (propaganda plots)

## Overlay of unrelated calculations:

Points from merger simulations with purely hadronic EoS [E. Most, L. Rezzolla, priv. comm.]  
and NJL phase diagrams [Rüster et al. (2005, 2006)]

“intermediate”  $qq$  coupling + neutrino chemical potential  $\mu_\nu = 200$  MeV



- ▶ **Proto-neutron stars:** neutrinos trapped during the first few seconds

- lepton number conserved

- more electrons:

$$\mu_e = \mu_d - \mu_u + \mu_\nu$$

- favors 2SC, suppresses CFL

[Steiner, Reddy, Prakash, PRD (2002)]

► Proto-neutron stars: neutrinos trapped during the first few seconds

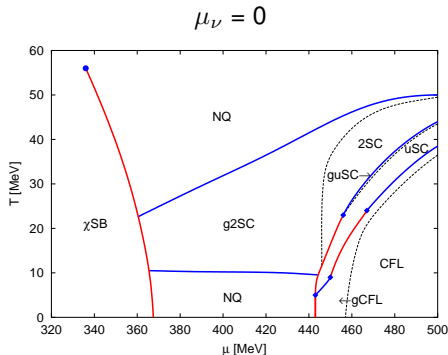
→ lepton number conserved

→ more electrons:

$$\mu_e = \mu_d - \mu_u + \mu_\nu$$

→ favors 2SC, suppresses CFL

[Steiner, Reddy, Prakash, PRD (2002)]



[Rüster, Werth, M.B., Shovkovy, Rischke, PRD (2006)]

► Proto-neutron stars: neutrinos trapped during the first few seconds

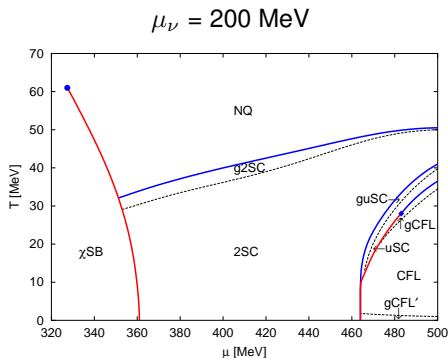
→ lepton number conserved

→ more electrons:

$$\mu_e = \mu_d - \mu_u + \mu_\nu$$

→ favors 2SC, suppresses CFL

[Steiner, Reddy, Prakash, PRD (2002)]



[Rüster, Werth, M.B., Shovkovy, Rischke, PRD (2006)]

► Proto-neutron stars: neutrinos trapped during the first few seconds

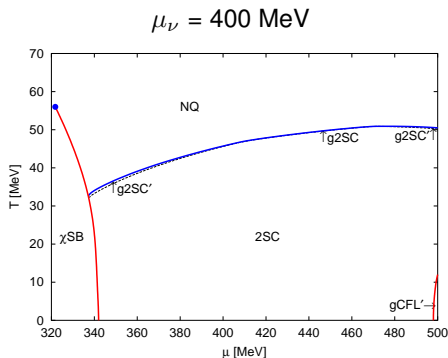
→ lepton number conserved

→ more electrons:

$$\mu_e = \mu_d - \mu_u + \mu_\nu$$

→ favors 2SC, suppresses CFL

[Steiner, Reddy, Prakash, PRD (2002)]



[Rüster, Werth, M.B., Shovkovy, Rischke, PRD (2006)]

- ▶ **Proto-neutron stars:** neutrinos trapped during the first few seconds

- lepton number conserved

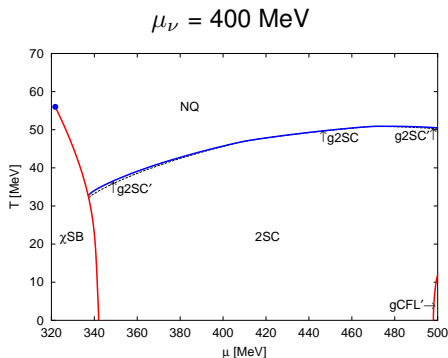
- more electrons:

$$\mu_e = \mu_d - \mu_U + \mu_\nu$$

- favors 2SC, suppresses CFL

[Steiner, Reddy, Prakash, PRD (2002)]

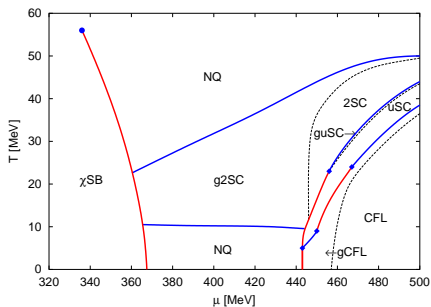
- ▶ also relevant for  
**neutron-star mergers!**



[Rüster, Werth, M.B., Shovkovy, Rischke, PRD (2006)]



# Main issues



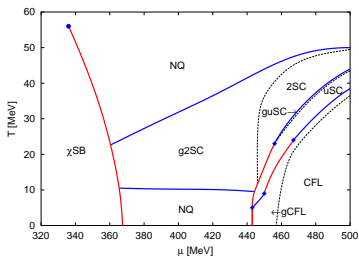
- ▶ strong parameter dependence
- ▶ unstable phases

# Kaon condensation in the CFL phase

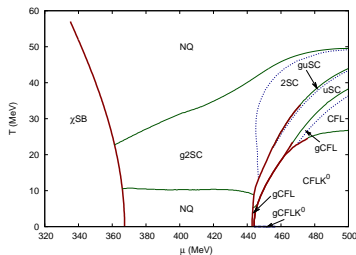
- ▶ CFL: chiral symmetry broken → Goldstone bosons  $\sim \mathcal{O}(10 \text{ MeV})$   
[Son, Stephanov, PRD (2000)]
- ▶  $\mu_s^{\text{eff}} \simeq \frac{m_s^2 - m_u^2}{2\mu} \rightarrow K^0$  condensation [T. Schäfer, PRL (2000); Bedaque, Schäfer, NPA (2002)]
- ▶ NJL model: include pseudoscalar diquark conds. [M.B., PLB (2005); M.M. Forbes, PRD (2005)]

# Kaon condensation in the CFL phase

- ▶ CFL: chiral symmetry broken  $\rightarrow$  Goldstone bosons  $\sim \mathcal{O}(10 \text{ MeV})$   
[Son, Stephanov, PRD (2000)]
- ▶  $\mu_s^{\text{eff}} \simeq \frac{m_s^2 - m_u^2}{2\mu} \rightarrow K^0$  condensation [T. Schäfer, PRL (2000); Bedaque, Schäfer, NPA (2002)]
- ▶ NJL model: include **pseudoscalar** diquark conds. [M.B., PLB (2005); M.M. Forbes, PRD (2005)]
- ▶ phase diagram:



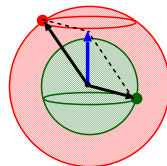
[Rüster, Werth, M.B., Shovkovy, Rischke, PRD (2005)]



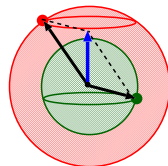
[H. Basler, M.B., PRD (2010); H. Warringa (2006)]

- ▶ BCS pairing disfavored for  $\delta p_F \gtrsim \frac{\Delta}{\sqrt{2}}$
- ▶ alternative: pairs with nonzero total momentum

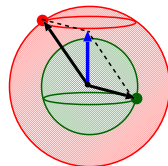
- ▶ BCS pairing disfavored for  $\delta p_F \gtrsim \frac{\Delta}{\sqrt{2}}$
- ▶ alternative: **pairs with nonzero total momentum**
- ▶ FF: [P. Fulde, R.A. Ferrell, Phys. Rev., 1964]
  - ▶ single plane wave,  $\langle q(\vec{x})q(\vec{x}) \rangle \sim \Delta e^{2i\vec{q}\cdot\vec{x}}$  for fixed  $\vec{q}$
  - ▶ disfavored by phase space



- ▶ BCS pairing disfavored for  $\delta p_F \gtrsim \frac{\Delta}{\sqrt{2}}$
- ▶ alternative: **pairs with nonzero total momentum**
- ▶ FF: [P. Fulde, R.A. Ferrell, Phys. Rev., 1964]
  - ▶ single plane wave,  $\langle q(\vec{x})q(\vec{x}) \rangle \sim \Delta e^{2i\vec{q}\cdot\vec{x}}$  for fixed  $\vec{q}$
  - ▶ disfavored by phase space
- ▶ LO: [Larkin, Ovchinnikov, Zh. Eksp. Teor. Fiz., 1964]
  - ▶ multiple plane waves (e.g.,  $\cos(2\vec{q} \cdot \vec{x})$ )



- ▶ BCS pairing disfavored for  $\delta p_F \gtrsim \frac{\Delta}{\sqrt{2}}$
- ▶ alternative: **pairs with nonzero total momentum**
- ▶ FF: [P. Fulde, R.A. Ferrell, Phys. Rev., 1964]
  - ▶ single plane wave,  $\langle q(\vec{x})q(\vec{x}) \rangle \sim \Delta e^{2i\vec{q}\cdot\vec{x}}$  for fixed  $\vec{q}$
  - ▶ disfavored by phase space
- ▶ LO: [Larkin, Ovchinnikov, Zh. Eksp. Teor. Fiz., 1964]
  - ▶ multiple plane waves (e.g.,  $\cos(2\vec{q} \cdot \vec{x})$ )
- ▶ LOFF in CSC ( $\rightarrow$  [Anglani et al., Rev. Mod. Phys. (2014)])

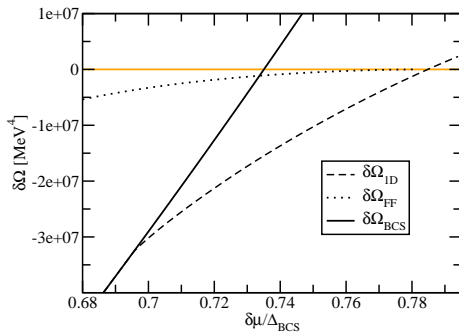


Indications:

chromomagnetic instabilities  $\leftrightarrow$  **instabilities towards LOFF phases**

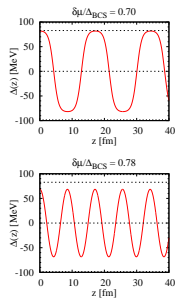
[Giannakis, Ren; Giannakis, Hou, Ren, PLB (2005)]

free-energy gain:



[D. Nickel, MB, PRD (2008)]

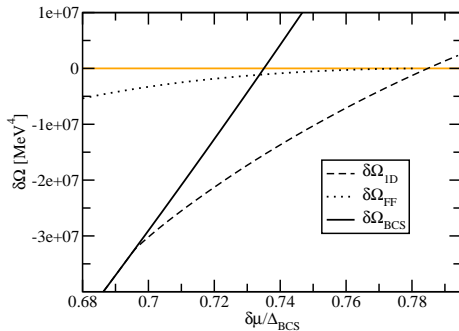
condensate shapes:





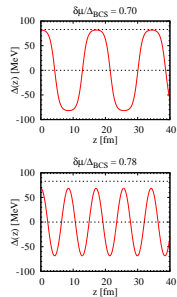
# NJL-model results

free-energy gain:



[D. Nickel, MB, PRD (2008)]

condensate shapes:



- ▶ still missing: comprehensive calculation of neutral phase diagram with LOFF phases



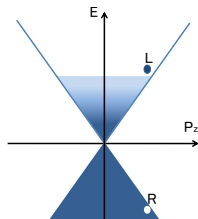
# INHOMOGENEOUS CHIRAL PHASES



- ▶  $\langle \bar{q}q \rangle = \langle \bar{q}_L q_R \rangle + \langle \bar{q}_R q_L \rangle$
- ▶ **chiral-symmetry breaking in vacuum:**  
pairing a left-handed quark with a right-handed antiquark (and vice versa)

# Stressed chiral-symmetry breaking

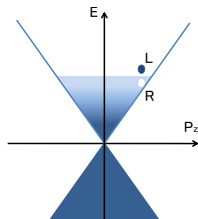
- ▶  $\langle \bar{q}q \rangle = \langle \bar{q}_L q_R \rangle + \langle \bar{q}_R q_L \rangle$
- ▶ chiral-symmetry breaking in vacuum:  
pairing a left-handed quark with a right-handed antiquark (and vice versa)
- ▶ finite chemical potential  $\rightarrow$  “pairing stress”  
(large excitation energy  $E \approx 2\mu$ )



[Kojo et al. (2010)]

# Stressed chiral-symmetry breaking

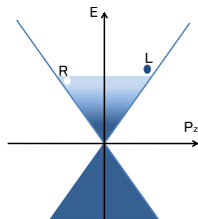
- ▶  $\langle \bar{q}q \rangle = \langle \bar{q}_L q_R \rangle + \langle \bar{q}_R q_L \rangle$
- ▶ chiral-symmetry breaking in vacuum:  
pairing a left-handed quark with a right-handed antiquark (and vice versa)
- ▶ finite chemical potential  $\rightarrow$  “pairing stress”  
(large excitation energy  $E \approx 2\mu$ )
- ▶ alternative: particle-hole pairing
  - ▶  $P_{tot} = 0 \Leftrightarrow P_{rel} \approx 2\mu$   
 $\rightarrow$  interaction weak  $\rightarrow$  not favored



[Kojo et al. (2010)]

# Stressed chiral-symmetry breaking

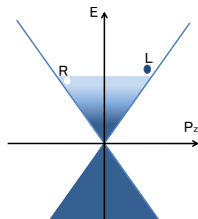
- ▶  $\langle \bar{q}q \rangle = \langle \bar{q}_L q_R \rangle + \langle \bar{q}_R q_L \rangle$
- ▶ chiral-symmetry breaking in vacuum:  
pairing a left-handed quark with a right-handed antiquark (and vice versa)
- ▶ finite chemical potential  $\rightarrow$  “pairing stress”  
(large excitation energy  $E \approx 2\mu$ )
- ▶ alternative: particle-hole pairing
  - ▶  $P_{tot} = 0 \Leftrightarrow P_{rel} \approx 2\mu$   
 $\rightarrow$  interaction weak  $\rightarrow$  not favored
  - ▶  $P_{rel} \approx 0 \Leftrightarrow P_{tot} \approx 2\mu \rightarrow$  interaction strong  $\rightarrow$  favored in some window



[Kojo et al. (2010)]

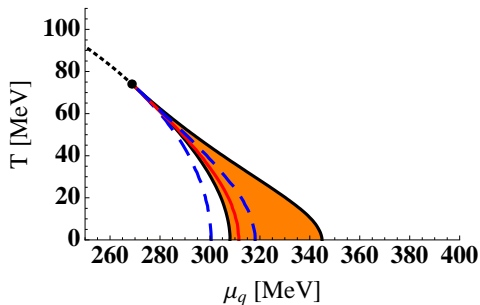
# Stressed chiral-symmetry breaking

- ▶  $\langle \bar{q}q \rangle = \langle \bar{q}_L q_R \rangle + \langle \bar{q}_R q_L \rangle$
- ▶ **chiral-symmetry breaking in vacuum:**  
pairing a left-handed quark with a right-handed antiquark (and vice versa)
- ▶ **finite chemical potential**  $\rightarrow$  “pairing stress”  
(large excitation energy  $E \approx 2\mu$ )
- ▶ alternative: **particle-hole pairing**
  - ▶  $P_{tot} = 0 \Leftrightarrow P_{rel} \approx 2\mu$   
 $\rightarrow$  interaction weak  $\rightarrow$  not favored
  - ▶  $P_{rel} \approx 0 \Leftrightarrow P_{tot} \approx 2\mu \rightarrow$  interaction strong  $\rightarrow$  favored in some window  
 $\rightarrow$  **inhomogeneous chiral condensates!**



[Kojo et al. (2010)]

NJL [D. Nickel (2009)]

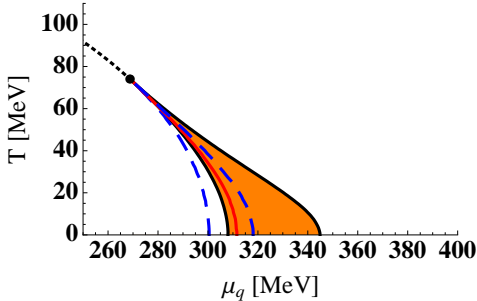


- ▶ **NJL:** inhomogeneous phase covers homogeneous first-order line

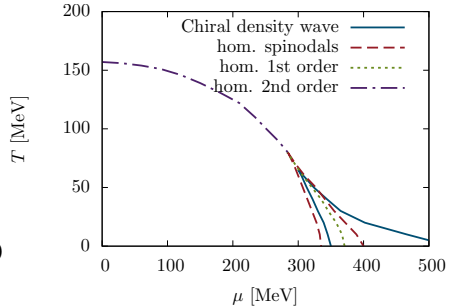


# Phase diagram

NJL [D. Nickel (2009)]



DSE [D. Müller et al. (2013)]



- ▶ **NJL:** inhomogeneous phase covers homogeneous first-order line
- ▶ **DSE:** phase-transition region qualitatively similar

# Methods to investigate inhomogeneous phases



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

- ▶ **Main difficulty:** **Functional minimization** of the free energy  
w.r.t. to arbitrarily shaped condensate functions  $\phi(\vec{x})$

$$\text{e.g., } \Omega[\phi(\vec{x})] = \frac{T}{V} \text{Tr} \log \frac{S^{-1}[\phi(\vec{x})]}{T} + \frac{1}{V} \int_V d^3x \left\{ \frac{1}{2} (\nabla \phi(\vec{x}))^2 + U(\phi(\vec{x})) \right\}$$

# Methods to investigate inhomogeneous phases

- ▶ **Main difficulty:** **Functional minimization** of the free energy  
w.r.t. to arbitrarily shaped condensate functions  $\phi(\vec{x})$

$$\text{e.g., } \Omega[\phi(\vec{x})] = \frac{T}{V} \text{Tr} \log \frac{S^{-1}[\phi(\vec{x})]}{T} + \frac{1}{V} \int_V d^3x \left\{ \frac{1}{2} (\nabla \phi(\vec{x}))^2 + U(\phi(\vec{x})) \right\}$$

- ▶ **strategy 1:** choose certain **ansatz functions**
  - ▶ chiral density wave (= single plane wave) [Nakano and Tatsumi (2005)]
  - ▶ 1D Jacobi elliptic functions (“real kink crystal”) [Nickel (2009)]
  - ▶ 2D modulations (quadratic, hexagonal) [Carignano, MB (2012)]

# Methods to investigate inhomogeneous phases

- ▶ **Main difficulty:** **Functional minimization** of the free energy  
w.r.t. to arbitrarily shaped condensate functions  $\phi(\vec{x})$

$$\text{e.g., } \Omega[\phi(\vec{x})] = \frac{T}{V} \text{Tr} \log \frac{S^{-1}[\phi(\vec{x})]}{T} + \frac{1}{V} \int_V d^3x \left\{ \frac{1}{2} (\nabla \phi(\vec{x}))^2 + U(\phi(\vec{x})) \right\}$$

- ▶ **strategy 1:** choose certain **ansatz functions**
  - ▶ chiral density wave (= single plane wave) [Nakano and Tatsumi (2005)]
  - ▶ 1D Jacobi elliptic functions (“real kink crystal”) [Nickel (2009)]
  - ▶ 2D modulations (quadratic, hexagonal) [Carignano, MB (2012)]
- ▶ **strategy 2:** **lattice field theory** [M. Wagner (2007), Lenz et al. 2020, ... ]

# Methods to investigate inhomogeneous phases



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

- ▶ **Main difficulty:** **Functional minimization** of the free energy  
w.r.t. to arbitrarily shaped condensate functions  $\phi(\vec{x})$

$$\text{e.g., } \Omega[\phi(\vec{x})] = \frac{T}{V} \text{Tr} \log \frac{S^{-1}[\phi(\vec{x})]}{T} + \frac{1}{V} \int_V d^3x \left\{ \frac{1}{2} (\nabla \phi(\vec{x}))^2 + U(\phi(\vec{x})) \right\}$$

- ▶ **strategy 1:** choose certain **ansatz functions**
  - ▶ chiral density wave (= single plane wave) [Nakano and Tatsumi (2005)]
  - ▶ 1D Jacobi elliptic functions (“real kink crystal”) [Nickel (2009)]
  - ▶ 2D modulations (quadratic, hexagonal) [Carignano, MB (2012)]
- ▶ **strategy 2:** **lattice field theory** [M. Wagner (2007), Lenz et al. 2020, ... ]

- ▶ **strategy 3:** **stability analysis**

split  $\phi(\vec{x}) = \bar{\phi} + \delta\phi(\vec{x})$ ,  $\bar{\phi}$  = **homogeneous** minimum

and expand in powers of the fluctuations  $\Omega = \sum_n \Omega^{(n)}$ ,  $\Omega^{(n)} = \mathcal{O}(\delta\phi^n)$

# Methods to investigate inhomogeneous phases



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

- ▶ **Main difficulty:** **Functional minimization** of the free energy  
w.r.t. to arbitrarily shaped condensate functions  $\phi(\vec{x})$

$$\text{e.g., } \Omega[\phi(\vec{x})] = \frac{T}{V} \text{Tr} \log \frac{S^{-1}[\phi(\vec{x})]}{T} + \frac{1}{V} \int_V d^3x \left\{ \frac{1}{2} (\nabla \phi(\vec{x}))^2 + U(\phi(\vec{x})) \right\}$$

- ▶ **strategy 1:** choose certain **ansatz functions**

- ▶ chiral density wave (= single plane wave) [Nakano and Tatsumi (2005)]
- ▶ 1D Jacobi elliptic functions (“real kink crystal”) [Nickel (2009)]
- ▶ 2D modulations (quadratic, hexagonal) [Carignano, MB (2012)]

- ▶ **strategy 2:** **lattice field theory** [M. Wagner (2007), Lenz et al. 2020, ... ]

- ▶ **strategy 3:** **stability analysis**

split  $\phi(\vec{x}) = \bar{\phi} + \delta\phi(\vec{x})$ ,  $\bar{\phi}$  = **homogeneous** minimum

and expand in powers of the fluctuations  $\Omega = \sum_n \Omega^{(n)}$ ,  $\Omega^{(n)} = \mathcal{O}(\delta\phi^n)$

$\Omega^{(2)} < 0 \rightarrow$  instability towards an **inhomogeneous phase** (sufficient condition)

- ▶ seen in NJL and Quark-Meson model in (extended) mean-field approximation, QCD with DSEs; indications from FRG (“moat regime”) [Fu et al. (2020)]
- ▶ rather robust against model extensions and variations:  
(review: [MB, S. Carignano (2015)])
  - ▶ 3 flavors
  - ▶ vector interactions
  - ▶ magnetic fields
  - ▶ explicit symmetry breaking

- ▶ seen in NJL and Quark-Meson model in (extended) mean-field approximation, QCD with DSEs; indications from FRG (“moat regime”) [Fu et al. (2020)]
- ▶ rather robust against model extensions and variations:  
(review: [MB, S. Carignano (2015)])
  - ▶ 3 flavors
  - ▶ vector interactions
  - ▶ magnetic fields
  - ▶ explicit symmetry breaking
- ▶ open questions:
  - ▶ Are the inhomogeneous phases stable against fluctuations beyond mean field?  
→ FRG approach, see Lennart’s talk (next)
  - ▶ What are the favored condensate shapes?
  - ▶ What is the role of the cutoff? (see Laurin’s talk on Friday)