



# Effective spectral functions via lifetime analysis

Renan Hirayama

in collaboration with

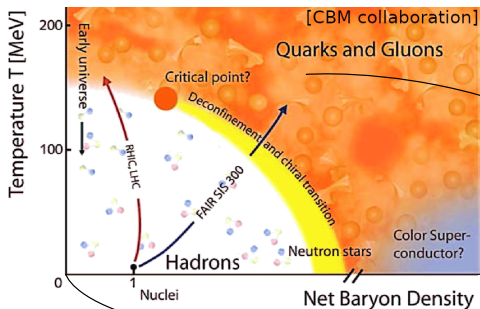
Jan Staudenmaier, Hannah Elfner

HFHF Retreat 2022, Castiglione della Pescaia



Helmholtz Forschungsakademie Hessen für FAIR

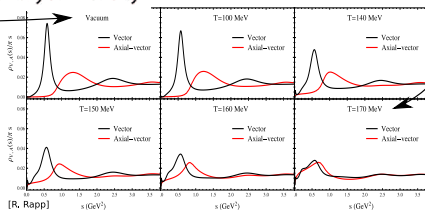
# Motivation - Theory



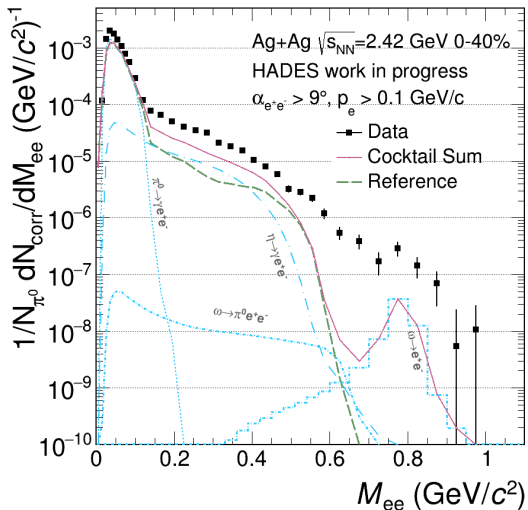
Chiral pairs degenerate

Vector :  $\rho(770)$

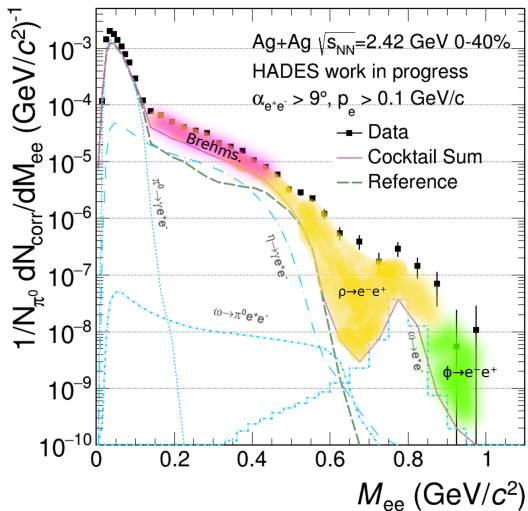
Axial :  $a_1(1260)$



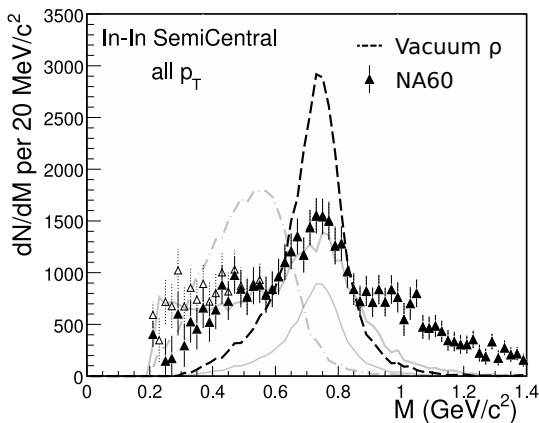
# Motivation - Experiment



# Motivation - Experiment



# Motivation - Experiment

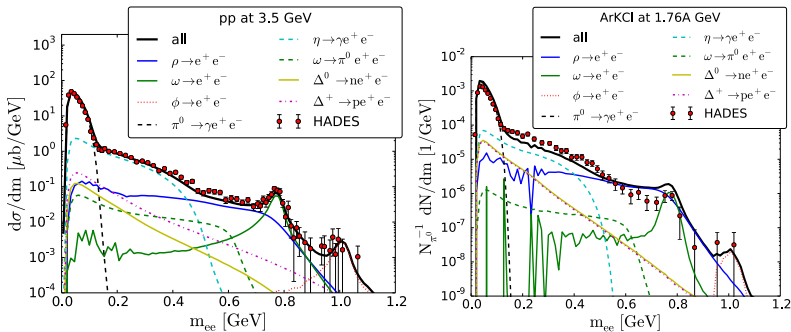


(Not the same experiment, but behavior is similar)

## Different approaches:

1. Vacuum spectral functions with *dynamic* broadening: UrQMD [S. Endres et al.: [PRC 92.1 \(2015\)](#)] and SMASH cascades [J. Weil et al.: [PRC 94 \(2016\)](#)]
2. Transport with *collisional* broadening dependent on local density: GiBUU [A.B. Larionov et al.: [PRC 102.6 \(2020\)](#)] and (P)HSD [E.L. Bratkovskaya et al.: [Nuc. Phys. A 807 \(2008\)](#)]
3. Full in-medium modifications: Rapp-Wambach model [H. van Hees and R. Rapp: [Eur. Phys. J. A6, 415–420 \(1999\)](#)] and FRG [R.A. Tripolt et al.: [PRD 104 \(2021\)](#)];

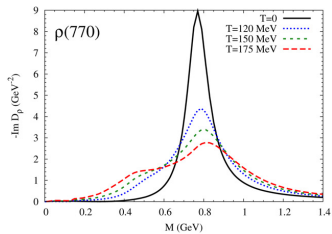
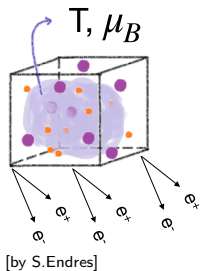
# SMASH dileptons



[J. Staudenmaier et al.: PRC 98 (2018)]

Dynamical broadening insufficient!

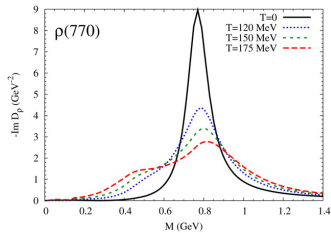
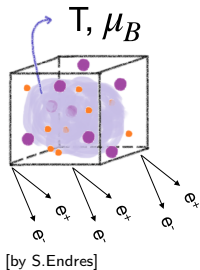
# SMASH dileptons



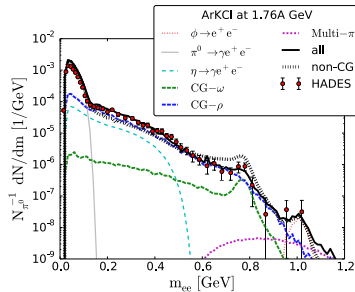
[Nuc.Phys. A 806 (2008)]



# SMASH dileptons



[Nuc.Phys. A 806 (2008)]



Quantify dynamic collisional broadening of  $\rho$  mesons

[2206.15166]

Addition of inelastic interactions where the  $\rho$  is absorbed

Vacuum

$$\rho^{\pm} \rightarrow \pi^0 \pi^{\mp}$$

$$\rho^0 \rightarrow \pi^+ \pi^-$$

$$\rho^0 \rightarrow l^+ l^-$$

Addition of inelastic interactions where the  $\rho$  is absorbed

Vacuum

$$\rho^\pm \rightarrow \pi^0 \pi^\mp$$

$$\rho^0 \rightarrow \pi^+ \pi^-$$

$$\rho^0 \rightarrow l^+ l^-$$

Medium

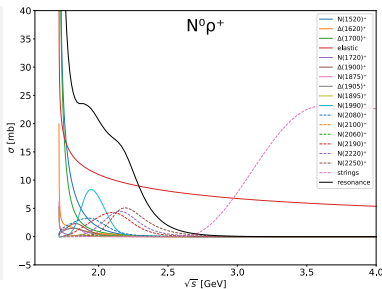
$$\rho N \rightarrow N^*, \Delta, \dots$$

$$\rho K \rightarrow K^*$$

$$\rho \pi \rightarrow \omega, \phi, f_1, a_1, \dots$$

# Medium modifications

Addition of inelastic interactions where the  $\rho$  is absorbed



Medium

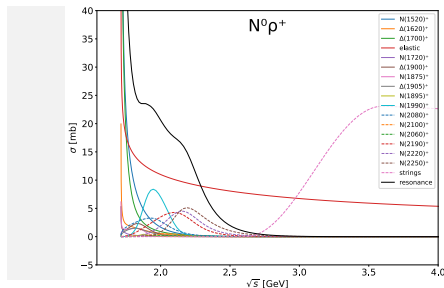
$$\rho N \rightarrow N^*, \Delta, \dots$$

$$\rho K \rightarrow K^*$$

$$\rho\pi \rightarrow \omega, \phi, f_1, a_1, \dots$$

# Medium modifications

Addition of inelastic interactions where the  $\rho$  is absorbed



Medium

$$\rho N \rightarrow N^*, \Delta, \dots$$

$$\rho K \rightarrow K^*$$

$$\rho \pi \rightarrow \omega, \phi, f_1, a_1, \dots$$

Collisional broadening: *Shortening the average lifetime of resonances due to absorptions by a medium*

Resonances are handled with vacuum properties:

- Decay width

$$\Gamma^{\text{vac}}(m) = \Gamma_{\rho \rightarrow \pi\pi}(m) + \Gamma_{\rho \rightarrow ll}(m)$$

- Breit-Wigner for mass sampling

$$\mathcal{A}(m) = \frac{2\mathcal{N}}{\pi} \frac{m^2 \Gamma^{\text{vac}}(m)}{(m^2 - M_0^2)^2 + m^2 \Gamma^{\text{vac}}(m)^2}$$

No *a priori* knowledge of medium!

Resonances are handled with vacuum properties:

- Decay width

$$\Gamma^{\text{vac}}(m) = \Gamma_{\rho \rightarrow \pi\pi}(m) + \Gamma_{\rho \rightarrow tt}(m) \xrightarrow{m \geq m_{2\pi}}$$

- Breit-Wigner for mass sampling

$$\mathcal{A}(m) = \frac{2\mathcal{N}}{\pi} \frac{m^2 \Gamma^{\text{vac}}(m)}{(m^2 - M_0^2)^2 + m^2 \Gamma^{\text{vac}}(m)^2}$$

No *a priori* knowledge of medium!

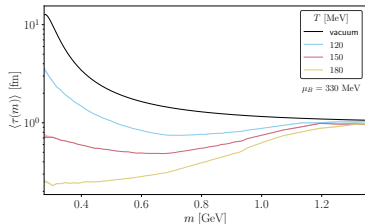


# An effective description

From the interaction history:

Effective width

$$\Gamma^{\text{eff}}(x) = \frac{1}{\langle \tau \rangle_x} = \left\langle \frac{\gamma}{t_f - t_i} \right\rangle_x$$



# An effective description

From the interaction history:

Effective width

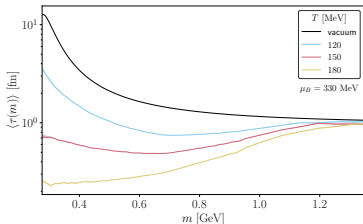
$$\Gamma^{\text{eff}}(x) = \frac{1}{\langle \tau \rangle_x} = \left\langle \frac{\gamma}{t_f - t_i} \right\rangle_x$$

Collisional width

$$\Gamma^{\text{col}} = \Gamma^{\text{eff}} - \Gamma^{\text{vac}}$$

Dynamic spectral function

$$\mathcal{A}^{\text{dyn}}(m) = \frac{2\mathcal{N}'}{\pi} \frac{m^2 \Gamma^{\text{eff}}(m)}{(m^2 - M_0^2)^2 + m^2 \Gamma^{\text{eff}}(m)^2}$$



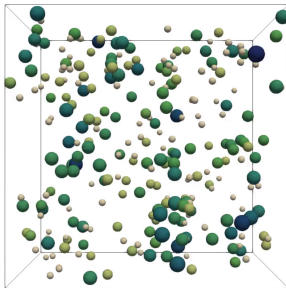
# Thermodynamic behavior

Box

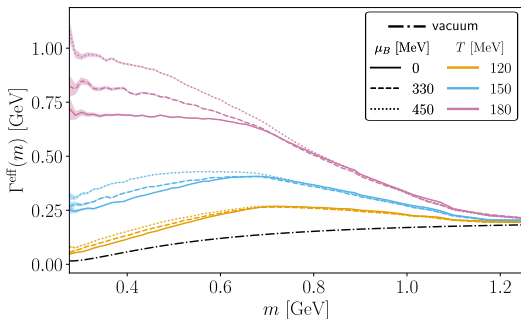
Time: 0 fm

Box Width: 10 fm

Temperature: 0.15 GeV



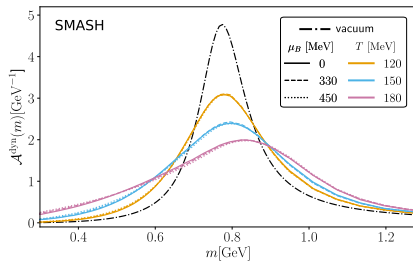
# Thermodynamic behavior



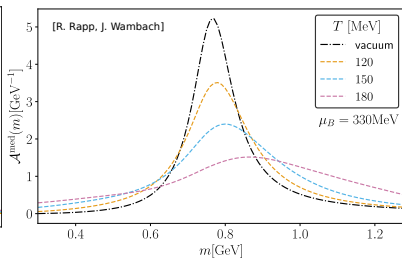
- High-mass  $\rho$  suffers little broadening ( $\sigma_{2 \rightarrow 1} \sim 1/M^4$ )
- Lower masses are mostly absorbed
- Medium temperature changes whole spectrum
- Baryochemical potential affects only  $m \leq M_0$  GeV

# Thermodynamic behavior

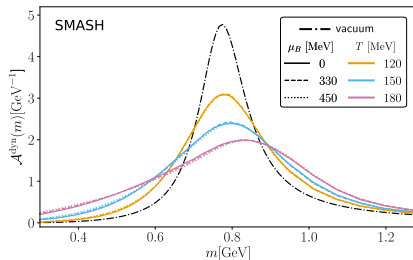
## SMASH



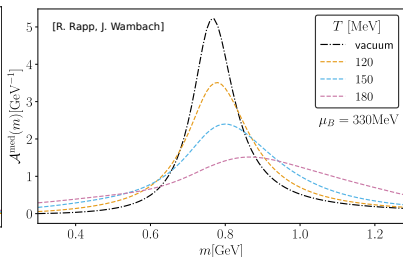
## full in-medium model



## SMASH



## full in-medium model



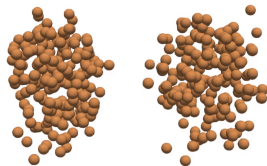
- Melting of  $\rho$  ✓
- Positive shift of peak mass ✓
- Differences: distinct processes present in SMASH ✗  
(intermediate resonances, no self-energy, ...)

# Nuclear collisions

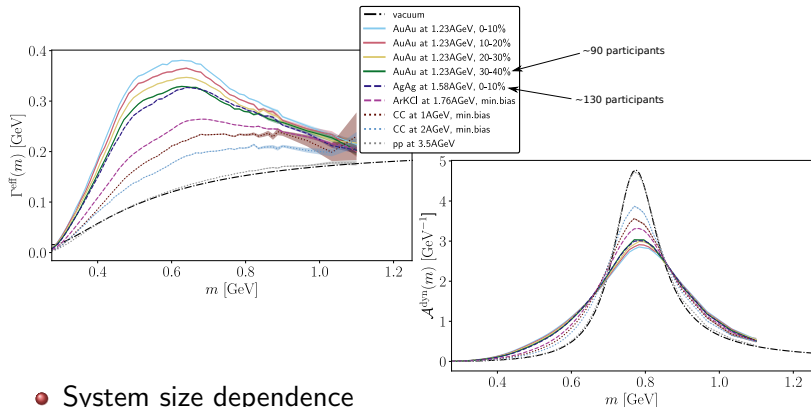
Au+Au at  $E_{\text{kin}} = 1.23A$  GeV

Impact: 0.0 fm

Time: -4 fm



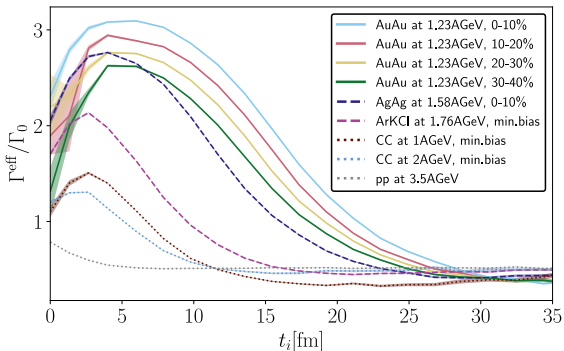
# Nuclear collisions



- System size dependence
- Higher energies disperse the medium
- No broadening at hadronic threshold  $m = m_{2\pi}$

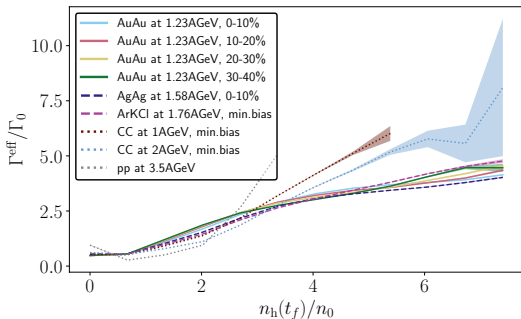


# Nuclear collisions



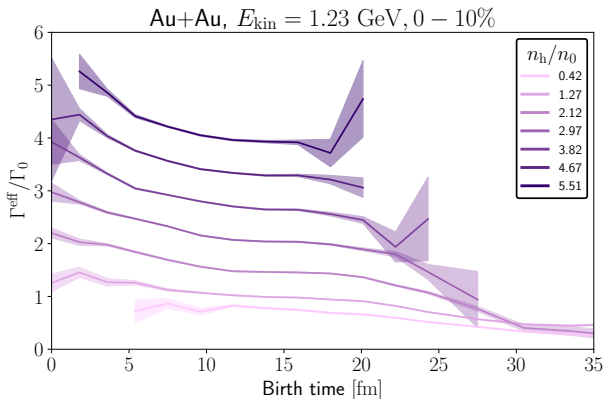
- “Cronometer” for the medium
- Highlights difference between AuAu 1.23A GeV 30-40% and AgAg 1.58A GeV 0-10%

# Nuclear collisions



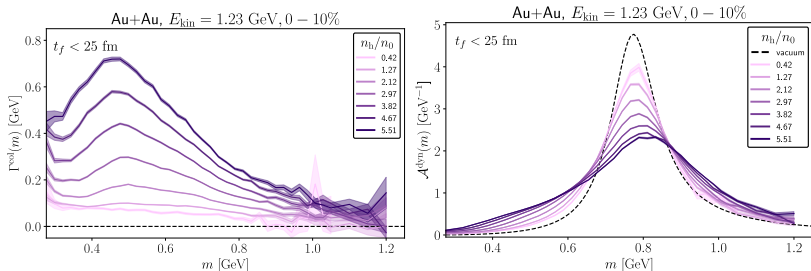
- Near universal dependence on the hadron density  $n_h$
- Deviations for high densities in small systems
- Off-shell models:  $\Gamma^{\text{coll}} = \gamma n_N \langle v \sigma_{VN}^{\text{tot}} \rangle$  (GiBUU, HSD)

# Nuclear collisions



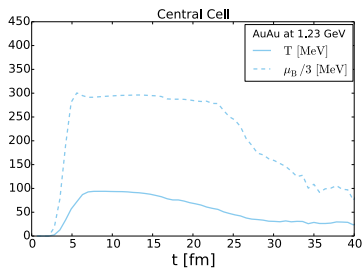
- No dense regions after  $t \approx 25 \text{ fm}$
- Almost constant (chemical composition changes)

# Nuclear collisions

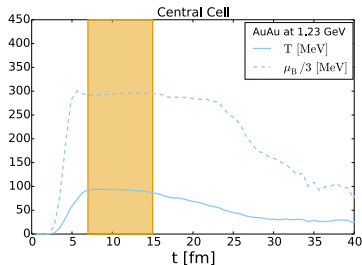


- Peak in  $m \approx 0.5$  GeV appears at high densities
- $\mathcal{A}^{\text{dyn}}$  resembles equilibrated gas

# Non-equilibrium effects

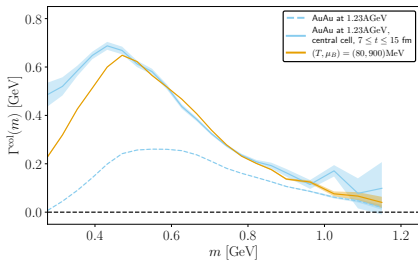
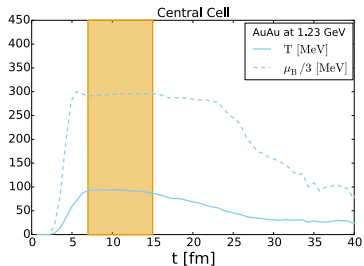


# Non-equilibrium effects



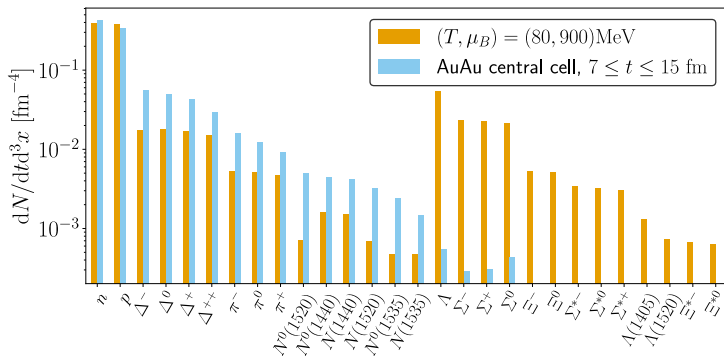
- Spacetime region of a collision where equilibrium is apparent
- Run a box w/ same thermodynamic conditions

# Non-equilibrium effects



- Spacetime region of a collision where equilibrium is apparent
- Run a box w/ same thermodynamic conditions
- Restriction to this region makes  $\Gamma^{\text{col}}(m_{2\pi}) \neq 0$
- Similar broadening, excess for small masses

# Non-equilibrium effects

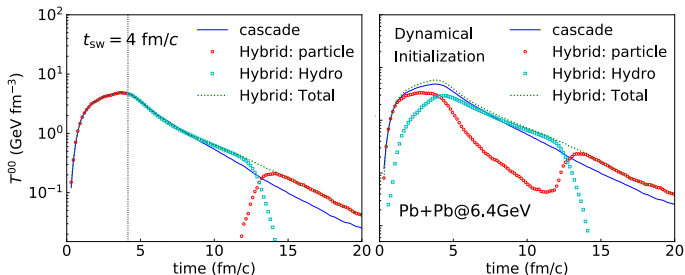


- Equilibrium populated also by strange particles ( $N \propto e^{-m/T}$ )
- Collision system has more particles that can absorb a  $\rho$ -meson



# Future projects

- Dynamic initialization of SMASH+vHLLX hybrid:  
intermediate beam energies
- Direct connection to full in-medium models (no CG)
- Electromagnetic signals of 1<sup>st</sup> order PT



[Akamatsu et al.: PRC 98.2 (2018)]

# Summary

- In SMASH: Vacuum properties  $\xrightarrow{\text{medium}}$  collisional broadening
- Equilibrium  $\mathcal{A}^{\text{dyn}}$  similar to full in-medium calculations
- Collision systems: clear setup dependence (mass number, centrality, beam energy), universality in density
- Effective width works as a cronometer for medium duration
- Out-of-equilibrium may increase collisional broadening

# Summary

- In SMASH: Vacuum properties  $\xrightarrow{\text{medium}}$  collisional broadening
- Equilibrium  $\mathcal{A}^{\text{dyn}}$  similar to full in-medium calculations
- Collision systems: clear setup dependence (mass number, centrality, beam energy), universality in density
- Effective width works as a cronometer for medium duration
- Out-of-equilibrium may increase collisional broadening



T.Hanks for the attention!

# BACKUP SLIDES

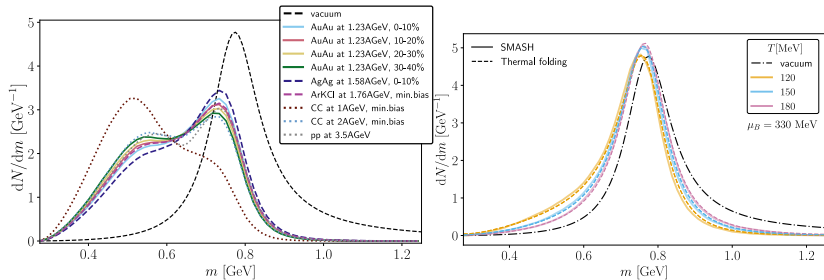
Following Manley et al.,

$$\Gamma_{\rho \rightarrow \pi\pi}^{\text{vac}}(m) = \Gamma^0 \frac{M_0}{m} \left( \frac{\frac{1}{4}m^2 - m_\pi^2}{\frac{1}{4}M_0^2 - m_\pi^2} \right)^{3/2} \left( \frac{\frac{1}{4}M_0^2 - m_\pi^2 + \Lambda^2}{\frac{1}{4}M_0^2 - m_\pi^2 + \Lambda^2} \right)$$

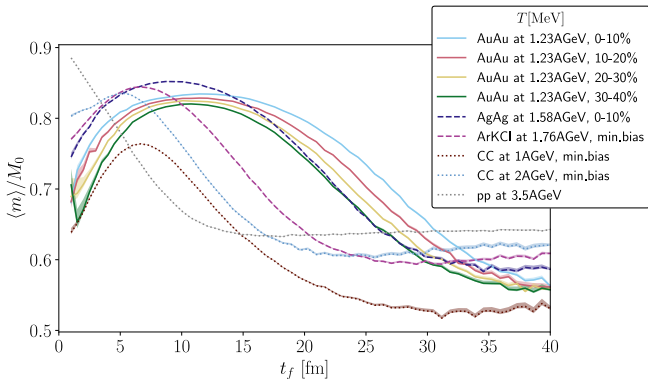
Under the *Vector Meson Dominance* model:

$$\Gamma_{\rho \rightarrow ll}^{\text{vac}}(m) = \Gamma_{\rho \rightarrow ll}^0 \left( \frac{M_0}{m} \right)^3 \left( 1 + \frac{2m_l^2}{m^2} \right) \sqrt{1 - \frac{4m_l^2}{m^2}}$$

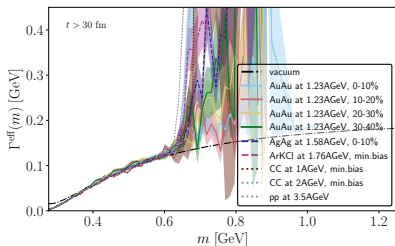
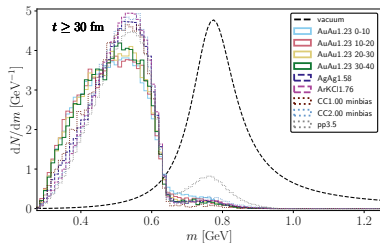
# Produced $\rho$ masses



# Evolution of average mass

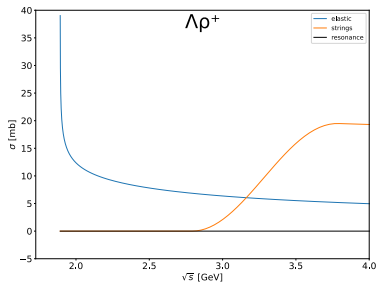
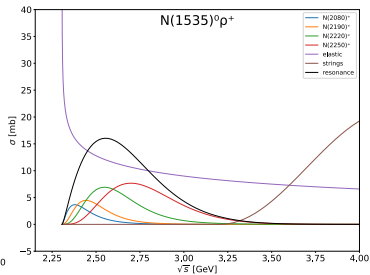
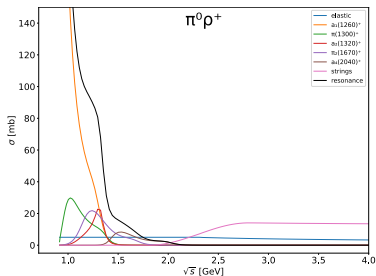


# Late stages of a collision





# $\rho$ -meson interactions



- R. Rapp and J. Wambach, Eur. Phys. J. A6, 415 (1999).
- P. Salabura and J. Stroth, Prog. Part. Nucl. Phys. 120, 103869 (2020).
- H. van Hees and R. Rapp, Nuc. Phys. A, 806 (1-4), 339-387 (2008).
- J. Weil, et al., Phys. Rev. C94 (5), 054905 (2016).
- J. Staudenmaier, et al., Phys. Rev. C98, 054908 (2018).
- D. M. Manley and E. M. Saleski, Phys. Rev. D 45, 4002 (1992).
- H. B. O'Connell, B. C. Pearce, A. W. Thomas, and A. G. Williams, Prog. Part. Nucl. Phys. 39, 201 (1997).