# Percolation and conformal limits in neutron stars







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### Solid Constraints

- Low density:  $\chi EFT$  ( $n \leq 1.1n_0$ ) Tews et al, 2013
- High density pQCD ( $n \gtrsim 40n_0$ ) Gorda et al, 2018

### Interpolation methods

• Polytropes, CSS, Linear Speed of Sound eg. Annala et al, 2018, 2020; Alford et al 2013, 2017, Li et al 2021

Deconfinement by polytropic index

#### $1.75 \rightarrow \text{Hadrons}$ $d\log p$ $\frac{1}{d \log \epsilon} \rightarrow \qquad \gamma < 1.75 \rightarrow \text{Quarks}$

Quark Matter in Neutron Stars?



Energy density (MeV/fm<sup>3</sup>)





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 $5 \times 10^5$  viable Equations of State



TOV Equations P(E)











# TOV Equations





# TOV Equations









## General Structure of Speed of Sound

 ${}^{C}_{s}$ 

#### • General peak-dip structure Altiparmak et al, 2022

#### • Peak similar to quarkyonic matter McLerran, Reddy, 2019; Pang et al, 2023



Local maximum at  $\epsilon_{\text{peak}} = 0.56^{+0.11}_{-0.09}$  GeV/fm<sup>3</sup> with  $c_s^2 = 0.82 \pm 0.08$ 







- Attractive interactions with resonance formation
- Chiral symmetry restoration and deconfinement

Non-monotonicity





- Dominance of repulsive interactions
- Onset of quark or quarkyonic matter?

Change of phase

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## Percolation theory vs speed of sound

see e.g. Satz, 1998; Castorina et al, 2009; Fukushima, 2020





#### Percolation theory: $n_c = 1.22/V_0$

Avg. proton radius:  $R_0 = 0.80 \pm 0.05$  fm Wang et al 2022 Pb-Pb collisions at  $\sqrt{s} = 2.76$  TeV Andronic et al 2018









## Speed of Sound as Trace Anomaly Fujimoto et al 2022



Trace anomaly more informative than speed of sound



## Measure of conformality

### $\Delta$ monotonic up to $\simeq \epsilon_{\rm TOV}$

 $c_s^2 = \frac{-}{3}$ 

## $\triangleleft -0.1$ Maximum in $c_s^2$ Fast approach to comformality



 $- p/\epsilon$ 



 $\Delta \simeq 0$  at  $\epsilon \simeq 1$  GeV/fm<sup>3</sup>  $\leftarrow \varepsilon_{TOV} = 1.16 \pm 0.01$  GeV/fm<sup>3</sup>







Conformality:  

$$c_s^2 = \frac{1}{3} \text{ and } \Delta = 0$$
 $c_s^2 = 0.28 \pm 0.16 \simeq \frac{1}{3}$ 
 $c_{0.4}^{0.4}$ 
 $c_{0.2}^{0.4}$ 
 $\Delta_{\text{TOV}} = -0.01 \pm 0.03 \simeq 0$ 
 $0$ 

Matter almost conformal in the cores of maximally massive NSs

## $c_{\rm s}^2$ and $\Delta$ in Heavy Neutron Stars



 $\Delta = 1/3 - p/\epsilon$ 





14/17

## Incompressiblity & curvature of binding energy MM, K. Redlich, C. Sasaki, to appear

0.2

-0.2

-0.4

$$K \equiv \frac{dP}{dn} = 9\mu(\hat{\alpha} + \hat{\beta}) \qquad 0.4$$

$$\hat{\alpha} \equiv 2 \frac{n}{\mu} \frac{d\epsilon/n}{dn} = 2 \frac{1/3 - \Delta}{4/3 - \Delta} \quad \stackrel{\text{(n)}}{\sim}$$

 $n^2 d^2 \epsilon / n$ 



# $\beta < 0 \rightarrow$ changeover to conformal regime

![](_page_14_Picture_7.jpeg)

![](_page_14_Figure_8.jpeg)

![](_page_14_Picture_9.jpeg)

![](_page_15_Figure_0.jpeg)

# Changeover consistent with percolation

 $K \equiv \frac{dP}{dn} = 9\mu(\hat{\alpha} + \hat{\beta})$ 0.4 $\hat{\alpha} \equiv 2 \frac{n \, d\epsilon / n}{\mu \, dn} = 2 \frac{1/3 - \Delta}{4/3 - \Delta}$ 0.1

 $n^2 d^2 \epsilon / n$  $\begin{array}{c} -0.2^{\bigsqcup}\\ 10^2 \end{array}$  $dn^2$ 

![](_page_16_Figure_3.jpeg)

![](_page_16_Figure_4.jpeg)

![](_page_16_Picture_5.jpeg)

# Summary

## Maximum of $c_s^2$ consistent with percolation threshold

### Matter seems to be conformal in the cores of massive NSs

# Thank You