

FLOW IN MAXWELL FIELD INITIAL VELOCITY FOR COLLECTIVITY

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OUTLINE

- Coarse graining scale and hydrodynamic observables
- Evaluating the flow profile in PHSD
- Quantum Flow from Initial Fields ?
- Coarse-Graining in Wigner Function

 "Success" of hydrodynamics in describing the experimental observations in heavy-ion collisions



P. Kolb and U. Heinz, arXiv:nucl-th/0305084



Oldenburg M.D. (STAR Collab.), J. Phys. G 31, S437

Ideal hydro:

- local thermal equilibrium
- equation of state

collective flow:

- hydro models can reproduce the anisotropic momentum distribution of the final particles
- conservation laws +
 the system behaves collectively (like a strongly interacting liquid)

• "Success" of hydrodynamics in describing the experimental observations in heavy-ion collisions



Event-by-event hydrodynamics:

 2-particle correlation analysis considering inhomogeneous initial condition + hydro evolution reproduces the ridge structure observed experimentally

Some puzzles in Flow...

 It is puzzling that pA and AA data cannot be described with the same set of parameters, since one expects that the same type of fluid is created in both collisions...

> I. Kozlov, M. Luzuma, G. S. Denicol, S. Jeon, Gale C, Signatures of collective behavior in small systems, arXiv:1412.3147v1

Direct Photon Puzzle

Taken from the presentation of K. Reygers,

"Ab initio approaches in many-body QCD confront heavy-ion experiments | December 15, 2014 |

Maybe many more photons from late stage close to *Tc* and hadron gas phase (need large increase in HG rates) [van Hees, He, Rapp, arXiv:1404.2846]

- Theoretical justification?
- Maybe just bremsstrahlung from the HG? ($m+m\rightarrow m+m+\gamma$, $m+B\rightarrow m+B+\gamma$) [Linnyk, Cassing, Bratkovskaya, arXiv:1311.0279]
- Important source in PHSD transport model
- Exotic new photon source, e.g., related to large initial *B* field? [Basar, Kharzeev, Skokov., arXiv:1206.1334]
- seems unlikely to me (centrality dependence, \sqrt{s} dependence, $\sqrt{3}$)
- Initial flow before hydro evolution starts, e.g., IPGlasma model?
- important, but does not address the missing photon yield
- Glasma photons, i.e., large photon production in very early gluon-rich phase? [McLerran, Schenke, arXiv:1403.7462], [Klein-Bösing, McLerran, arXiv:1403.1174]
- promising, but so far based on simplified models
- calculations from first principles needed

Coarse graining scale of hydrodynamic modeling

"... why at all the hydrodynamic approach works so well for such a violent and almost microscopic collisional process?"

(Ph. Mota, et al, Eur. Phys. J A, 48, 165)



S. Pratt et al., Constraining the Equation of State of Superhadronic Matter from Heavy-Ion Collisions, Phys. Rev. Lett. 114 (2015) 202301 (v_2 (Pt-weighted, 2-cenralities), HBT, Spectrum), without viscosity.

Coarse graining scale of hydrodynamic MODELING

"... why at all the hydrodynamic approach works so well for such a violent and almost microscopic collisional process?"

(Ph. Mota, et al, Eur. Phys. J A, 48, 165)

Important questions: J. Berges, J.-P. Blaizot, F. Gelis, J. Phys. G, 39 085115 (2012)

- What is the degree of local thermal equilibrium required for hydrodynamic behavior?
- Can collective flow measurements provide any signal about local equilibration/isotropization?
- What is the coarse graining scale of hydrodynamic modeling for such violent collisional process?

HYDRO IN RELATIVISTIC HEAVY ION COLLISION HOW QUANTITATIVELY PRECISE ?

Uncertainties associated

- EoS, Transport Coefficients (?)
- Freezeout Mechanism (Tough)
- Initial Condition (Challenging)
- Event-by-Event vs. Ensemble Average? (To be clarified)

WE NEED TO KEEP SOME CARE,...

 In a Japanese popular- saying, "Typically in the following three conditions,



- in a twilight, - from far, - half-hidden by a hat

WE NEED TO KEEP SOME CARE,...

 In a Japanese popular- saying, "Typically in the following three conditions,



 - in a twilight, - from far, - half-hidden by a hat make a boy (girl) looking nice,.... "

(We usually see what we WANT to see)

COUNTER-EXAMPLES OF REAL HYDRO ("PSEUDO HYDRO")

- Schrödinger Equation Quantum Hydro
- Isotropic massless gas Non Equilibrium
- Initial state correlation in free streaming case
- Event average -> Effective EoS



Non-equilibrium dynamics – PHSD model

Parton-Hadron-String Dynamics



W. Cassing, E. Bratkovskaya, PRC 78 (2008) 034919;
NPA831 (2009) 215;
W. Cassing, EPJ ST 168 (2009) 3
A. Peshier, W. Cassing, PRL 94 (2005) 172301;
Cassing, NPA 791 (2007) 365: NPA 793 (2007)

microscopic transport description of the partonic and hadronic phase in terms of strongly interacting dynamical quasi-particles and off-shell hadrons Non-equilibrium dynamics – PHSD model

Parton-Hadron-String Dynamics

Initial A+A collisions:

LUND string model

- String formation in primary NN collisions
- String decay to pre-hadrons (B baryons; m mesons)

• Formation of QGP phase: $\varepsilon > \varepsilon_{critical}$



 dissolution of pre-hadrons into massive colored quarks + mean field energy

$$B \rightarrow qqq, m \rightarrow q\overline{q} \quad \forall \quad U_q$$

Dynamical QuasiParticle Model
 (DQPM)

defines quark spectral functions, i.e. masses $M_q(\varepsilon)$ and widths $\Gamma_q(\varepsilon)$ + mean field potential at a given ε (local energy density) (ε related by IQCD EoS to T in the local cell)



Bratkovskaya, PRC 78 (2008) 034919; NPA831 (2009) 215; EPJ ST 168 (2009) 3; NPA856 (2011) 162.

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Non-equilibrium dynamics – PHSD model



Parton-Hadron-String Dynamics

Partonic phase – QGP:



- quarks and gluons = dynamical quasi-particles with off-shell spectral functions (width, mass) defined by DQPM
- ✓ self generated mean field potential for quarks and gluons U_q, U_g from DQPM
- **EoS of partonic phase**: crossover from lattice QCD (fitted by DQPM)
- (quasi-) elastic and inelastic parton-parton interactions using effective cross-section from the DQPM



Hadronization:

 massive, off-shell (anti-)quarks with broad spectral functions hadronize to off-shell mesons and baryons or color neutral excited states – 'strings'

 $g \rightarrow q + \overline{q}, \quad q + \overline{q} \leftrightarrow meson \ ('string')$



PHSD

Parton-Hadron-String Dynamics



W. Cassing & E. Bratkovskaya, NPA 831 (2009) 215

E. Bratkovskaya, W. Cassing, V. Konchakovski, O. Linnyk, NPA856 (2011) 162























$$T^{\mu\nu}(\vec{r},t) = \sum_{i} \frac{p_{i}^{\mu}(t)p_{i}^{\nu}(t)}{p_{i}^{0}(t)} \delta^{3}(\vec{r}-\vec{r_{i}})$$

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$$T^{\mu\nu}(\vec{r},t) = \sum_{i} \frac{p_i^{\mu}(t)p_i^{\nu}(t)}{p_i^0(t)} \delta^3(\vec{r}-\vec{r}_i)$$

Replace the delta by a kernel function:

$$\begin{split} \delta^3(\vec{r}-\vec{r_i}) &\to W(\vec{r}-\vec{r_i}(t);\Delta\vec{r}) \\ \text{rectangular} \begin{cases} W=k & \text{inside a box} \\ W=0 & \text{outside a box} \end{cases} \end{split}$$



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$$\left(\int W(\vec{r};\Delta\vec{r})d^3\vec{r}=1\right)$$

> Define a smoothing (gaussian) kernel:

$$W(\vec{r} - \vec{r_i}; h_T, h_L) = A e^{-\frac{1}{2} \left[\left(\frac{x - x_i}{h_T} \right)^2 + \left(\frac{y - y_i}{h_T} \right)^2 + \left(\frac{z - z_i}{h_L} \right)^2 \right]}$$

$$h_T \to \text{transverse scale} \qquad \qquad h_L \to \text{longitudinal scale}$$

 Diagonalizing the energy-momentum tensor (solve the eigenvalue/eigenvector problem)



- > The four velocity u^{μ} can be identified with the eigenvector associated with the eigenvalue ε $T^{\mu\nu}u_{\nu} = \varepsilon u^{\mu}$ (time-like eigenvector)
- The flow profile:

 $\vec{\beta}, \varepsilon, P$

"Rectangular box" vs "Gaussian kernel"



PRESSURE COMPONENTS (SPATIAL EIGENVALUES)



MEAN FIELD EFFECTS

• Varying the number of parallel events in PHSD



MEAN FIELD EFFECTS

• Varying the number of parallel events in PHSD



COARSE-GRAINING SCALE

• Varying cell size in the transverse plane



COARSE-GRAINING SCALE

• Varying cell size in the transverse plane


COARSE-GRAINING SCALE

• Varying cell size in the transverse plane



COARSE-GRAINING SCALE

• Varying cell size in the longitudinal direction



COARSE-GRAINING SCALE

• Varying cell size in the longitudinal direction



Transverse and Longitudinal position dependence

- Varying the position of the cell along x direction
- Varying the position of the cell along z direction



Transverse and Longitudinal position dependence

- Varying the position of the cell along x direction
- Varying the position of the cell along z direction



INITIAL ECCENTRICITIES AND FLOW COEFFICIENTS

spatial anisotropy

momentum anisotropy





Note that Ψ_n is not the event plane angle usually obtained experimentally, which is the final momentum event plane







Flow coefficients seem to reach asymptotic values around 5-6 fm/c



o 2nd harmonic:



ECCENTRICITY -> FLOW IN PHSD • 3rd harmonic: (Event average) (event-by-event) 0.06 NUM=10 ---ε₃ (×0.17) (b)=5.7 [fm] NUM = 100.3 0.05 ---- V₃ event-by-event mean 0.04 0.2 0.03 0.02 0.1 0.01

-0.1

0.2

0.1

0.3

t [fm]

8

6

12

14

16

10

ν_n, ε_n

-0.01

-0.02

n

2

ε3

0.5

0.6

0.7

0.4

0.9

0.8

• 4th harmonic:







PARTICLE VS FIELD PRESSURE





• T. Epelbaum, QM2014

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• T. Epelbaum, F. Gelis, PRL 111, 232301 (2013)



PARTIAL SUMMARY

 PHSD model provides a convenient way to test the coarse-graining scale of hydrodynamics within a scenario of microscopic dynamics

- So far, we have observed a clear separation of the "flow profile" into longitudinal and transverse components
- Only for some very specific situations the system evolution seems to approach "equilibrium"
- On the other hand, event average eccentricities/flow coefficients seem to follow the hydrodynamic behavior in PHSD

Coarse graining may depend on Observables

🗸 Particle vs. Field

NUM=1



NUM=2



NUM=10



WEYL PROJECTION, WIGNER FUNCTION, HUSIMI FUNCTION AND COARSE GRAINING

Weyl Projection Operator

$$\hat{\Delta}(q,p) = \frac{1}{2\pi\hbar} \int du dv \, e^{i\left\{u(q-\hat{Q})+v(p-\hat{P})\right\}/\hbar},$$
(generalization of $|q\rangle\langle q| \equiv \hat{\delta}(q-\hat{Q}) = \frac{1}{2\pi\hbar} \int du \, e^{i\left\{u(q-\hat{Q})\right\}/\hbar}$)
$$Tr\left[\hat{\Delta}(q,p)\right] = 1,$$

$$Tr\left[\hat{\Delta}(q,p)\hat{\Delta}(q',p')\right] = 2\pi\hbar\delta(q-q')\delta(p-p'),$$

$$O_{W}(q,p) = Tr\left[\hat{\Delta}(q,p)\hat{O}\right] \rightleftharpoons \hat{O} = \frac{1}{2\pi\hbar} \int dq dp O_{W}(q,p)\hat{\Delta}(q,p)$$

$$Tr\left[\hat{A}\hat{B}\right] = \frac{1}{2\pi\hbar} \int dq dp A_{W}(q,p) B_{W}(q,p)$$
⁵⁸

WEYL PROJECTION, WIGNER FUNCTION, HUSIMI FUNCTION AND COARSE GRAINING

Density Matriz

Pure state

 $\hat{\rho} = |\Psi\rangle\langle\Psi|$

Mixed state

 $\hat{\rho} = \sum_{\alpha} \omega_{\alpha} |\Psi_{\alpha}\rangle \langle \Psi_{\alpha}|$

Wigner Function

$$f_{W}(q,p) = Tr\left[\hat{\Delta}(q,p)\hat{\rho}\right] \iff \hat{\rho} = \frac{1}{2\pi\hbar}\int dqdp f_{W}(q,p)\hat{\Delta}(q,p)$$
$$\langle O \rangle = Tr\left[\hat{\rho}\hat{O}\right] = \frac{1}{2\pi\hbar}\int dqdp f_{W}(q,p)O_{W}(q,p)$$
$$f_{W}(q,p) = \frac{1}{2\pi\hbar}\int du \, e^{iuq/\hbar} \left\langle p + \frac{1}{2}u \left| \hat{\rho} \right| p - \frac{1}{2}u \right\rangle$$
$$= \int dv \, e^{ivp/\hbar} \left\langle q - \frac{1}{2}v \left| \hat{\rho} \right| q + \frac{1}{2}v \right\rangle$$
$$= \int dv \, e^{ivp/\hbar} \psi \left(q - u/2 \right) \psi^{*} \left(q + u/2 \right)$$

WEYL PROJECTION, WIGNER FUNCTION, HUSIMI FUNCTION AND COARSE GRAINING

$$\langle O \rangle = Tr \left[\hat{\rho} \hat{O} \right] = \frac{1}{2\pi\hbar} \int dq dp f_w(q, p) O_w(q, p)$$

Wigner Function is not necessarily positive.

Husimi Function Smoothing (coarse graining in q and p wih Gaussian weight,

$$f_{H}(q,p) = \int dq' dp' G(q-q';h) G\left(p-p';\frac{\hbar}{2h}\right) f_{W}(q,p)$$
$$= \langle q,p | \hat{\rho} | q,p \rangle$$

Where

$$G(x,h) = \frac{h}{\sqrt{\pi}} e^{-\left(\frac{x}{h}\right)^2}$$

and $|q, p\rangle$ is the coherent state.

WEYL PROJECTION, WIGNER FUNCTION, HUSIMI FUNCTION AND COARSE GRAINING

Still valid

$$\langle O \rangle = Tr \left[\hat{\rho} \hat{O} \right] = \frac{1}{2\pi\hbar} \int dq dp f_H(q, p) O_H(q, p)$$

In the limit of, for example, $h \rightarrow \infty$

$$G(p,\hbar/2h) \rightarrow \delta(p),$$

so that

$$f_H(q,p) \rightarrow |\varphi(p)|^2 = |\langle p | \psi \rangle|^2,$$

and the corresponding density matrix reconstructed becomes $\hat{\rho}_{rec} = \int dp |\varphi(p)|^2 |p\rangle \langle p|$

which is a mixed state! ...

P. Carruthers and F. Zachariasen Rev. Mod. Phys., Vol. 55, No. 1, 1983

$$i\hbar\partial_{t}\psi(\vec{x},t) = -\frac{\hbar^{2}}{2m}\nabla^{2}\psi(\vec{x},t) + V(\vec{x})\psi(\vec{x},t)$$

$$\partial_{t}\rho + \nabla \cdot (\rho\vec{u}) = 0,$$

$$\partial_{t}(\rho u_{i}) + \partial_{j}p_{ij} = -\frac{1}{m}\partial_{i}V$$

$$\rho = \int d^{3}x f(\vec{x},\vec{p},t)$$
with
$$\rho\vec{u} = \int d^{3}p \frac{\vec{p}}{m}f(\vec{x},\vec{p},t),$$

$$p_{ij} = \int d^{3}p \frac{p_{i}p_{j}}{m^{2}}f(\vec{x},\vec{p},t).$$

P. Carruthers and F. Zachariasen Rev. Mod. Phys., Vol. 55, No. 1, 1983

This is equivalent to that of E. Madelung, 1926, Z. Phys. 40, 322.

$$i\hbar\partial_{t}\psi\left(\vec{x},t\right) = -\frac{\hbar^{2}}{2m}\nabla^{2}\psi\left(\vec{x},t\right) + V\left(\vec{x}\right)\psi\left(\vec{x},t\right)$$

$$\partial_{t}\rho + \nabla\cdot\left(\rho\vec{u}\right) = 0,$$

$$\partial_{t}\vec{u} + (\vec{u}\cdot\nabla)\vec{u} = -\frac{1}{m}\nabla V - \frac{1}{\rho}\nabla\left(\frac{1}{\sqrt{\rho}}\nabla^{2}\ln\sqrt{\rho}\right)$$
with
$$\rho = |\psi\left(\hat{x},t\right)|^{2}$$

$$\vec{u} = \frac{\hbar}{u}\operatorname{Im}(\ln\psi\left(\hat{x},t\right))$$

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Quantum pressure
$$\partial_{t}\rho + \nabla\cdot\left(\rho\vec{u}\right) = 0,$$

$$\partial_{t}\vec{u} + \left(\vec{u}\cdot\nabla\right)\vec{u} = -\frac{1}{m}\nabla V - \frac{1}{\rho}\nabla\left(\frac{1}{\sqrt{\rho}}\nabla^{2}\ln\sqrt{\rho}\right)$$
with
$$\rho = |\psi\left(\hat{x},t\right)|^{2}$$

$$\vec{u} = \frac{\hbar}{m}\operatorname{Im}(\ln\psi\left(\hat{x},t\right))$$

$$f(x,t) = \frac{1}{\rho}\left(\frac{1}{\sqrt{\rho}}\nabla^{2}\ln\sqrt{\rho}\right)$$

P. Carruthers and F. Zachariasen Rev. Mod. Phys., Vol. 55, No. 1, 1983

E. Madelung, 1926, Z. Phys. 40, 322.

Takabayashi-Wallstrom $i\hbar\partial_t\psi(\vec{x},t) = -\frac{\hbar^2}{2m}\nabla^2\psi(\vec{x},t) + V(\vec{x})\psi(\vec{x},t)$ $\partial_t \rho + \nabla \cdot (\rho \vec{u}) = 0,$ $\partial_t \vec{u} + (\vec{u} \cdot \nabla)\vec{u} = -\frac{1}{m}\nabla V - \frac{1}{\rho}\nabla \left(\frac{1}{\sqrt{\rho}}\nabla^2 \ln \sqrt{\rho}\right)$ $\rho = |\psi(\hat{x},t)|^2$ with 65 $\vec{u} = \frac{\hbar}{\ln(\ln\psi(\hat{x},t))}$

When exists a large inhomogeniety, the quantum pressure may generate a collective flow...

P. Carruthers and F. Zachariasen Rev. Mod. Phys., Vol. 55, No. 1, 1983

$$i\hbar\partial_{t}\psi(\vec{x},t) = -\frac{\hbar^{2}}{2m}\nabla^{2}\psi(\vec{x},t) + V(\vec{x})\psi(\vec{x},t)$$

$$\partial_{t}\rho + \nabla \cdot (\rho\vec{u}) = 0,$$

$$\partial_{t}(\rho u_{i}) + \partial_{j}p_{ij} = -\frac{1}{m}\partial_{i}V$$

$$\rho = \int d^{3}x f(\vec{x},\vec{p},t)$$
with
$$\rho\vec{u} = \int d^{3}p \frac{\vec{p}}{m}f(\vec{x},\vec{p},t),$$

$$p_{ij} = \int d^{3}p \frac{p_{i}p_{j}}{m^{2}}f(\vec{x},\vec{p},t).$$

Peter Holland, Proc. R. Soc. A 2005 461,

1st step:

rewrite the Maxwell Eqs.

$$\nabla \times \vec{E} = -\partial_t \vec{B}, \qquad \nabla \times \vec{B} = \frac{1}{c^2} \partial_t \vec{E},$$
$$\nabla \cdot \vec{E} = 0, \qquad \nabla \cdot \vec{B} = 0.$$

$$i\hbar\partial_t \vec{F} = -ic\left(\vec{\Sigma}\cdot\nabla\right)\vec{F}$$
$$\vec{F} \equiv \sqrt{\frac{\varepsilon_0}{2}}\left(\vec{E} + ic\vec{B}\right)$$

as

where

and
$$\Sigma$$
 is (3x3) matrix,
corresponding to the Pauli
matrix for spin 1, satisfying

together with the initial condition

$$\vec{\Sigma} \times \vec{\Sigma} = i\hbar\vec{\Sigma}$$

$$\nabla \cdot \vec{F}(t=0) = 0.$$

Peter Holland, Proc. R. Soc. A 2005 461,

2nd step:

where

rewrite this Dirac-like Eqs. in the Cartesian base into the spherical base,

$$\begin{split} i\hbar\partial_t \vec{F} &= -ic\left(\vec{\Sigma}\cdot\nabla\right)\vec{F} \\ i\hbar\partial_t \vec{G} &= -ic\left(\vec{J}\cdot\nabla\right)\vec{G}, \\ \vec{G} &= \begin{pmatrix} G_{+1} \\ G_0 \\ G_{-1} \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} -F_x + iF_y \\ \sqrt{2}F_z \\ F_x + iF_y \end{pmatrix} \end{split}$$

and J is the matrix given by

$$J_{x} = \frac{\hbar}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \ J_{y} = \frac{\hbar}{\sqrt{2}} \begin{pmatrix} 0 & -i & 0 \\ i & 0 & -i \\ 0 & i & 0 \end{pmatrix}, \ J_{z} = \hbar \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & -1 \end{pmatrix}$$
⁶⁸

MAXWELL EQUATION IN SCHRÖDINGER FORM P. Holland, Proc. R. Soc. A 2005 461,

3rd step:

Introduce the derivative operator whose matrix elements coincide with those of J, $J_x \rightarrow \hat{M}_x = i\hbar \left(\cos\beta\partial_{\alpha} - \sin\beta\cot\alpha\partial_{\beta} + \sin\beta\csc\alpha\partial_{\gamma}\right),$ $J_y \rightarrow \hat{M}_y = i\hbar \left(-\sin\beta\partial_{\alpha} - \cos\beta\cot\alpha\partial_{\beta} + \cos\beta\csc\alpha\partial_{\gamma}\right),$ $J_z \rightarrow \hat{M}_z = i\hbar\partial_{\beta},$ where $\alpha = (\alpha, \beta, \gamma)$ are Euler angles (in the convention of Holland) in the sense that

with

$$\int d\Omega \ u_{m}^{*}(\vec{\alpha}) \hat{M}_{i}u_{m}(\vec{\alpha}) = (J_{i})_{mn}, \ m, n = -1, 0, 1$$

$$u_{1}(\vec{\alpha}) = \frac{\sqrt{3}}{4\pi} \sin \alpha \ e^{-i\beta}, \ u_{0}(\vec{\alpha}) = i\sqrt{\frac{3}{2}} \frac{1}{2\pi} \cos \alpha, \ u_{-1}(\vec{\alpha}) = \frac{\sqrt{3}}{4\pi} \sin \alpha \ e^{i\beta}$$
(69)

and

 $d\Omega = \sin \alpha d\alpha d\beta d\gamma, \quad \alpha \in [0, \pi], \ \beta, \gamma \in [0, 2\pi]$

P. Holland, Proc. R. Soc. A 2005 461,

4th step:

Introduce the scalar wavefunction

$$\psi(x,\vec{\alpha},t) = \sum_{m=1,0,-1} G_m(x,t) u_m(\alpha)$$

the Schrödinger equation in R³ X SO(3) space,

$$i\hbar\partial_t\psi(x,\vec{\alpha},t) = -ic\sum_i \hat{M}_i\partial_i\psi(x,\vec{\alpha},t)$$

reduces to the Maxwell equation.

P. Holland, Proc. R. Soc. A 2005 461,

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reduces to the Maxwell equation.

Hydrodynamic form ?

$$\rho = \int d^{3}x f\left(\vec{x}, \vec{p}, t\right)$$

$$\rho \vec{u} = \int d^{3}p \, \frac{\vec{p}}{m} f\left(\vec{x}, \vec{p}, t\right),$$

$$p_{ij} = \int d^{3}p \, \frac{P_{i}P_{j}}{m^{2}} f\left(\vec{x}, \vec{p}, t\right).$$

P. Holland, Proc. R. Soc. A 2005 461,

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$$i\hbar\partial_t\psi(x,\vec{\alpha},t) = -ic\sum_i \hat{M}_i\partial_i\psi(x,\vec{\alpha},t)$$

reduces to the Maxwell equation.....

Particle interpretation with the quantum pressure ?
P. Holland, Proc. R. Soc. A 2005 461,

Starting point:

Schrödinger equation in R³ X SO(3) space, with second order differencial in the space, $t \rightarrow 0$ $(t \rightarrow t)$ $t \rightarrow 0$ $(t \rightarrow t)$

$$i\hbar\partial_t\psi(x,\vec{\alpha},t) = -\hbar c \sum_i \lambda_i \partial_i\psi(x,\vec{\alpha},t)$$

where $\hat{\lambda}_i = i\hbar \hat{M}_i$ are real differential operators with respect to the three Euler angles $\vec{\alpha}$.

Energy density:

$$\frac{\varepsilon_0}{2} \left(\vec{E}^2 + c^2 \vec{B}^2 \right) = \int d^3 \Omega \left| \psi(\vec{x}, \vec{\alpha}) \right|^2,$$

and the Poynting vector

$$\frac{\varepsilon_0}{2} \left(\vec{E}^2 + c^2 \vec{B}^2 \right) = \int d^3 \Omega \left| \psi(\vec{x}, \vec{\alpha}) \right|^2,$$

P. Holland, Proc. R. Soc. A 2005 461,

Second step:

Following Madelung, write

$$\psi(x,\vec{\alpha},t)=\sqrt{\rho}\,e^{iS/\hbar},$$

Then we get

$$\frac{\partial S}{\partial t} + \frac{c}{\hbar} \left(\vec{\hat{\lambda}} \cdot \nabla_x \right) S + Q = 0,$$

and

$$\frac{\partial \rho}{\partial t} + \frac{c}{\hbar} \nabla_x \cdot \left(\rho \vec{\hat{\lambda}}\right) S + \frac{c}{\hbar} \vec{\hat{\lambda}} \cdot \left(\rho \nabla S\right) = 0.$$

where

$$Q = -c\hbar \left(\vec{\hat{\lambda}} \cdot \nabla_x\right) \ln \sqrt{\rho},$$

P. Holland, Proc. R. Soc. A 2005 461,

Like Schrödinger Equation's case, we need Takabayashi-Wallstrom constraints,

$$i\hbar\partial_t\psi(x,\vec{\alpha},t) = -\hbar c \sum_i \hat{\lambda}_i \partial_i\psi(x,\vec{\alpha},t)$$

$$\begin{aligned} \frac{\partial S}{\partial t} &+ \frac{c}{\hbar} \left(\hat{\hat{\lambda}} \cdot \nabla_x \right) S + Q = 0, \\ \frac{\partial \rho}{\partial t} &+ \frac{c}{\hbar} \nabla_x \cdot \left(\rho \hat{\hat{\lambda}} \right) S + \frac{c}{\hbar} \hat{\hat{\lambda}} \cdot \left(\rho \nabla S \right) = 0, \\ \psi \left(x, \vec{\alpha}, t \right) &= \sqrt{\rho} \, e^{iS/\hbar}, \quad Q = -c\hbar \left(\hat{\hat{\lambda}} \cdot \nabla_x \right) \ln \sqrt{\rho}, \end{aligned}$$

P. Holland, Proc. R. Soc. A 2005 461,

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 $|i\hbar\partial_t\psi(x,\vec{\alpha},t)| = -\hbar c \sum \hat{\lambda}_i \partial_i\psi(x,\vec{\alpha},t)$ $\oint \nabla_{\xi} S\left(\vec{\xi},t\right) \cdot d\vec{\xi} = n\hbar, \quad n \in \mathbb{Z}, \ \vec{\xi} \in \left\{R^3 \otimes \Omega^3\right\}$ $\frac{\partial S}{\partial t} + \frac{c}{\hbar} \left(\vec{\lambda} \cdot \nabla_x \right) S + Q = 0 \quad \mathbf{D} \mathbf{C} \quad \mathbf{Q} \quad$

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- Maxwell's Equations can be written in the form of hydrodynamic flow in R³ X SO(3) space with the quantum pressure (which is not visible in the original equation).
- 2. Energy-momentum tensor can be obtained as the coarse grained over the angular state (What happens for the generalized Husimi states?)
- 3. Circulations are quantized. Vortexes in polarization.
- 4. How to introduce the Gauge transformation?

SUMMARY

- How to describe "Particulization" from the Intense Initial Field consistently ? (Analogy to Quantum Optics)
- Effects of Initial Velocity Field for the Initial Condition ?
- To find the initial condition for hydromodel, several levels of Coarse Graining is necessary (Total wavefunction to the single-particle states, localization in space, etc). All of them introduce the mixed states and may affect the momentum distribution,...
- Can non-Abelian field be described in the form of Schrödinger form? Inhomonegeities in Quantum Pressure affects the initial momentum distribution?

In 60's...

Where have all the flowers gone...











And thank you for Helea, Jörg and Marcus ...