

The global $U(2)_L \times U(2)_R$ Linear σ -Model with Vector and Axialvector Fields and Electroweak Interaction

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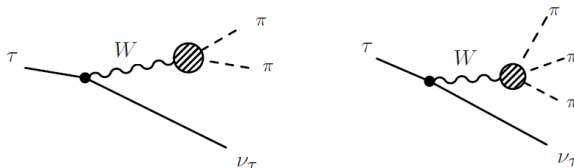
- 1 Introduction
- 2 Vector Spectral Density
- 3 Axialvector Spectral Density
- 4 Conclusion and Outlook

Motivation

- τ spectral functions from ALEPH collaboration
ALEPH Collaboration, S. Schael et al., Branching Ratios and Spectral Functions of Tau Decays: Final ALEPH measurements and physics implications, Phys.Rept. 421 (2005) 191284, [hep-ex/0506072]
- determine electroweak interactions for hadronic degrees of freedom in the vacuum
- results can be used to perform calculations at nonzero temperature and density (e.g. dilepton decay rate)
- understand the nature of resonances like a_1 , e.g. $\bar{q}q$ or $\rho\pi$ -state?

τ -Decay

- weak semileptonic decay:
decay rate $\Gamma_{\tau \rightarrow W \nu_\tau}$ well known from Standard Model of Particle Physics



- a hadronic model with electroweak interaction is needed

- Linear σ Model with global $U(2)_L \times U(2)_R$ chiral symmetry
- hadronic degrees of freedom are the light $N_f = 2$ meson multiplets, represented by the matrix-valued fields Φ , L^μ , R^μ
 - Φ : scalar S (σ, \vec{a}_0) , pseudoscalar P $(\eta, \vec{\pi})$,
 - $L^\mu(R^\mu)$: vector $V^\mu (\omega^\mu, \vec{\rho}^\mu)$, axialvector $A^\mu (f_1^\mu, \vec{a}_1^\mu)$.
- Strong isospin algebra with generators

$$t_0 = \frac{1}{2} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \quad t_i = \frac{\sigma_i}{2}, \quad i = 1, 2, 3$$

- τ -decay can be calculated within linear σ -model after including electroweak interactions

Representation of the Fields

S scalar, P pseudoscalar:

$$\begin{aligned}\Phi &= \sum_{a=0}^3 (S_a + iP_a) t_a \left(\sim \sqrt{2} \sum_{a=0}^3 (\bar{q} t_a q + \bar{q} t_a \gamma_5 q) t_a \right) \\ &= \frac{1}{2} \begin{pmatrix} \sigma + a_0^0 + i(\eta + \pi^0) & a_0^+ + i\pi^+ \\ a_0^- + i\pi^- & \sigma - a_0^0 + i(\eta - \pi^0) \end{pmatrix}\end{aligned}$$

V_μ vector, A_μ axialvector:

$$\begin{aligned}(L, R)_\mu &= \sum_{a=0}^3 (V_\mu^a \pm A_\mu^a) t_a \left(\sim \sqrt{2} \sum_{a=0}^3 (\bar{q} \gamma_\mu t_a q \pm \bar{q} t_a \gamma_\mu \gamma_5 q) t_a \right) \\ &= \frac{1}{2} \begin{pmatrix} (\omega_\mu + \rho_\mu^0) \pm (f_{1\mu} + a_{1\mu}^0) & \rho_\mu^+ \pm a_{1\mu}^+ \\ \rho_\mu^- \pm a_{1\mu}^- & (\omega_\mu - \rho_\mu^0) \pm (f_{1\mu} - a_{1\mu}^0) \end{pmatrix}\end{aligned}$$

Lagrangian

$$\begin{aligned}
 \mathcal{L} = & Tr[(D^\mu \Phi)^\dagger (D^\mu \Phi)] - m^2 Tr(\Phi^\dagger \Phi) - \lambda_1 [Tr(\Phi^\dagger \Phi)]^2 - \lambda_2 Tr(\Phi^\dagger \Phi)^2 \\
 & - \frac{1}{4} Tr[(L_0^{\mu\nu})^2 + (R_0^{\mu\nu})^2] + \frac{m_1^2}{2} Tr[(L^\mu)^2 + (R^\mu)^2] + Tr[H(\Phi + \Phi^\dagger)] \\
 & + c(\det \Phi + \det \Phi^\dagger) + \text{higher order couplings of } \Phi, L^\mu, R^\mu
 \end{aligned}$$

$$\begin{aligned}
 D_\mu \Phi &= \partial_\mu \Phi - ig_1(L_\mu \Phi - \Phi R_\mu), \\
 (L, R)_0^{\mu\nu} &= \partial^\mu (L, R)^\nu - \partial^\nu (L, R)^\mu
 \end{aligned}$$

1 invariant under global $U(2)_L \times U(2)_R$ transformations

$$\begin{aligned}
 \Phi &\rightarrow \Phi' = U_L \Phi U_R^\dagger \\
 L^\mu &\rightarrow L^{\mu'} = U_L L^\mu U_L^\dagger \\
 R^\mu &\rightarrow R^{\mu'} = U_R R^\mu U_R^\dagger
 \end{aligned}$$

2 symmetry is broken to $U(1)_V$, baryon number conservation

$SU(2)_L \times U(1)_Y$ Transformation

- it turns out that $SU(2)_L \times U(1)_Y$ is simply a subgroup of $U(2)_L \times U(2)_R$
- transformation laws

$$\begin{aligned}
 \Phi &\longrightarrow \Phi' &= U_L \Phi U_Y^\dagger \\
 L^\mu &\longrightarrow L^{\mu'} &= U_L L^\mu U_L^\dagger, \\
 R^\mu &\longrightarrow R^{\mu'} &= U_Y R^\mu U_Y^\dagger, \\
 B^\mu &\longrightarrow B^{\mu'} &= U_Y B^\mu U_Y^\dagger + \frac{i}{g'} U_Y \partial^\mu U_Y^\dagger, \\
 W^\mu &\longrightarrow W^{\mu'} &= U_L W^\mu U_L^\dagger + \frac{i}{g} U_L \partial^\mu U_L^\dagger.
 \end{aligned}$$

- covariant derivative

$$D^\mu \Phi = \partial^\mu \Phi - ig_1 (L^\mu \Phi - \Phi R^\mu) + ig' \Phi B^\mu - ig W^\mu \Phi.$$

Weinberg Mixing and Cabibbo Mixing

- neutral bare $SU(2)_L \times U(1)_Y$ gauge fields B^μ , W_3^μ are related to the physical fields A^μ , Z^μ by

$$\begin{pmatrix} W_3^\mu \\ B^\mu \end{pmatrix} = \begin{pmatrix} \cos\theta_W & \sin\theta_W \\ -\sin\theta_W & \cos\theta_W \end{pmatrix} \begin{pmatrix} Z^\mu \\ A^\mu \end{pmatrix}$$

and

$$e = g' \cos\theta_W = g \sin\theta_W$$

- strong isospin eigenstates of d , s , b are related to the weak eigenstates by the CKM matrix
- Cabibbo mixing

$$\begin{pmatrix} d' \\ s' \end{pmatrix} = \begin{pmatrix} \cos\theta_C & \sin\theta_C \\ -\sin\theta_C & \cos\theta_C \end{pmatrix} \begin{pmatrix} d \\ s \end{pmatrix}$$

- covariant derivative in terms of physical weak interaction fields and weak eigenstates

$$D^\mu \Phi = \partial^\mu \Phi - ig_1(L^\mu \Phi - \Phi R^\mu) - ig \cos \theta_C (W_1^\mu t_1 + W_2^\mu t_2) \Phi - ieA^\mu [t_3, \Phi] - ig \cos \theta_W (Z^\mu \Phi + \tan^2 \theta_W \Phi Z^\mu)$$

- also redefinition of the field strength tensors for left- and righthanded fields

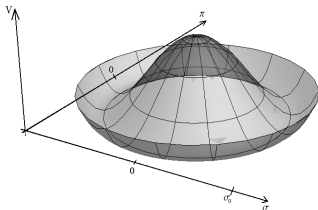
$$L^{\mu\nu} = \frac{i}{g} [D_L^\mu, D_L^\nu], \quad D_L^\mu = \partial^\mu - ig W^\mu \cdot *$$

and

$$R^{\mu\nu} = \frac{i}{g'} [D_R^\mu, D_R^\nu], \quad D_R^\mu = \partial^\mu - ig * \cdot B^\mu$$

Spontaneous Symmetry Breaking

- for $m^2 < 0$ we obtain infinite number of degenerate groundstates



$\sigma \rightarrow \sigma + \phi_0$ generates mass shift
between σ and π -triplet

π massless Goldstone bosons

- explicit symmetry breaking term $\mathcal{L}_{eSB} = Tr[H(\Phi + \Phi^\dagger)]$
renders π -triplet massive

- gives rise to vertices proportional to ϕ_0 , ϕ_0^2 and generates mixing terms between weak bosons and vector and axialvector fields
- we encounter two problems:
 - a_1 - π -mixing term $-g_1\phi_0\vec{a}_1^\mu \cdot \partial_\mu\vec{\pi}$ requires shifting \vec{a}_1^μ and renormalizing $\vec{\pi}$

$$\left. \begin{array}{l} \vec{a}_{1\mu} \rightarrow \vec{a}_{1\mu} + Zw\partial_\mu\vec{\pi} \\ \vec{\pi} \rightarrow Z\vec{\pi} \end{array} \right\} \text{eliminate } a_1\text{-}\pi\text{-mixing}$$

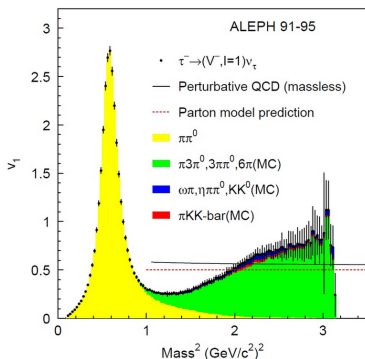
- model does not contain W_ρ mixing terms, therefore introduce

$$\mathcal{L}_{W\rho} = \frac{\delta}{2}g \cos\theta_C \text{Tr}[W_{\mu\nu}L^{\mu\nu}]$$

ALEPH Inclusive Spectral Functions

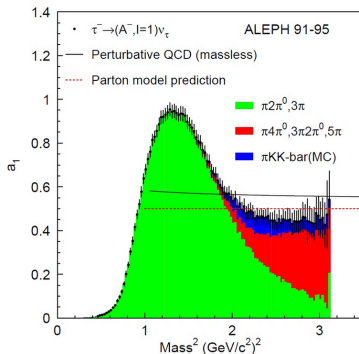
vector channel

$$\tau \rightarrow \nu_\tau 2\pi \nu_\tau (4\pi \nu_\tau)$$



axialvector channel

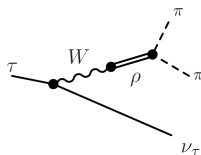
$$\tau \rightarrow \nu_\tau 3\pi \nu_\tau (5\pi \nu_\tau)$$



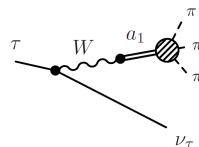
ALEPH Collaboration / Physics Reports 421 (2005) 191–284

dominant decay channels are 2π vector channel and 3π axialvector channel

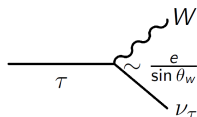
Components τ -decay



Vectorchannel



Axialvectorchannel



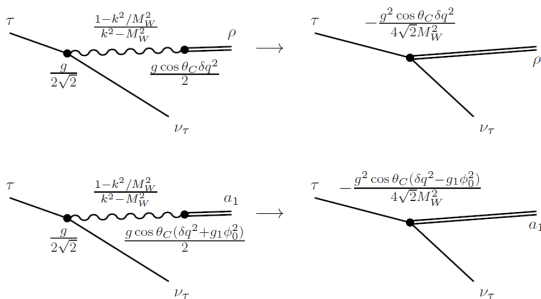
- $\mathcal{L}_{\tau \rightarrow W \nu_\tau} = \frac{g}{2\sqrt{2}} \bar{\nu}_\tau W_\mu^- \gamma^\mu (1 - \gamma^5) \tau + h.c.$
- $\Gamma_{\tau \rightarrow W \nu_\tau}(s) = \frac{1}{8\pi} \left(\frac{g}{2\sqrt{2}} \right)^2 \frac{m_\tau^3}{M_W^2} \left(1 - \frac{s}{m_\tau^2} \right)^2 \left(1 + \frac{2s}{m_\tau^2} \right)$

Weak Interaction Vertices and Effective Couplings

Tree-level weak interaction vertices with ρ and a_1 :

$$W\rho: \quad \frac{g \cos \theta_C}{2} \delta s W_\mu^- \rho^{\mu+} + \text{h.c.}$$

$$W a_1: \quad \frac{g \cos \theta_C}{2} (\delta s - g_1 \phi_0^2) W_\mu^- a_1^{\mu+} + \text{h.c.}$$



- decay rates

$$\Gamma_{\tau \rightarrow \rho \nu_\tau}(s) = \frac{g_\rho^2 m_\tau^3}{8\pi} \frac{1}{s} \left(1 - \frac{s}{m_\tau^2}\right)^2 \left(1 + \frac{2s}{m_\tau^2}\right),$$

$$\Gamma_{\tau \rightarrow a_1 \nu_\tau}(s) = \frac{g_{a_1}^2 m_\tau^3}{8\pi} \frac{1}{s} \left(1 - \frac{s}{m_\tau^2}\right)^2 \left(1 + \frac{2s}{m_\tau^2}\right)$$

- spectral densities (in a form suitable to compare to ALEPH data)

$$v_1(s) = \frac{(2\pi)^2}{S_{EW}} (\delta s)^2 \frac{1}{s} \frac{\rho_V(s)}{N}$$

$$a_1(s) = \frac{(2\pi)^2}{S_{EW}} (\delta s - g_1 \phi_0^2)^2 \frac{1}{s} \frac{\rho_A(s)}{N}$$

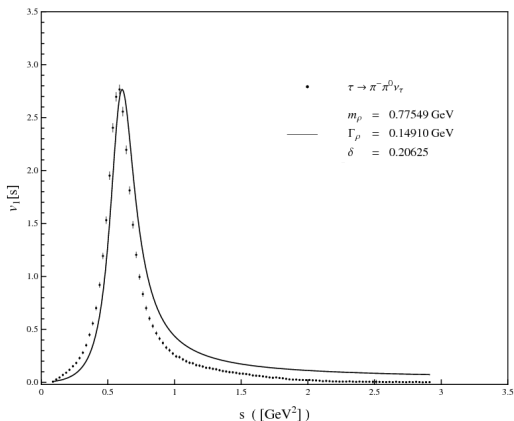
The s -dependent Decay Width $\Gamma_{\rho^- \rightarrow \pi^- \pi^0}(s)$

- $\Gamma_{\rho^- \rightarrow \pi^- \pi^0}(s) = \frac{m_\rho^5}{48\pi m_{a_1}^4} \left[1 - \left(\frac{2m_\pi}{\sqrt{s}} \right)^2 \right]^{\frac{3}{2}} \left[g_1 Z^2 + \frac{g_2}{2} (1 - Z^2) \right]^2$
- $v_1(s) = \frac{(2\pi)^2}{NS_{EW}} (\delta s)^2 \frac{1}{s} \frac{\sqrt{s} \Gamma_{\rho^- \rightarrow \pi^- \pi^0}(s)}{(s - m_\rho^2)^2 + (\sqrt{s} \Gamma_{\rho^- \rightarrow \pi^- \pi^0}(s))^2}$
- decay width calculated from $U(2)_L \times U(2)_R$ Linear Sigma Model

D. Parganlija, F. Giacosa, and D. H. Rischke, Vacuum Properties of Mesons in a Linear Sigma Model with Vector Mesons and Global Chiral Invariance, Phys. Rev. D82 (2010) 054024, [arXiv:1003.4934]

Parameters from hadronic sector

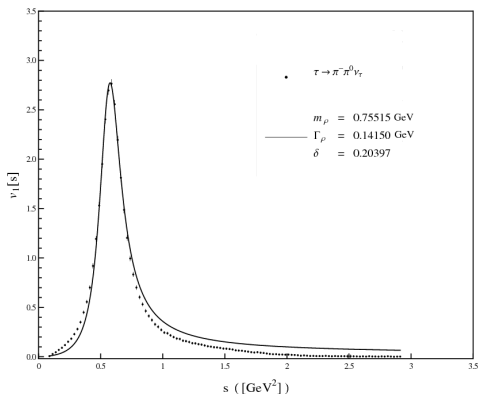
PDG: $m_\rho = 0.77549 \text{ GeV}$, $\Gamma_{\rho^- \rightarrow \pi^- \pi^0}(m_\rho^2) = 0.1491 \text{ GeV}$
one free parameter δ , fixed to obtain peak value



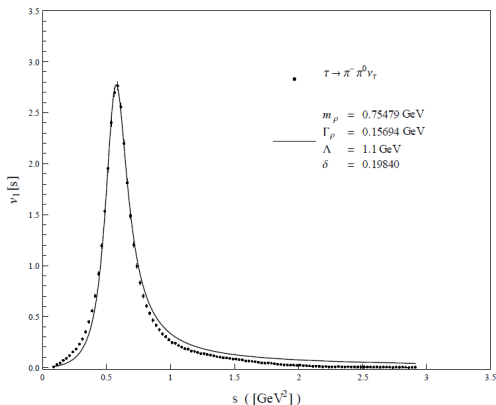
Fitted Parameters

improve agreement by refitting ρ mass and width:

$$\delta = 0.20397, \quad m_\rho = 0.75515 \text{ GeV}, \quad \Gamma_{\rho^- \rightarrow \pi^- \pi^0}(m_\rho) = 0.1415 \text{ GeV}$$



we can even further improve the result by removing the contributions from the 4 pion continuum by introducing a regularisation function $e^{-\frac{4m_\pi^2 - s}{\Lambda^2}}$ with cut-off parameter $\Lambda = 1.1$ GeV

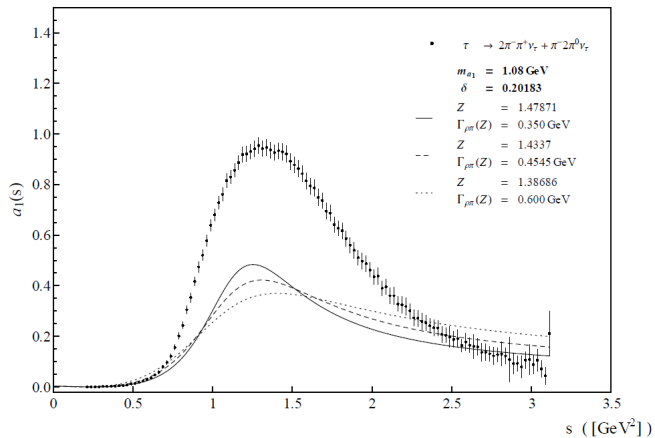


The s -dependent Decay Width $\Gamma_{a_1 \rightarrow \rho\pi}(s)$

- $$\Gamma_{a_1 \rightarrow \rho\pi}(s) = \frac{k(\sqrt{s}, m_\rho, m_\pi)}{12\pi m_{a_1}^2} \left[h_{\mu\nu}^2 - \frac{(h_{\mu\nu} K_1^\nu)^2}{m_\rho^2} - \frac{(h_{\mu\nu} P^\mu)^2}{m_{a_1}^2} + \frac{(h_{\mu\nu} P^\mu K_1^\nu)^2}{m_\rho^2 m_{a_1}^2} \right]$$
 also calculated by Parganlija, Giacosa and Rischke
- $$a_1(s) = \frac{(2\pi)^2}{NS_{EW}} (\delta s + g_1 \phi_0^2)^2 \frac{1}{s} \frac{\sqrt{s} \Gamma_{a_1 \rightarrow \rho\pi}(s)}{(s - m_{a_1}^2)^2 + (\sqrt{s} \Gamma_{a_1 \rightarrow \rho\pi}(s))^2}$$

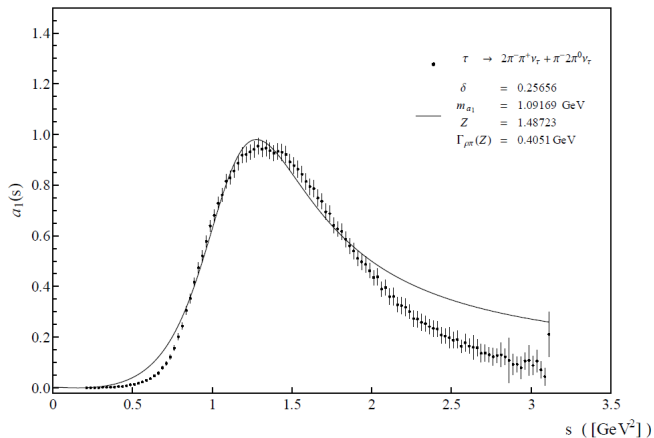
Axialvector Spectral Density

$$m_{a_1} = 1.08 \text{ GeV}, \quad \delta = 0.20183$$



Axialvector Spectral Density

Fitted axialvector spectral density



Conclusion and Outlook

- Vector Spectral Density can be described well with weak hadron-coupling $\delta \approx 0.2$ within the global $U(2)_L \times U(2)_R$ Linear Sigma Model
- Axialvector Spectral Density is still causing problems, however
 - δ may in fact be larger due to destructive interference effects from additional contributions in the vectorchannel, that will be considered
 - there are also other contributions in the axialvector channel that have not been considered yet

$$v_1(s) = \frac{m_\tau^2}{6 \cos \theta_C^2 S_{EW}} \frac{B(\tau^- \rightarrow V^- \nu_\tau)}{B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)} \times \frac{dN_V}{N_V ds} \left[\left(1 - \frac{s}{m_\tau^2}\right)^2 \left(1 + \frac{2s}{m_\tau^2}\right) \right]^{-1} \quad (1)$$

$$a_1(s) = \frac{m_\tau^2}{6 \cos \theta_C^2 S_{EW}} \frac{B(\tau^- \rightarrow A^- \nu_\tau)}{B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)} \times \frac{dN_A}{N_A ds} \left[\left(1 - \frac{s}{m_\tau^2}\right)^2 \left(1 + \frac{2s}{m_\tau^2}\right) \right]^{-1} \quad (2)$$

$$\frac{dN_V}{N_V ds} = \frac{\rho_V(s) \Gamma_{\tau^- \rightarrow V^- \nu_\tau}(s)}{\Gamma_{\tau^- \rightarrow V^- \nu_\tau}}, \quad \frac{dN_A}{N_A ds} = \frac{\rho_A(s) \Gamma_{\tau^- \rightarrow A^- \nu_\tau}(s)}{\Gamma_{\tau^- \rightarrow A^- \nu_\tau}} \quad (3)$$

where $\Gamma_{\tau^- \rightarrow V^- \nu_\tau}$ and $\Gamma_{\tau^- \rightarrow A^- \nu_\tau}$ are the partial widths in the corresponding channel and related to the branching fraction as

$$\Gamma_{\tau^- \rightarrow V^-(A^-) \nu_\tau} = B(\tau^- \rightarrow V^-(A^-) \nu_\tau) \cdot \Gamma_\tau^{\text{full}}$$

- First estimate for the parameter δ is taken from the convolution

$$\Gamma_{\tau \rightarrow 2\pi\nu_\tau} = \int_0^\infty ds \rho_{\rho \rightarrow 2\pi}(s) \Gamma_{\tau \rightarrow \rho\nu_\tau}(s) \quad (4)$$

- purely hypothetical assumption that ρ were stable

$$\begin{aligned} \Gamma_{\tau^- \rightarrow \pi^- \pi^0 \nu_\tau} &= \int_0^\infty ds \delta(s - m_\rho^2) \Gamma_{\tau^- \rightarrow \rho^- \nu_\tau}(s) \\ &= \frac{m_\tau^3}{8\pi m_\rho^2} \left(\frac{g^2 \cos \theta_C}{4\sqrt{2} M_W^2} \delta m_\rho \right)^2 \left(1 - \frac{m_\rho^2}{m_\tau^2} \right)^2 \left(1 + \frac{2m_\rho^2}{m_\tau^2} \right) \end{aligned}$$

- from experiment we know

$$\begin{aligned} \Gamma_{\tau^- \rightarrow \pi^- \pi^0 \nu_\tau}^{\text{exp.}}(m_\rho^2) &= 5.7811 \cdot 10^{-13} \text{ GeV} \\ \Rightarrow \delta &\approx 0.2 \end{aligned}$$