# Production of light nuclei in relativistic HIC via rate equations (arxiv:2108.13151) 

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## Introduction

- at LHC, the ALICE collaboration measured the yields of light nuclei (Jaroslav Adam et al., Phys. Rev. C, 93(2):024917, 2016)


(a) Hadron abundances and (b) Space-time diagram of a HIC statistical hadronization model (P. Braun-Munzinger et.al., Nucl. predictions (A. Andronic et.al., Phys. A, 987:144201, 2019) Nature 561, 321 (2018))


## Introduction

- the binding energies of light nuclei are much smaller then temperature of the environment
- the nucleosynthesis in heavy-ion collisions can be described by the Saha equation ( Volodymyr Vovchenko et. al., Phys. Lett. B, 800:135131, 2020)
- we use the principle of detailed balance to construct rate equations for the light nuclei
- the important reactions are of the following type

$$
\begin{aligned}
& \frac{d N_{A}}{d t}=\frac{\left\langle\sigma_{A+X \rightarrow a \cdot N+X} v_{r e l}\right\rangle}{V} N_{X}\left(-N_{A}+R \cdot N_{N}^{a}\right) \\
& R=\frac{N_{A}^{e q u} N_{X}^{\text {equ }}}{N_{N}^{\text {equ }}}
\end{aligned}
$$

## Introduction

- As an example, consider $\rho \leftrightarrow \pi+\pi$

$$
\begin{aligned}
& \frac{\mathrm{d} N_{\rho}}{\mathrm{d} t}=-\Gamma_{\rho \rightarrow 2 \pi} N_{\rho}+\frac{\left\langle\sigma_{\pi+\pi \rightarrow \rho} v_{\text {rel }}\right\rangle}{V} N_{\pi}^{2} \\
& \frac{\mathrm{~d} N_{\pi}}{\mathrm{d} t}=2 \Gamma_{\rho \rightarrow 2 \pi} N_{\rho}-2 \frac{\left\langle\sigma_{\pi+\pi \rightarrow \rho} v_{\text {rel }}\right\rangle}{V} N_{\pi}^{2}
\end{aligned}
$$

- in equilibrium, the Ihs is zero, thus we have

$$
\frac{\left\langle\sigma_{\pi+\pi \rightarrow \rho} v_{r e l}\right\rangle}{V}=\Gamma_{\rho \rightarrow 2 \pi} \frac{N_{\rho}^{\text {equ }}}{N_{\pi}^{e q u}}
$$

- by introducing fugacities $\lambda_{i}=e^{\frac{\mu_{i}(T)}{T}}=\frac{N_{i}(T)}{N_{i}^{e q u}(T)}$, we finaly get

$$
\begin{aligned}
\frac{\mathrm{d} \lambda_{\rho}}{\mathrm{d} t} & =-\Gamma_{\rho \rightarrow 2 \pi}\left(\lambda_{\rho}+\lambda_{\pi}^{2}\right) \\
\frac{\mathrm{d} \lambda_{\pi}}{\mathrm{d} t} & =2 \Gamma_{\rho \rightarrow 2 \pi} \frac{N_{\rho}^{\text {equ }}}{N_{\pi}^{\text {equ }}}\left(\lambda_{\rho}-\lambda_{\pi}^{2}\right)
\end{aligned}
$$

## Introduction

- we first have to determine the averaged cross sections, the volume and the multiplicities in chemical equilibrium in dependence of $T$
- particles: nucleons, the light nuclei and their corresponding anti-particles, $\pi, \rho, \omega, K, K^{*}, \Delta, \Lambda, \Sigma, \equiv$ and $\Omega$
- the catalysing particles $X$ are just $\pi$ and $K$, because they will have the largest contribution ( large abundances and cross sections)


## Thermal averaged cross sections

- average over Boltzmann distribution:

$$
\left\langle\sigma_{A+X \rightarrow a \cdot N+X} v_{r e l}\right\rangle=\frac{\iint \frac{d \vec{p}_{A}^{3}}{(2 \pi)^{3}} \frac{d \vec{p}_{X}^{3}}{(2 \pi)^{3}} e^{-\left(E_{A}+E_{X}\right) / T} \sigma\left(p_{\text {lab }}\right) v_{r e l}\left(\vec{p}_{A}, \vec{p}_{X}\right)}{\iint \frac{d \vec{p}_{A}^{3}}{(2 \pi)^{3}} \frac{d \vec{p}_{X}^{3}}{(2 \pi)^{3}} e^{-\left(E_{A}+E_{X}\right) / T}}
$$

- the known cross sections are taken from the PDG (Particle Data Group and P A et. al., Progress of Theoretical and Experimental Physics, 2020(8), 082020)
- we are interested in the case were the nuclei are splited in their nucleonic constituents $\rightarrow$ inelastic cross sections


## Thermal averaged cross sections

- as an example the results for $\pi^{+}+d$ scattering:


Figure: Total (blue) and inelastic (orange) thermal cross section for $\pi^{+}+d$ scattering as function of the tempertaure $T$.

## Thermal Model and Saha equation

- it is usefull to consider a simplified (analytical) example
- system is dominated by effectively massles pions
- relation between T and V (isentropic expansion): $V \propto T^{-3}$
- for all particles without the pions the non-relativistic approximation is used:

$$
N_{i}(T) \approx g_{i}\left(\frac{m_{i} T}{2 \pi}\right)^{\frac{3}{2}} e^{-m_{i} / T} \lambda_{i} V
$$

- here $\lambda_{i}$ are the fugacities for $\mu_{i}$
- a simplified expression for the $\mu_{i}$ 's by using $N_{i}\left(T_{c}\right)=N_{i}(T)$ and $\mu_{i}\left(T_{c}\right)=0$ ( Volodymyr Vovchenko et. al., Phys. Lett. B, 800:135131, 2020):

$$
\mu_{i}(T)=\frac{3}{2} T \ln \left(\frac{T}{T_{c}}\right)+m_{i}\left(1-\frac{T}{T_{c}}\right)
$$

## Thermal Model and Saha equation

- now we are able to calculate the normalised ratio $\frac{N_{A}(T)}{N_{A}\left(T_{c}\right)}$

$$
\begin{aligned}
\frac{N_{A}(T)}{N_{A}\left(T_{c}\right)} & =\left(\frac{T}{T_{c}}\right)^{\frac{3}{2}} e^{-m_{A}\left(\frac{1}{T}-\frac{1}{T_{c}}\right)} e^{\frac{\mu_{A}(T)}{T}} \frac{V(T)}{V_{c}} \\
& =\left(\frac{T}{T_{c}}\right)^{\frac{3}{2}} e^{-m_{A}\left(\frac{1}{T}-\frac{1}{T_{c}}\right)} e^{\mathrm{a} \cdot\left(\frac{3}{2} \ln \left(\frac{T}{T_{c}}\right)+m_{N}\left(\frac{1}{T}-\frac{1}{T_{c}}\right)\right)} \frac{T_{c}^{3}}{T^{3}} \\
& =\left(\frac{T}{T_{c}}\right)^{\frac{3}{2}(a-1)} e^{\left(a \cdot m_{N}-m_{A}\right)\left(\frac{1}{T}-\frac{1}{T_{c}}\right)} \\
& =\left(\frac{T}{T_{c}}\right)^{\frac{3}{2}(a-1)} e^{B_{A}\left(\frac{1}{T}-\frac{1}{T_{c}}\right)}
\end{aligned}
$$

- here we introduced the binding energy of a nucleus

$$
B_{A}=a \cdot m_{N}-m_{A}
$$

## Thermal Model and Saha equation

- this result is different to the standard thermal model result

$$
\left.\frac{N_{A}(T)}{N_{A}\left(T_{c}\right)}\right|_{\text {stand. }}=\left(\frac{T}{T_{c}}\right)^{\frac{3}{2}} e^{-m_{A}\left(\frac{1}{T}-\frac{1}{T_{c}}\right)}
$$

- the major difference is clearly the value in the exponential: $2 \mathrm{MeV} \approx B_{A} \ll m_{A} \approx 1000 \mathrm{MeV}$
- we see, that the exponential behaviour is strongly weakened


## Thermal Model and Saha equation

- to gain the full solution (HRG in PCE) we need to consider also the contributions of the other particles (Volodymyr Vovchenko et. al., Phys. Lett. B, 800:135131, 2020)

$$
\begin{aligned}
S_{\text {eff }}\left(T_{c}\right) & =V \sum_{j \in \text { all particles }} s_{j}\left(T, \tilde{\mu}_{j}, \mu_{B}, \mu_{S}\right) \\
N_{i} \text { eff }\left(T_{c}\right) & =V \sum_{j \in \text { all particles }}\left\langle n_{i}\right\rangle_{j} n_{j}\left(T, \tilde{\mu}_{j}, \mu_{B}, \mu_{S}\right) \\
B_{\text {eff }}\left(T_{c}\right) & =V \sum_{j \in \text { all particles }} B_{j} n_{j}\left(T, \tilde{\mu}_{j}, \mu_{B}, \mu_{S}\right) \\
0 & =V \sum_{j \in \text { all particles }} S_{j} n_{j}\left(T, \tilde{\mu}_{j}, \mu_{B}, \mu_{S}\right)
\end{aligned}
$$

## Thermal Model and Saha equation

- $\left\langle n_{i}\right\rangle_{j}$ is the averaged number of stable hadrons $i$ which came from the decay(-chain) of hadron $j$
- the chemical potentials are given as

$$
\tilde{\mu}_{j}=\sum_{i \in \text { stable }}\left\langle n_{i}\right\rangle_{j} \mu_{i} ; j \in \text { all particles }
$$

- by solving the set of non-linear equations we will get $V(T)$, $\mu_{i}(T), \mu_{B}(T)$ and $\mu_{S}(T)$
- all relativistic Boltzmann particles have their "normal" degeneracy factors with exception the $\Delta$-baryon $\left(g_{\Delta}^{\text {eff }}=2 g_{\Delta}\right)$
- fit model to experimental data (ALICE $0-10 \%$ central $\mathrm{Pb}-\mathrm{Pb}$ (2,72 TeV ) e.g. J. Adam et. a., Physics Letters B, 754:360372, 2016) to obtain: $V\left(T_{c}\right)=4017.5 \mathrm{fm}^{3}$, $\mu_{B}\left(T_{c}\right)=2.98 \mathrm{MeV}$ and $\mu_{S}\left(T_{c}\right)=0.39 \mathrm{MeV}$ at $T_{c}=155 \mathrm{MeV}$


## Thermal Model and Saha equation



Figure: The $\mu_{i}$ 's of the as stable considered hadrons in dependence of T .

## Thermal Model and Saha equation



Figure: The volume ratio in dependence of T .

## Solving the rate equations

$$
\begin{aligned}
& \begin{aligned}
\frac{\mathrm{d} N_{N}}{\mathrm{~d} t}= & 2 \bar{\alpha}_{D+\cdots \rightarrow 2 N+\cdots} N_{\pi}\left(N_{D}-R_{01} N_{N}^{2}\right)+3 \tilde{\alpha}_{r+\cdots \rightarrow 3 N+\pi} N_{\pi}\left(N_{T}-R_{02} N_{N}^{3}\right) \\
& +3 \bar{a}_{H+\cdots+\cdots \rightarrow N+\cdots} N_{\pi}\left(N_{\mathrm{He}^{3}}-R_{03} N_{N}^{3}\right)+4 \bar{\alpha}_{H++\cdots \rightarrow 4 N+\sigma} N_{\pi}\left(N_{\mathrm{He}^{t}}-R_{04} N_{N}^{4}\right)
\end{aligned} \\
& +\dot{\alpha}_{\Delta \rightarrow N+\cdots}\left(N_{\Delta}-R_{05} N_{N} N_{\pi}\right)+2 \tilde{\alpha}_{b+\kappa / \pi \rightarrow 2 N+k / K} N_{K / K}\left(N_{D}-R_{01} N_{N}^{2}\right) \\
& +3 \hat{\alpha}_{T+K / K \rightarrow 3 N+K / \pi^{N}} N_{K / K}\left(N_{T}-R_{02} N_{N}^{3}\right)+3 \hat{\alpha}_{H+{ }^{3}+K / \pi \rightarrow 3 N+K / \pi} N_{K / K}\left(N_{H_{e}}-R_{03} N_{N}^{3}\right) \\
& +4 \bar{\sigma}_{\mathrm{He}+\mathrm{H} / \kappa \rightarrow \mathrm{N}+\kappa / \kappa} N_{K / \kappa^{( }}\left(N_{\mathrm{He}^{4}}-R_{\mathrm{Ot}} N_{N}^{4}\right) \\
& \frac{\mathrm{d} N_{D}}{\mathrm{~d} t}=\tilde{\alpha}_{D+\cdots \rightarrow N+\square} N_{\pi}\left(-N_{D}+R_{01} N_{\mathrm{N}}^{2}\right)+\tilde{\alpha}_{D+K / \pi \rightarrow 2 N+N / \pi} N_{K / K}\left(-N_{D}+R_{01} N_{N}^{2}\right) \\
& \frac{\mathrm{d} N_{T}}{\mathrm{~d} t}=\bar{\alpha}_{T+\pi \rightarrow \mathrm{a} N+\pi} N_{\pi}\left(-N_{T}+R_{02} N_{N}^{3}\right)+\bar{\alpha}_{T+K / \pi \rightarrow \mathrm{N}+\kappa / \kappa} N_{K / K}\left(-N_{T}+R_{\mathrm{O} 2} N_{N}^{3}\right) \\
& \frac{\mathrm{d} N_{\mathrm{He}^{2}}}{\mathrm{~d} t}=\bar{\alpha}_{\mathrm{H}^{3}+\pi \rightarrow 3 \mathrm{~N}+=} N_{\mathrm{\pi}}\left(-N_{\mathrm{He}^{3}}+R_{03} N_{N}^{3}\right)+\bar{\alpha}_{\mathrm{nu}}{ }^{3}+K / \pi-3 N+K / K^{2} / K / K ~\left(-N_{\mathrm{He}^{2}}+R_{03} N_{N}^{3}\right) \\
& \frac{\mathrm{d} N_{\mathrm{He}^{4}}}{\mathrm{~d} t}=\tilde{\alpha}_{\mathrm{in}, 4+\pi \rightarrow 4 \mathrm{~N}+\pi} N_{\pi}\left(-N_{\mathrm{He}^{4}}+R_{\mathrm{0} 4} N_{N}^{4}\right)+\dot{\alpha}_{\mathrm{H} \mathrm{n}^{4}+\kappa / \pi \rightarrow 4 N+K / \pi} N_{K / K}\left(-N_{\mathrm{He}^{4}}+R_{\mathrm{O4}} N_{N}^{4}\right) \\
& \frac{\mathrm{d} N_{\bar{N}}}{\mathrm{~d} t}=2 \bar{\alpha}_{\overline{0}+\cdots \rightarrow 2 \pi+\cdots} N_{\pi}\left(N_{\bar{D}}-R_{\mathrm{O}} N_{N}^{2}\right)+3 \bar{\alpha}_{T+\cdots \rightarrow 2 \pi+\pi} N_{\pi}\left(N_{T}-R_{07} N_{N}^{3}\right)
\end{aligned}
$$

$$
\begin{aligned}
& +\dot{\alpha}_{J \rightarrow \pi+=}\left(N_{\bar{\Delta}}-R_{10} N_{\bar{N}} N_{\pi}\right)+2 \dot{\alpha}_{T+K / \pi \rightarrow / \pi+\kappa / \pi} N_{K / K}\left(N_{\bar{D}}-R_{06} N_{N}^{2}\right)
\end{aligned}
$$

$$
\begin{aligned}
& +4 \tilde{\alpha}_{\text {Hल+ }+K / K \rightarrow 4 \bar{N}+\kappa / K^{2}} N_{K / K}\left(N_{\overline{\mathrm{He}^{4}}}-R_{o g} N_{\frac{4}{N}}\right) \\
& \frac{\mathrm{d} N_{\bar{D}}}{\mathrm{~d} t}=\bar{\alpha}_{\bar{D}+\cdots \rightarrow 2 \bar{N}+\cdots} N_{\pi}\left(-N_{\bar{D}}+R_{06} N_{\bar{N}}^{2}\right)+\bar{\alpha}_{\bar{B}+\kappa / \pi \rightarrow 2 \bar{N}+N / \pi} N_{K / \pi}\left(-N_{\bar{D}}+R_{06} N_{\bar{N}}\right) \\
& \frac{\mathrm{d} N_{\bar{T}}}{\mathrm{~d} t}=\bar{\alpha}_{\bar{T}+n \rightarrow \mathrm{~N}+\cdots} N_{\pi}\left(-N_{T}+R_{07} N_{\bar{N}}^{3}\right)+\bar{\alpha}_{T+\kappa / \pi \rightarrow 3 \bar{N}+\kappa / \pi} N_{K / K}\left(-N_{T}+R_{07} N_{N}^{3}\right)
\end{aligned}
$$

$$
\begin{aligned}
& \frac{\mathrm{d} N_{\Delta}}{\mathrm{d} t}=\tilde{\alpha}_{\Delta \rightarrow N+\cdots}\left(-N_{\Delta}+R_{05} N_{N} N_{\pi}\right) \\
& \frac{\mathrm{d} N_{\bar{\Delta}}}{\mathrm{d} t}=\bar{\alpha}_{\overline{\mathrm{S}} \rightarrow++\cdots}\left(-N_{\bar{\Delta}}+R_{10} N_{\bar{N}} N_{\pi}\right) \\
& \frac{\mathrm{d} N_{\pi}}{\mathrm{d} t}=\bar{\alpha}_{\Delta \rightarrow N+\eta}\left(N_{\Delta}-R_{05} N_{N} N_{\pi}\right)+\bar{\alpha}_{\overline{3}+\pi+\cdots}\left(N_{\Delta}-R_{10} N_{\bar{N}} N_{\pi}\right) \\
& +2 \hat{\alpha}_{n \rightarrow 2 ⿱}\left(N_{\rho}-R_{11} N_{\pi}^{2}\right)+3 \tilde{\alpha}_{-\rightarrow 1+}\left(N_{\omega}-R_{12} N_{\pi}^{3}\right) \\
& \text { (3.19) } \\
& \frac{\mathrm{d} N_{\rho}}{\mathrm{d} t}=\bar{\alpha}_{p \rightarrow 3 ⿱}\left(-N_{p}+R_{11} N_{n}^{2}\right) \\
& \frac{\mathrm{d} N_{\omega}}{\mathrm{d} t}=\bar{\alpha}_{\omega} \ldots\left(-N_{\nu}+R_{12} N_{\pi}^{3}\right) \\
& \frac{\mathrm{d} N_{K}}{\mathrm{~d} t}=\bar{\alpha}_{\kappa_{* *} \rightarrow \kappa+\cdots}\left(N_{K *}-R_{13} N_{\pi} N_{K}\right) \\
& \frac{\mathrm{d} N_{\bar{R}}}{\mathrm{~d} t}=\bar{\alpha}_{\bar{K}^{*} \rightarrow \pi_{+-}}\left(N_{\bar{K}^{*}}-R_{14} N_{\pi} N_{\bar{K}}\right) \\
& \frac{\mathrm{d} N_{K^{*}}}{\mathrm{~d} t}=\tilde{\alpha}_{K^{*} \rightarrow K+\cdots}\left(-N_{K^{*}}+R_{13} N_{\pi} N_{K}\right) \\
& \frac{\mathrm{d} N_{K^{*}}}{\mathrm{~d} t}=\bar{\alpha}_{\kappa^{*} \rightarrow \pi_{+-}}\left(-N_{\bar{K}^{*}}+R_{14} N_{\pi} N_{K^{\prime}}\right)
\end{aligned}
$$

## Solving the rate equations

- for all light nuclei up to $\mathrm{He}^{4}$ rate equations has been implemented, but also the decays of $\rho, \omega, K^{*}$ and $\Delta$ has been considered
- we have just related the volume and temperature, but the system contains ODE's in time
- here we consider a parametrisation $V(t)$ (Yinghua Pan and Scott Pratt, Phys. Rev. C, 89(4):044911, 2014):

$$
\frac{V(t)}{V_{c h}}=\frac{t}{t_{c h}} \frac{t_{\perp}^{2}+t^{2}}{t_{\perp}^{2}+t_{c h}^{2}} ; t_{\perp}=6.5 \frac{\mathrm{fm}}{c} ; \quad t_{c h}=9 \frac{\mathrm{fm}}{c}
$$

## Solving the rate equations

- first we want to look at the rates:


Figure: The rates $\alpha=\frac{\left\langle v_{\text {rel }} \sigma_{\text {tot }}\right\rangle}{V} N_{X}^{\text {eq }}$ in a fixed volume $V=4000 \mathrm{fm}^{3}$ for different temperatures (but fixed during the equilibration of the system) for different deuteron break up reactions

## Solving the rate equations

- now we solve the full set of equations and look at the evolution of the deuteron number:


Figure: Normalised particle number of deuterons to the value at $T_{c}=155 \mathrm{MeV}$ for $g_{\Delta}^{\text {eff }}=2 g_{\Delta}$

## Solving the rate equations

- now we want to check how fast the system equilibrates when starting out of equilibrium:


Figure: The ratio of deuterons to protons normalized to the same ratio at equilibrium for different initial conditions with $g_{\Delta}^{\text {eff }}=2 g_{\Delta}$.

## Solving the rate equations



Figure: Solid lines represent the results of the rate equations, while dashed curves show the result of the HRG in PCE. The colored bands represent the experimental data (ALICE)

## Effect of the $N+\bar{N} \rightleftharpoons 5 \pi$ reaction

- a big advantage of the rate equation approach is the possibility of the annihilation of stable hadrons e.g. nucleons

$$
\begin{aligned}
\frac{d N_{N}}{d t} & =\frac{\left\langle\sigma_{N+\bar{N}=5 \pi} v_{r e l}\right\rangle}{V}\left(-N_{N} N_{\bar{N}}+R_{15} N_{\pi}^{5}\right) \\
\frac{d N_{\bar{N}}}{d t} & =\frac{\left\langle\sigma_{N+\bar{N}=5 \pi} v_{r e l}\right\rangle}{V}\left(-N_{N} N_{\bar{N}}+R_{15} N_{\pi}^{5}\right) \\
\frac{d N_{\pi}}{d t} & =5 \frac{\left\langle\sigma_{N+\bar{N}=5 \pi} v_{r e l}\right\rangle}{V}\left(N_{N} N_{\bar{N}}-R_{15} N_{\pi}^{5}\right)
\end{aligned}
$$

$$
R_{15}=\frac{N_{N}^{e q} N_{\bar{N}}^{e q}}{N_{\pi}^{e q 5}}
$$

## Effect of the $N+\bar{N} \rightleftharpoons 5 \pi$ reaction

- the averaged cross section is about 50 mb for $p+\bar{p}$ scattering
- this type of reaction exlicitly violates the conservation of stable hadrons, but the net baryon number is still conserved


Figure: Normalised particle number of deuterons to the value at $T_{c}=155 \mathrm{MeV}$ and $g_{\Delta}^{\text {eff }}=2 g_{\Delta}$.

## Conclusions and Outlook

- both approaches are in great agreement with each other and also in the error range of the experimental data
- the same procedure could be done for RHIC or SPS energies
- the annihilation of nucleon and anti-nucleon into five pions only leads to a $4-5 \%$ decrease in the effective nucleon number
- under-occupation in the nucleons leads to a suppression of the light nuclei


## Conclusions and Outlook

- calculations also support earlier assumptions, that the nuclei do not need to be formed at the chemical freeze-out
- this approach neglects the formation time of the nuclei
- a quantum mechanical description of creation and decay of bound states (the nuclei) in an open thermal system (fireball) is needed

