

Lecture

Models for heavy-ion collisions: (Part 3): transport models

SS2024: ‚Dynamical models for relativistic heavy-ion collisions‘

BUU (VUU) equation

Reminder:
Lecture 4

Boltzmann (Vlasov)-Uehling-Uhlenbeck equation (NON-relativistic formulation!)

- free propagation of particles in the self-generated HF mean-field potential with an **on-shell collision term**:

$$\frac{d}{dt} f(\vec{r}, \vec{p}, t) \equiv \frac{\partial}{\partial t} f(\vec{r}, \vec{p}, t) + \frac{\vec{p}}{m} \vec{\nabla}_{\vec{r}} f(\vec{r}, \vec{p}, t) - \vec{\nabla}_{\vec{r}} U(\vec{r}, t) \vec{\nabla}_{\vec{p}} f(\vec{r}, \vec{p}, t) = \left(\frac{\partial f}{\partial t} \right)_{coll}$$

Collision term for 1+2→3+4 (let's consider fermions) :

$$I_{coll} \equiv \left(\frac{\partial f}{\partial t} \right)_{coll} \Rightarrow \frac{1}{((2\pi)^3)^3} \int d^3 p_2 d^3 p_3 d^3 p_4 \cdot w(1+2 \rightarrow 3+4) \cdot P$$

$$\times (2\pi)^3 \delta^3(\vec{p}_1 + \vec{p}_2 - \vec{p}_3 - \vec{p}_4) (2\pi) \delta\left(\frac{\vec{p}_1}{2m_1} + \frac{\vec{p}_2}{2m_2} - \frac{\vec{p}_3}{2m_3} - \frac{\vec{p}_4}{2m_4}\right)$$

Probability including
Pauli blocking of fermions

Transition probability for 1+2→3+4: $w(1+2 \rightarrow 3+4) \Rightarrow v_{12} \cdot \frac{d^3 \sigma}{d^3 q}$

where $v_{12} = \frac{\hbar}{m} / |\vec{p}_1 - \vec{p}_2|$ - relative velocity of the colliding nucleons

$\frac{d^3 \sigma}{d^3 q}$ - differential cross section, q - momentum transfer $\vec{q} = \vec{p}_1 - \vec{p}_3$

2. Density-matrix formalism: Correlation dynamics

Density-matrix formalism

□ **Schrödinger equation** for a **system of N fermions**:

$$i\hbar \frac{\partial}{\partial t} \Psi_N(1, \dots, N; t) = H_N(1, \dots, N) \Psi_N(1, \dots, N; t) \quad (1)$$

antisymmetric N-body wave function

or in Dirac notation:
$$i\hbar \frac{\partial}{\partial t} |\Psi_N(t)\rangle = H_N |\Psi_N(t)\rangle$$

Notation:
i – particle index of
 many body system
 (*i=1, ..., N*):

$$i \equiv \mathbf{r}_i, \sigma_i, \tau_i$$

i = coordinate, spin,
 isospin,...

□ **Hamiltonian operator**:

$$H_N = \sum_{i=1}^N h^0(i) + \sum_{i<j}^{N-1} v(ij),$$

2-body interaction

One-body Hamiltonian:

$$h^0(i) = t(i) + U^0(i)$$

**kinetic energy operator + (possible) external mean-field potential
 (e.g. external electromagnetic field)**

Hermitian Hamiltonian: $H_N = H_N^\dagger$

Consider (1) - hermitean conjugate:

$$-i\hbar \frac{\partial}{\partial t'} \Psi_N^*(1', \dots, N'; t) = H_N(1', \dots, N') \Psi_N^*(1', \dots, N'; t') \quad (2)$$

(1)*Ψ_N* - Ψ_N*(2) :
$$i\hbar \left(\frac{\partial}{\partial t} + \frac{\partial}{\partial t'} \right) \Psi_N(1, \dots, N; t) \Psi_N^*(1', \dots, N'; t') =$$

$$(H_N(1, \dots, N) - H(1', \dots, N'; t')) \Psi_N(1, \dots, N; t) \Psi_N^*(1', \dots, N'; t') \quad (2.1)$$

Density-matrix formalism

- Introduce two time density :

$$\rho_N(1, \dots, N, 1', \dots, N'; t, t') = \Psi_N(1, \dots, N; t) \Psi_N^*(1', \dots, N'; t') \quad (2.2)$$

or in Dirac notation:

$$\rho_N(1, \dots, N, 1', \dots, N'; t, t') = \langle 1', \dots, N' | \rho_N(t, t') | 1, \dots, N \rangle$$

Substitute (2.2) in (2.1):

$$i\hbar \left(\frac{\partial}{\partial t} + \frac{\partial}{\partial t'} \right) \rho_N = (H_N(1, \dots, N) - H(1', \dots, N'; t')) \rho_N$$

$$i\hbar \left(\frac{\partial}{\partial t} + \frac{\partial}{\partial t'} \right) \rho_N = [H_N, \rho_N]$$

- Restrict to $t'=t$:

$$\rho_N(1, \dots, N, 1', \dots, N'; t) = \rho_N(1, \dots, N, 1', \dots, N'; t, t) \delta(t - t')$$

- Schrödinger eq. in density-operator representation →

von Neumann (or Liouville) eq. (in matrix representation) describes an N-particle system in- or out-of equilibrium :

$$i\hbar \frac{\partial}{\partial t} \rho_N(1, \dots, N; 1' \dots N'; t) = [H_N, \rho_N(1, \dots, N; 1' \dots N'; t)] \quad (2.3)$$

Density-matrix formalism

□ Introduce a **reduced density matrices** $\rho_n(1\dots n, 1'\dots n'; t)$ by taking the trace (integrate) over particles $n+1, \dots, N$ of ρ_N :

$$\rho_n = \frac{1}{(N-n)!} \text{Tr}_{n+1, \dots, N} \rho_N = \frac{1}{n+1} \text{Tr}_{n+1} \{\rho_{n+1}\} \quad (3) \quad \text{Recurrence}$$

Here the relative **normalization** between ρ_n and ρ_{n+1} is fixed and it is useful to choose the normalization

$$\text{Tr}_{1, \dots, N} \rho_N = N!$$

which leads to the following **normalization for the one-body density matrix**:

$$\text{Tr}_{(1=1')} \rho(11'; t) = \sum_i \langle a_i^\dagger a_i \rangle = N \quad \text{Tr} \rightarrow \int \frac{d^3p}{(2\pi\hbar)^3} \int d^3r$$

i.e. the particle number for the **N-body Fermi system**.

The **normalization of the two-body density matrix** then reads as

$$\begin{aligned} \text{Tr}_{(1,2)} \rho_2 &= \sum_{i,j} \langle a_i^\dagger a_j^\dagger a_j a_i \rangle = - \sum_{i,j} \langle a_i^\dagger a_j^\dagger a_i a_j \rangle = \sum_{i,j} \{ \langle a_i^\dagger a_i a_j^\dagger a_j \rangle - \langle a_i^\dagger a_j \rangle \delta_{ij} \} \\ &= (N-1) \sum_j \langle a_j^\dagger a_j \rangle = N(N-1) \end{aligned}$$

The traces of the density matrices ρ_n (for $n < N$): $\text{Tr}_{(1, \dots, n)} \rho_n = \frac{N!}{(N-n)!}$

Density matrix formalism: BBGKY-Hierarchy

Taking corresponding traces (i.e. $\text{Tr}_{(n+1, \dots, N)}$) of the von-Neumann equation we obtain the **BBGKY-Hierarchy** (Bogolyubov, Born, Green, Kirkwood and Yvon)

$$i \frac{\partial}{\partial t} \rho_n = \left[\sum_{i=1}^n h^0(i), \rho_n \right] + \left[\sum_{1=i \langle j}^{n-1} v(ij), \rho_n \right] + \sum_{i=1}^n \text{Tr}_{n+1} [v(i, n+1), \rho_{n+1}] \quad (4)$$

for $1 \leq n \leq N$ with $\rho_{N+1} = 0$.

- This set of equations is **equivalent to the von-Neumann equation**
- The **approximations or truncations** of this set will reduce the information about the system

□ The explicit equations for $n=1, n=2$ read:

$$i \frac{\partial}{\partial t} \rho_1 = [h^0(1), \rho_1] + \text{Tr}_2 [v(12), \rho_2], \quad (5)$$

$$i \frac{\partial}{\partial t} \rho_2 = \left[\sum_{i=1}^2 h^0(i), \rho_2 \right] + [v(12), \rho_2] + \text{Tr}_3 [v(13) + v(23), \rho_3] \quad (6)$$

Eqs. (5,6) are **not closed** since eq. (6) for ρ_2 requires information from ρ_3 . Its equation reads:

$$i \frac{\partial}{\partial t} \rho_3 = \left[\sum_{i=1}^3 h^0(i), \rho_3 \right] + [v(12) + v(13) + v(23), \rho_3] + \text{Tr}_4 [v(14) + v(24) + v(34), \rho_4] \quad (7)$$

Correlation dynamics

➔ Introduce the **cluster expansion** ➔ Correlation dynamics:

□ **1-body density matrix:** $\rho_1(11') = \rho(11')$,

1 – initial state of particle „1“
1' – final state of the same particle „1“

□ **2-body density matrix (consider fermions):**

$$(8) \quad \rho_2(12, 1'2') = \rho(11')\rho(22') - \rho(12')\rho(21') + c_2(12, 1'2') = \rho_{20}(12, 1'2') + c_2(12, 1'2')$$

2PI= 2-particle-irreducible approach

$$(9) \quad \rho_2(12, 1'2') = \mathcal{A}_{12}\rho(11')\rho(22') + c_2(12, 1'2')$$

2-body antisymmetrization operator:

$$\mathcal{A}_{ij} = 1 - P_{ij}$$

Permutation operator

1PI = 1-particle-irreducible approach + **2-body correlations**
(TDHF approximation)

By neglecting c_2 in (9) we get the **limit** of independent particles (**Time-Dependent Hartree-Fock**). This implies that all effects from **collisions or correlations are incorporated in c_2** and higher orders in c_2 etc.

$$\rho_3(123, 1'2'3') = \rho(11')\rho(22')\rho(33') - \rho(12')\rho(21')\rho(33')$$

□ **3-body density matrix:**

$$\begin{aligned} & -\rho(13')\rho(22')\rho(31') - \rho(11')\rho(32')\rho(23') + \rho(13')\rho(21')\rho(32') + \rho(12')\rho(31')\rho(23') \\ & + \rho(11')c_2(23, 2'3') - \rho(12')c_2(23, 1'3') - \rho(13')c_2(23, 2'1') + \rho(22')c_2(13, 1'3') \\ & - \rho(21')c_2(13, 2'3') - \rho(23')c_2(13, 1'2') + \rho(33')c_2(12, 1'2') - \rho(31')c_2(12, 3'2') \\ & - \rho(32')c_2(12, 1'3') + c_3(123, 1'2'3'). \end{aligned} \quad (10)$$

3-body correlations

Correlation dynamics

The goal: from BBGKY-hierarchy obtain closed equation **for 1-body density matrix** within 2PI **discarding explicit three-body correlations c_3**

□ for that we reformulate eq. (5) for ρ_1 using the cluster expansion (correlation dynamics):

$$(5) \quad i\hbar \frac{\partial}{\partial t} \rho_1 = [h^0(1), \rho_1] + \text{Tr}_2[v(12), \rho_2]$$

substitutute eq. (8) for ρ_2 in eq. (5)

$$(8) \quad \underline{\rho_2(12, 1'2')} = \rho(11')\rho(22') - \rho(12')\rho(21') + c_2(12, 1'2') = \rho_{20}(12, 1'2') + c_2(12, 1'2')$$

→ we obtain **EoM for the one-body density matrix:**

$$i \frac{\partial}{\partial t} \underline{\rho(11'; t)} = [h^0(1) - h^0(1')] \rho(11'; t) + \text{Tr}_{(2=2')} [v(12) \mathcal{A}_{12} - v(1'2') \mathcal{A}_{1'2'}] \rho(11'; t) \rho(22'; t) + \text{Tr}_{(2=2')} [v(12) - v(1'2')] \underline{c_2(12, 1'2'; t)}$$

2-body correlations

* How to obtain the 2-body correlation matrix c_2 ?

Correlation dynamics

□ In order to obtain the 2-body correlation matrix c_2 , we start with eq. (6) for ρ_2

$$(6) \quad i\hbar \frac{\partial}{\partial t} \rho_2 = \left[\sum_{i=1}^2 h^0(i), \rho_2 \right] + [v(12), \rho_2] + \text{Tr}_3 [v(13) + v(23), \rho_3]$$

substitute eq. (10) for ρ_3
and discarding explicit 3-body correlations c_3

$$\begin{aligned} \rho_3(123, 1'2'3') = & \rho(11')\rho(22')\rho(33') - \rho(12')\rho(21')\rho(33') \quad (10) \\ & - \rho(13')\rho(22')\rho(31') - \rho(11')\rho(32')\rho(23') + \rho(13')\rho(21')\rho(32') + \rho(12')\rho(31')\rho(23') \\ & + \rho(11')c_2(23, 2'3') - \rho(12')c_2(23, 1'3') - \rho(13')c_2(23, 2'1') + \rho(22')c_2(13, 1'3') \\ & - \rho(21')c_2(13, 2'3') - \rho(23')c_2(13, 1'2') + \rho(33')c_2(12, 1'2') - \rho(31')c_2(12, 3'2') \\ & - \rho(32')c_2(12, 1'3') + c_3(123, 1'2'3') \end{aligned}$$

→ we obtain EoM for the two-body correlation matrix c_2 :

$$\begin{aligned} i \frac{\partial}{\partial t} c_2(12, 1'2'; t) = & [h^0(1) + h^0(2) - h^0(1') - h^0(2')] c_2(12, 1'2'; t) \quad (12) \\ & + \text{Tr}_{(3=3')} [v(13)\mathcal{A}_{13} + v(23)\mathcal{A}_{23} - v(1'3')\mathcal{A}_{1'3'} - v(2'3')\mathcal{A}_{2'3'}] \rho(33'; t) c_2(12, 1'2'; t) \\ & + [v(12) - v(1'2')] \rho_{20}(12, 1'2') \quad \leftarrow \rho(11')\rho(22') - \rho(12')\rho(21') \\ & - \text{Tr}_{(3=3')} \{ v(13)\rho(23'; t)\rho_{20}(13, 1'2'; t) - v(1'3')\rho(32'; t)\rho_{20}(12, 1'3'; t) \\ & + v(23)\rho(13'; t)\rho_{20}(32, 1'2'; t) - v(2'3')\rho(31'; t)\rho_{20}(12, 3'2'; t) \} \\ & = \rho_{20}(12, 1'2') \\ & + [v(12) - v(1'2')] c_2(12, 1'2'; t) \\ & - \text{Tr}_{(3=3')} \{ v(13)\rho(23'; t)c_2(13, 1'2'; t) - v(1'3')\rho(32'; t)c_2(12, 1'3'; t) \\ & + v(23)\rho(13'; t)c_2(32, 1'2'; t) - v(2'3')\rho(31'; t)c_2(12, 3'2'; t) \} \\ & + \text{Tr}_{(3=3')} \{ [v(13)\mathcal{A}_{13}\mathcal{A}_{1'2'} - v(1'3')\mathcal{A}_{1'3'}\mathcal{A}_{12}] \rho(11'; t)c_2(32, 3'2'; t) \\ & + [v(23)\mathcal{A}_{23}\mathcal{A}_{1'2'} - v(2'3')\mathcal{A}_{2'3'}\mathcal{A}_{12}] \rho(22'; t)c_2(13, 1'3'; t) \}. \end{aligned}$$

Correlation dynamics

To reduce the complexity we introduce **useful notations**:

□ a **one-body Hamiltonian** by

$$h(i) = t(i) + U^s(i) = t(i) + \text{Tr}_{(n=n')} v(in) \mathcal{A}_{in} \rho(nn'; t), \quad (13)$$

$$h(i') = t(i') + U^s(i') = t(i') + \text{Tr}_{(n=n')} v(i'n') \mathcal{A}_{i'n'} \rho(nn'; t)$$

kinetic term + interaction with the **self-generated time-dependent mean field**

□ **Pauli-blocking operator** is uniquely defined by (14)

$$Q_{ij}^- = 1 - \text{Tr}_{(n=n')} (P_{in} + P_{jn}) \rho(nn'; t); \quad Q_{i'j'}^- = 1 - \text{Tr}_{(n=n')} (P_{i'n'} + P_{j'n'}) \rho(nn'; t),$$

□ **Effective 2-body interaction in the medium**:

$$V^-(ij) = Q_{ij}^- v(ij); \quad V^-(i'j') = Q_{i'j'}^- v(i'j'), \quad (15)$$

Resummed interaction → **G-matrix approach**

Correlation dynamics

□* EoM for the **one-body density matrix**:

(16)

$$i \frac{\partial}{\partial t} \rho(11'; t) = [h(1) - h(1')] \rho(11'; t) + \text{Tr}_{(2=2')} [v(12) - v(1'2')] c_2(12, 1'2'; t)$$

TDHF

2-body correlations

EoM (16) describes the propagation of a particle in the **self-generated mean field U^s** with additional 2-body correlations that are further specified in the EoM (17) for c_2 :

□* EoM for the **2-body correlation matrix**:

$$(17) \quad i \frac{\partial}{\partial t} \underline{c_2(12, 1'2'; t)} = \left[\sum_{i=1}^2 h(i) - \sum_{i'=1'}^{2'} h(i') \right] \underline{c_2(12, 1'2'; t)} + [V^=(12) - V^=(1'2')] \rho_{20}(12, 1'2'; t) + [V^=(12) - V^=(1'2')] \underline{c_2(12, 1'2'; t)} + \text{Tr}_{(3=3')} \{ [v(13) \mathcal{A}_{13} \mathcal{A}_{1'2'} - v(1'3') \mathcal{A}_{1'3'} \mathcal{A}_{12}] \rho(11'; t) \underline{c_2(32, 3'2'; t)} + [v(23) \mathcal{A}_{23} \mathcal{A}_{1'2'} - v(2'3') \mathcal{A}_{2'3'} \mathcal{A}_{12}] \rho(22'; t) \underline{c_2(13, 1'3'; t)} \}.$$

Propagation of two particles 1 and 2 in the **mean field U^s**
 Born term: bare 2-body scattering
 resummed in-medium interaction with intermediate Pauli blocking (**G-matrix theory**)
 2-Particle-2-hole interactions (important for groundstate correlations) and damping of low energy modes

Note: Time evolution of c_2 depends on the distribution of a **third particle**, which is integrated out in the trace! The third particle is interacting as well!


*: EoM is obtained after the 'cluster' expansion and neglecting the explicit 3-body correlations c_3

Vlasov equation

BBGKY-Hierarchie - 1PI →
eq.(11) with $c_2(1,2,1',2')=0$

$$i\hbar \frac{\partial}{\partial t} \rho(11'; t) = [h^0(1) - h^0(1')] \rho(11'; t)$$

TDHF



$$\frac{\partial}{\partial t} \rho(\vec{r}, \vec{r}', t) + \frac{i}{\hbar} \left[\frac{\hbar^2}{2m} \vec{\nabla}_r^2 + U(\vec{r}, t) - \frac{\hbar^2}{2m} \vec{\nabla}_{r'}^2 - U(\vec{r}', t) \right] \rho(\vec{r}, \vec{r}', t) = 0$$

➤ perform **Wigner transformation** of one-body density distribution function $\rho(r,r',t)$ →

$$f(\vec{r}, \vec{p}, t) = \int d^3s \exp\left(-\frac{i}{\hbar} \vec{p} \vec{s}\right) \rho\left(\vec{r} + \frac{\vec{s}}{2}, \vec{r} - \frac{\vec{s}}{2}, t\right) \quad (18)$$

$f(r,p,t)$ is the **single particle phase-space distribution function**

After the **1st order gradient expansion** → **Vlasov equation of motion**

- free propagation of particles in the self-generated HF mean-field potential $U(r,t)$:

$$\frac{\partial}{\partial t} f(\vec{r}, \vec{p}, t) + \frac{\vec{p}}{m} \vec{\nabla}_{\vec{r}} f(\vec{r}, \vec{p}, t) - \vec{\nabla}_{\vec{r}} U(\vec{r}, t) \vec{\nabla}_{\vec{p}} f(\vec{r}, \vec{p}, t) = 0 \quad (19)$$

$$U(\vec{r}, t) = \frac{1}{(2\pi\hbar)^3} \sum_{\beta_{occ}} \int d^3r' d^3p V(\vec{r} - \vec{r}', t) f(\vec{r}', \vec{p}, t)$$

Uehling-Uhlenbeck equation: collision term

$$i \frac{\partial}{\partial t} \rho(11'; t) = [h(1) - h(1')] \rho(11'; t) + \underbrace{\text{Tr}_{(2=2')} [v(12) - v(1'2')] c_2(12, 1'2'; t)}_{\text{2-body correlations}} \quad (21)$$

TDHF – Vlasov equation

2-body correlations

Collision term:

$$I(11', t) = \text{Tr}_{(2=2')} [v(12) - v(1'2')] c_2(12, 1'2'; t) \quad (22)$$

□ perform Wigner transformation

□ Formally solve the EoM for c_2 (with some approximations in momentum space):

$$\begin{aligned} i \frac{\partial}{\partial t} \underline{c_2(12, 1'2'; t)} &= \left[\sum_{i=1}^2 h(i) - \sum_{i'=1'}^{2'} h(i') \right] c_2(12, 1'2'; t) \\ &+ [V^=(12) - V^=(1'2')] \rho_{20}(12, 1'2'; t) \\ &+ [V^=(12) - V^=(1'2')] c_2(12, 1'2'; t) \\ &+ \text{Tr}_{(3=3')} \{ [v(13) \mathcal{A}_{13} \mathcal{A}_{1'2'} - v(1'3') \mathcal{A}_{1'3'} \mathcal{A}_{12}] \rho(11'; t) c_2(32, 3'2'; t) \\ &+ [v(23) \mathcal{A}_{23} \mathcal{A}_{1'2'} - v(2'3') \mathcal{A}_{2'3'} \mathcal{A}_{12}] \rho(22'; t) c_2(13, 1'3'; t) \}. \end{aligned} \quad (23)$$

□ and insert obtained c_2 in the expression (22) for $I(11', t)$: → BUU EoM

Boltzmann (Vlasov)-Uehling-Uhlenbeck (B(V)UU) equation : Collision term

$$\frac{d}{dt} f(\vec{r}, \vec{p}, t) \equiv \frac{\partial}{\partial t} f(\vec{r}, \vec{p}, t) + \frac{\vec{p}}{m} \vec{\nabla}_{\vec{r}} f(\vec{r}, \vec{p}, t) - \vec{\nabla}_{\vec{r}} U(\vec{r}, t) \vec{\nabla}_{\vec{p}} f(\vec{r}, \vec{p}, t) = \left(\frac{\partial f}{\partial t} \right)_{coll} \quad (24)$$

Collision term for $1+2 \rightarrow 3+4$ (let's consider fermions) :

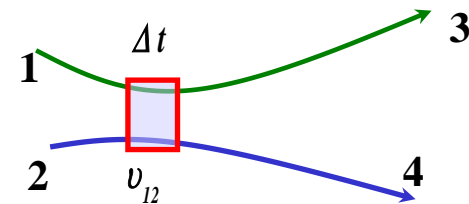
$$I_{coll} = \frac{4}{(2\pi)^3} \int d^3 p_2 d^3 p_3 \int d\Omega |v_{12}| \delta^3(\vec{p}_1 + \vec{p}_2 - \vec{p}_3 - \vec{p}_4) \cdot \frac{d\sigma}{d\Omega} (1+2 \rightarrow 3+4) \cdot P \quad (25)$$

Probability including Pauli blocking of fermions:

$$P = \underbrace{f_3 f_4 (1 - f_1) (1 - f_2)}_{\text{Gain term}} - \underbrace{f_1 f_2 (1 - f_3) (1 - f_4)}_{\text{Loss term}} \quad (26)$$

Gain term
 $3+4 \rightarrow 1+2$

Loss term
 $1+2 \rightarrow 3+4$



For particle 1 and 2:

Collision term = **Gain term** - **Loss term**

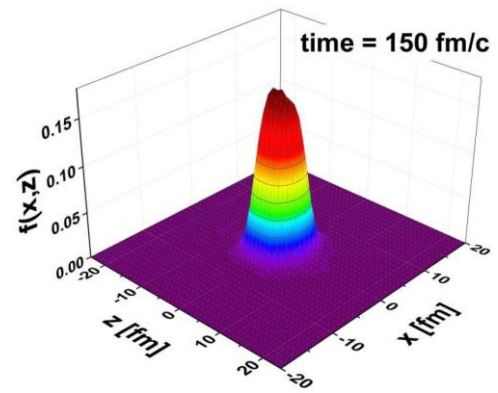
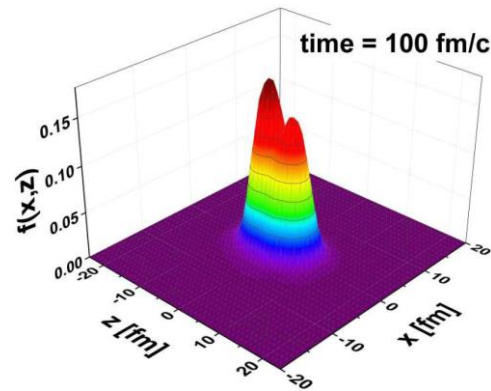
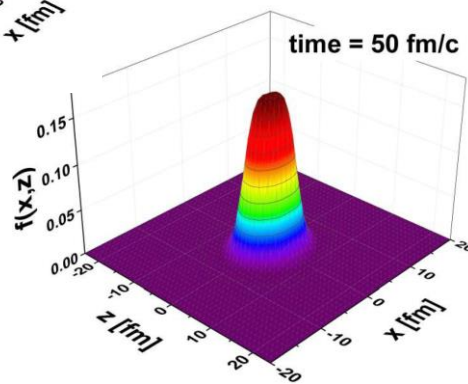
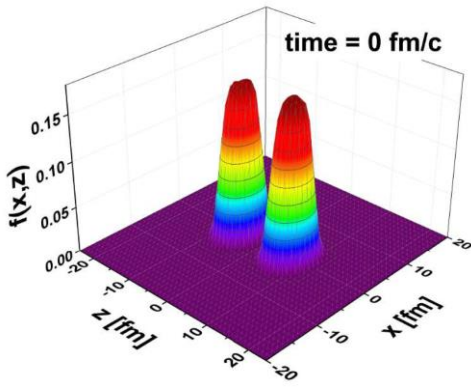
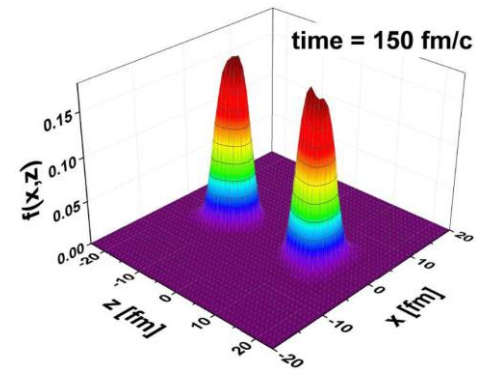
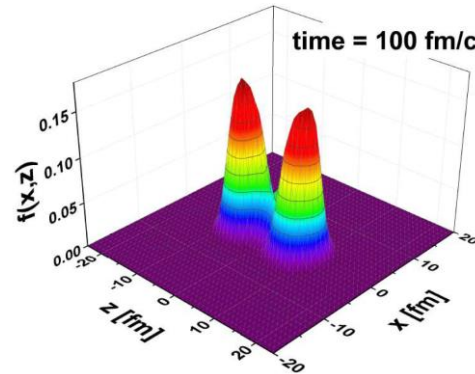
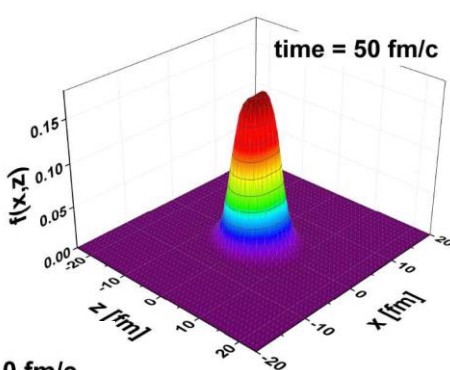
$$I_{coll} = G - L$$

The **BUU equations** (24) describes the propagation in the **self-generated mean-field** $U(\vec{r}, t)$ as well as mutual **two-body interactions** respecting the **Pauli-principle**

Density distribution: Vlasov vs. Boltzmann equation

Ca+Ca, 40 A MeV

Vlasov



Boltzmann

Covariant transport equation

From non-relativistic to a relativistic formulation of on-shell transport equations:

Non-relativistic Schrödinger equation

→ relativistic Dirac equation

Non-relativistic dispersion relation:

$$E = \frac{\vec{p}^2}{2m} + U(\vec{r})$$

$U(r)$ – density dependent potential
(with attractive and repulsive parts)

! Not Lorentz invariant, i.e.
dependent on the frame

Relativistic dispersion relation:

$$E^{*2} = m^{*2} + \vec{p}^{*2}$$

$$m^* = m + U_S \quad \leftarrow \quad U_S - \text{scalar potential (attractive)}$$

$$\vec{p}^* = \vec{p} + \vec{U}_V \quad \leftarrow \quad U_\mu = (U_0, \vec{U}_V)$$

$$E^* = E - U_0 \quad \leftarrow \quad \text{vector 4-potential (repulsive)}$$

$\mu = 0, 1, 2, 3$

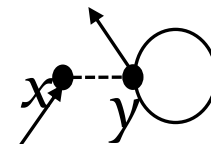
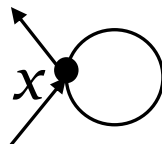
! Lorentz invariant, i.e.
independent on the frame

→ Consider the **Dirac equation** with **local** and **non-local** mean fields:

$$(i\gamma^\mu \partial_\mu - m)\psi(x) - U^{MF}(x)\psi(x) - \int d^4y U^{MD}(x,y)\psi(y) = 0 \quad (12)$$

here

$$x \equiv (t, \vec{r}) \quad y \equiv (t, \vec{r}')$$



Covariant transport equation

Perform a **Wigner transformation** of eq. (12) → in phase-space:

$$U_S^{MD}(x, p) = C_S \frac{4}{(2\pi)^3} \int d^4 p' D_S^{MD}(p - p') m^*(x, p') f(x, p')$$

$$U_{V\mu}^{MD}(x, p) = C_V \frac{4}{(2\pi)^3} \int d^4 p' D_V^{MD}(p - p') \Pi_\mu(x, p') f(x, p')$$

where $f(x, p)$ is the **single particle phase-space distribution function**

$$m^*(x, p) = m + U_S^{MF}(x) + U_S^{MD}(x, p) \quad - \text{effective mass}$$

$$\Pi_\mu(x, p) = p_\mu - U_{V\mu}^{MF}(x) - U_{V\mu}^{MD}(x, p) \quad - \text{effective momentum}$$

where

$$D_{S,V}^{MD}(p - p') \equiv \int d^4 z D_{S,V}^{MD}(z) e^{iz(p - p')}$$

Fourier transform

,Ansatz' for D -function:

$$D_{S(V)}^{MD}(p - p') = \frac{\Lambda_{S(V)}^2}{\Lambda_{S(V)}^2 - (p - p')^2}$$

Covariant transport equation



□ Covariant relativistic on-shell BUU equation :

from many-body theory by connected Green functions in phase-space + mean-field limit for the propagation part (VUU)

$$\left\{ \left(\Pi_\mu - \Pi_\nu (\partial_\mu^p U_\nu^\nu) - m^* (\partial_\mu^p U_S^\nu) \right) \partial_x^\mu + \left(\Pi_\nu (\partial_\mu^x U_\nu^\nu) + m^* (\partial_\mu^x U_S^\nu) \right) \partial_p^\mu \right\} f(x, p) = I_{coll}$$

$$I_{coll} \equiv \sum_{2,3,4} \int d2 d3 d4 [G^+ G]_{1+2 \rightarrow 3+4} \delta^4(\Pi + \Pi_2 - \Pi_3 - \Pi_4)$$

$$d2 \equiv \frac{d^3 p_2}{E_2}$$

$$\times \{ f(x, p_3) f(x, p_4) (1 - f(x, p)) (1 - f(x, p_2))$$

Gain term
3+4 → 1+2

$$- f(x, p) f(x, p_2) (1 - f(x, p_3)) (1 - f(x, p_4)) \}$$

Loss term
1+2 → 3+4

where $\partial_\mu^x \equiv (\partial_t, \vec{\nabla}_r)$

$$m^*(x, p) = m + U_S(x, p) \quad - \text{effective mass}$$

$$\Pi_\mu(x, p) = p_\mu - U_\mu(x, p) \quad - \text{effective momentum}$$

$U_S(x, p)$, $U_\mu(x, p)$ are scalar and vector part of particle self-energies

$\delta(\Pi_\mu \Pi^\mu - m^{*2})$ – mass-shell constraint

Nuclear equation of state (EoS)

EoS:

$$\frac{E_B}{A} = \frac{\varepsilon}{\rho_N} - M_0$$

□ **Energy density:** $\varepsilon = \langle : T^{00} : \rangle,$

$$P = \langle : T^{ii} : \rangle = \frac{1}{3} \sum_{i=1}^3 \langle : T^{ii} : \rangle$$

□ **Nucleon density:**

$$\rho_N(x) = 2 \langle : \bar{\Psi} \gamma^0 \Psi : \rangle = d \int \frac{d^3 \Pi}{(2\pi)^3} \left(f_p^*(x, \Pi) - f_a^*(x, \Pi) \right)$$

$f^*(\mathbf{x}, \Pi)$ – Fermi distribution with effective masses and energies

Momentum dependence of the Schrödinger-equivalent potential:

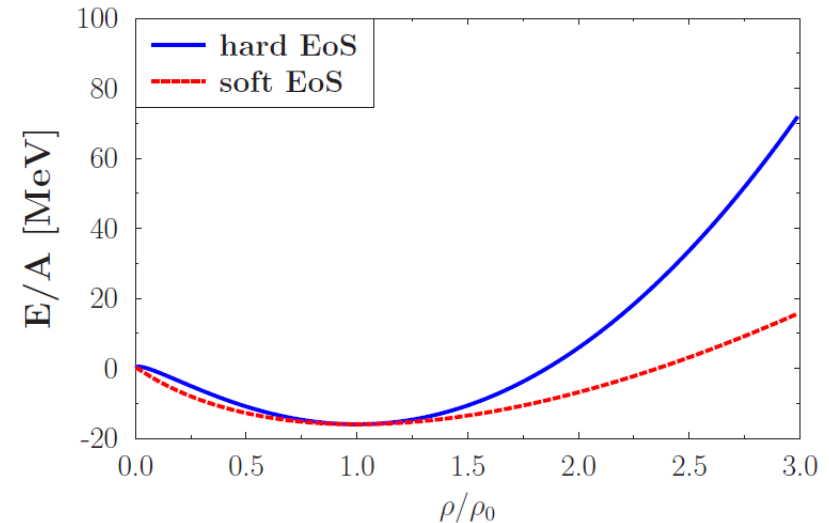
$$U_{\text{SEP}} = U_s(\rho_0, P) + U_0(\rho_0, P) + \frac{1}{2M_N} (U_s(\rho_0, P)^2 - U_0(\rho_0, P)^2) + U_0(\rho_0, P) \frac{\sqrt{P^2 + M_N^2} - M_N}{M_N}$$

Compression modulus:

$$K = -V \frac{dP}{dV} = 9\rho^2 \left. \frac{\partial^2 [E/A(\rho)]}{(\partial \rho)^2} \right|_{\rho=\rho_0}$$

Soft EoS: K=200 MeV
Hard EoS: K=380 MeV

EoS for infinite nuclear matter



Useful literature

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