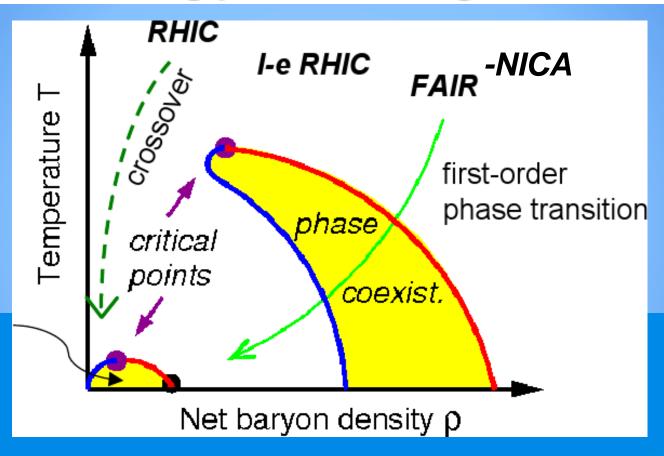
Lecture 5: Some examples of experimental data analysis

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Frankfurt Institute for Advanced Studies

Expected phase diagram of strongly-interacting matter



Such a phase diagram is still a beautiful dream! We hope that future FAIR-NICA experiments will help to establish what is the reality.

Effects of fast dynamics

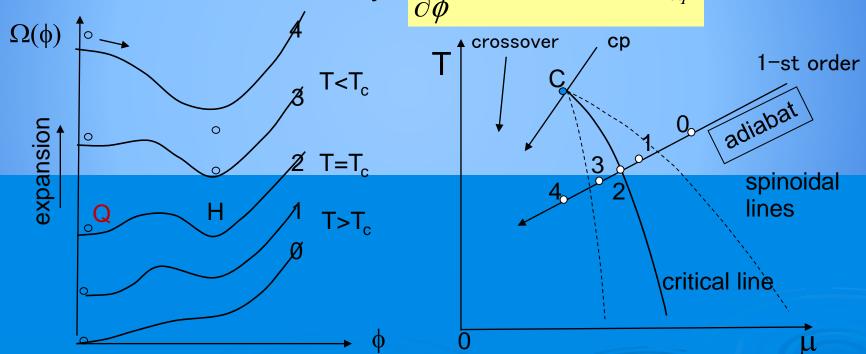
Effective thermodynamic potential for a 1st order transition

$$\Omega(\phi; T, \mu) = \Omega_0(T, \mu) + \frac{a}{2}\phi^2 + \frac{b}{4}\phi^4 + \frac{c}{6}\phi^6$$

a,b,c are functions of T and μ

Equilibrium ϕ is determined by

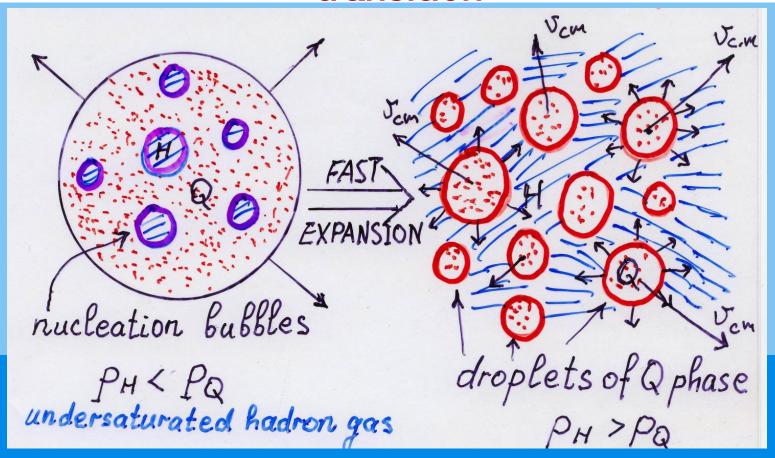
$$\frac{\partial \Omega}{\partial \phi} = 0 \Rightarrow P = -\Omega(\phi_{eq})$$



In rapidly expanding system 1-st order transition is delayed until the barrier between two competing phases disappears - spinodal decomposition

I. Mishustin, Phys. Rev. Lett. 82 (1999) 4779; Nucl. Phys. A681 (2001) 56

Rapid expansion through a 1st order phase transition

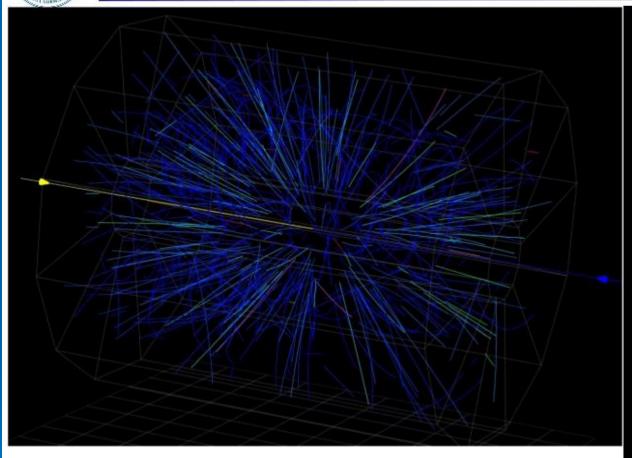


The system is trapped in a metastable state until it enters the spinodal instability region, when Q phase becomes unstable and splits into droplets

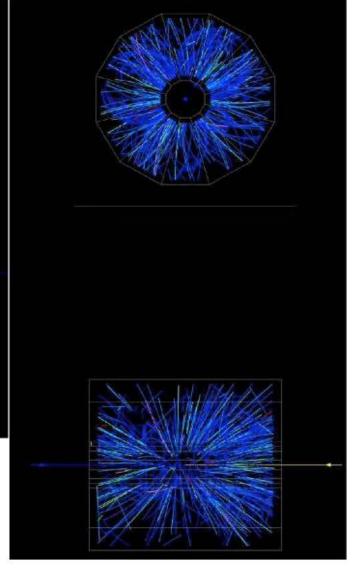
Csernai&Mishustin, 1995; Mishustin, 1999; Rafelski et al. 2000; Randrup, 2003; Peach&Stoecker, 2003; Stephanov, 2005, 2009; Steinheimer&Randrup 2013; Nahrgang, Herold, Mishustin, Bleicher, 2013-p.t.; Liang, Li, Song, 2016-p.t.. ...



3D Event Display at STAR



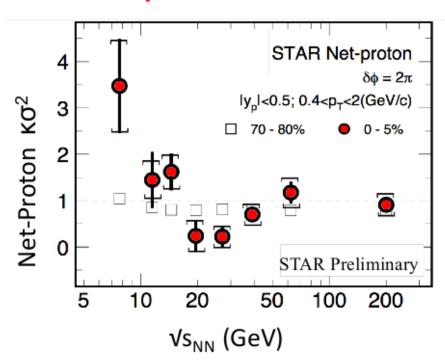
BES-II, Au+Au collisions at 19.6 GeV.





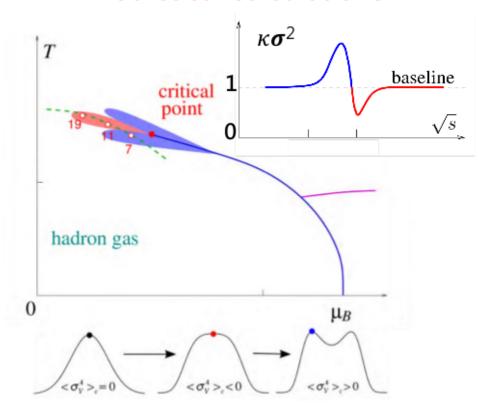
Net-Proton Fluctuations

Experimental Measure



STAR: Phys. Rev. Lett. 105, 022302 (2010). Phys. Rev. Lett. 112, 032302 (2014). PoS CPOD2014 (2015) 019.

Theoretical calculations



M. Stephanov, PRL107, 052301(2011) J. Phys. G: 38, 124147 (2011).

➤ First observation of the non-monotonic energy dependence of fourth order net-proton fluctuations. Hint of entering Critical Region?

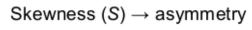


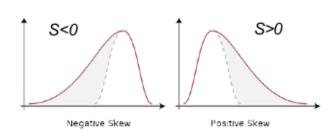
Higher Moments of Conserved Quantities (B, Q, S)

1. Higher order cumulants/moments: describe the shape of distributions and quantify fluctuations. (sensitive to the correlation length (ξ))

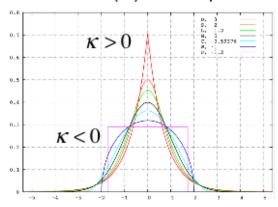
$$<\delta N> = N - < N>$$
 $C_1 = M = < N>$
 $C_2 = \sigma^2 = <(\delta N)^2>$
 $C_3 = S\sigma^3 = <(\delta N)^3>$
 $C_4 = \kappa\sigma^4 = <(\delta N)^4> -3<(\delta N)^2>^2$

$$\langle (\delta N)^3 \rangle_c \approx \xi^{4.5}, \quad \langle (\delta N)^4 \rangle_c \approx \xi^7$$





Kurtosis (K) \rightarrow Sharpness



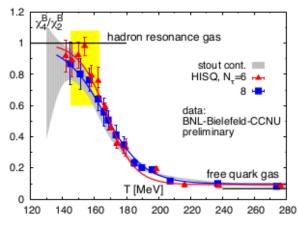
M. A. Stephanov, Phys. Rev. Lett. 102, 032301 (2009); 107, 052301 (2011). M.Asakawa, S. Ejiri and M. Kitazawa, Phys. Rev. Lett. 103, 262301 (2009).

2. Direct connect to the susceptibility of the system.

$$\frac{\chi_q^4}{\chi_q^2} = \kappa \sigma^2 = \frac{C_{4,q}}{C_{2,q}} \qquad \frac{\chi_q^3}{\chi_q^2} = S\sigma = \frac{C_{3,q}}{C_{2,q}},$$

$$\chi_q^{(n)} = \frac{1}{VT^3} \times C_{n,q} = \frac{\partial^n (p/T \wedge 4)}{\partial (\mu_q)^n}, q = B, Q, S$$

S. Ejiri et al, Phys. Lett. B 633 (2006) 275. Cheng et al, PRD (2009) 074505. B. Friman et al., EPJC 71 (2011) 1694. F. Karsch and K. Redlich, PLB 695, 136 (2011).S. Gupta, et al., Science, 332, 1525(2012). A. Bazavov et al., PRL109, 192302(12) // S. Borsanyi et al., PRL111, 062005(13)



Conclusions

- Phase transitions in relativistic heavy-ion collisions will most likely proceed out of equilibrium
- > 2nd order phase transition (with CEP) is too weak to produce significant observable effects in fast dynamics
- Non-equilibrium effects in a1st order p.t. (spinodal decomposition, dynamical domain formation) may help to identify the chiral/deconfinement phase transition
- If QGP domains (droplets) survive until the freeze-out stage, they will show up by large non-statistical fluctuations of hadron multiplicities in phase space (in single events)
- Exotic objects like strangelets have a better chance to be formed in such a non-equilibrium scenario