

Investigating Proton Number Fluctuations with HADES at SIS18

EMMI Fluctuation Workshop - Nov 2015
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- Baryon number fluncts. in the few-GeV regime
- The HADES experiment at SIS18
- Experimental artifacts

Net proton nb. fluctuations: $\delta(\Delta N_p)$

Net number of protons: $\Delta N_p = N_p - N_{\bar{p}}$

- $\delta(\Delta N_p)$ used as estimate of net baryon nb. fluctuations
- justified for $\sqrt{s} \geq 10 \text{ GeV}$ Kitazawa & Asakawa, PRC 86, 024904 (2012)
- baryon nb. is conserved quantity \rightarrow fluctuations in y- p_t bin
- likewise for net charge & net strangeness fluctuations

- At **SIS18**:
 - fixed target expt. ($\sqrt{s} \leq 2.7 \text{ GeV}$)
 - what about fragments (d, t, He, etc.) ?
 - what about neutrons ?

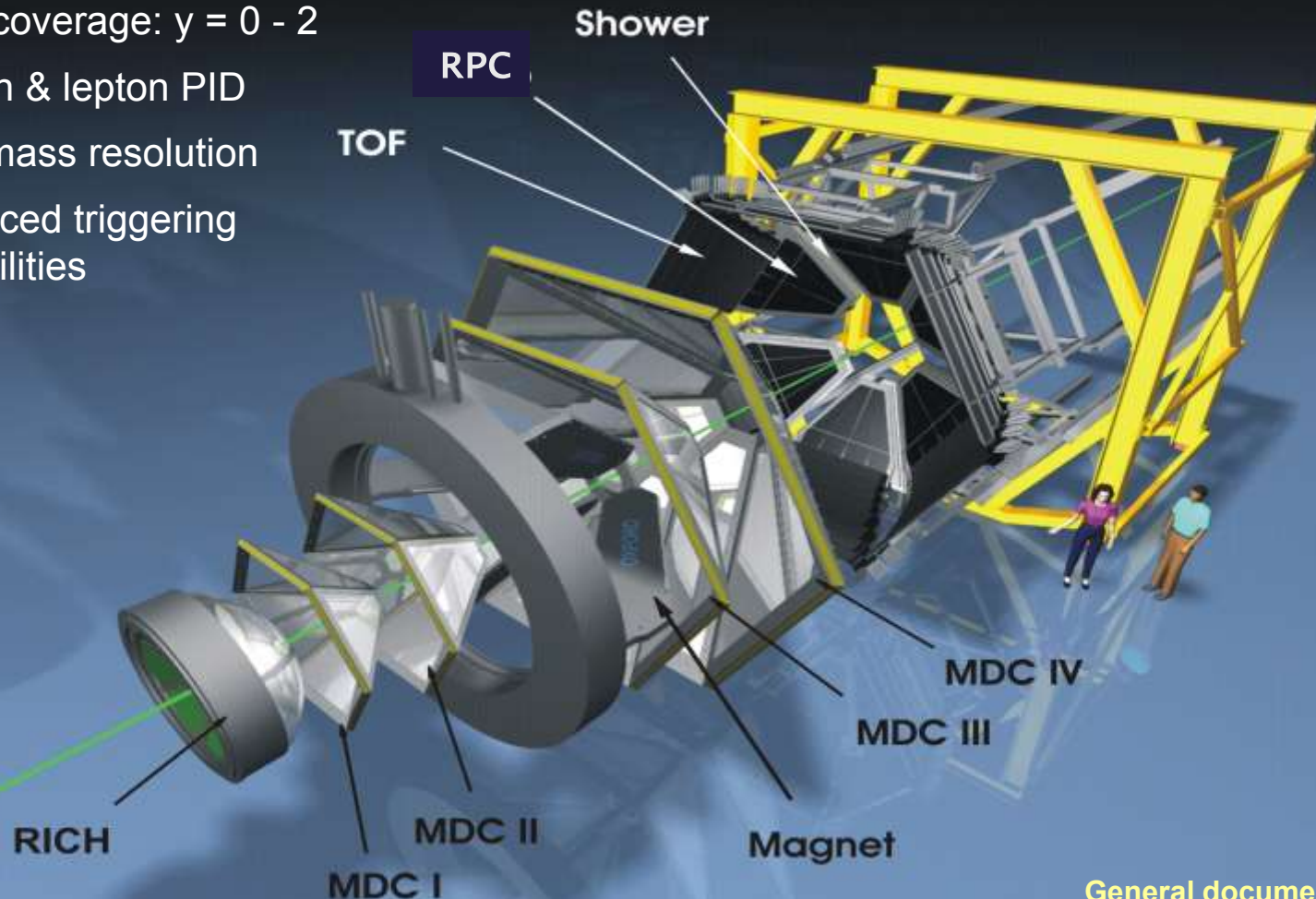
\rightarrow Here focus is on proton nb. fluctuations $\delta(N_p)$

The HADES experiment at GSI



High Acceptance DiElectron Spectrometer

- azimuth. symmetry
- large coverage: $y = 0 - 2$
- hadron & lepton PID
- $< 2\%$ mass resolution
- advanced triggering capabilities



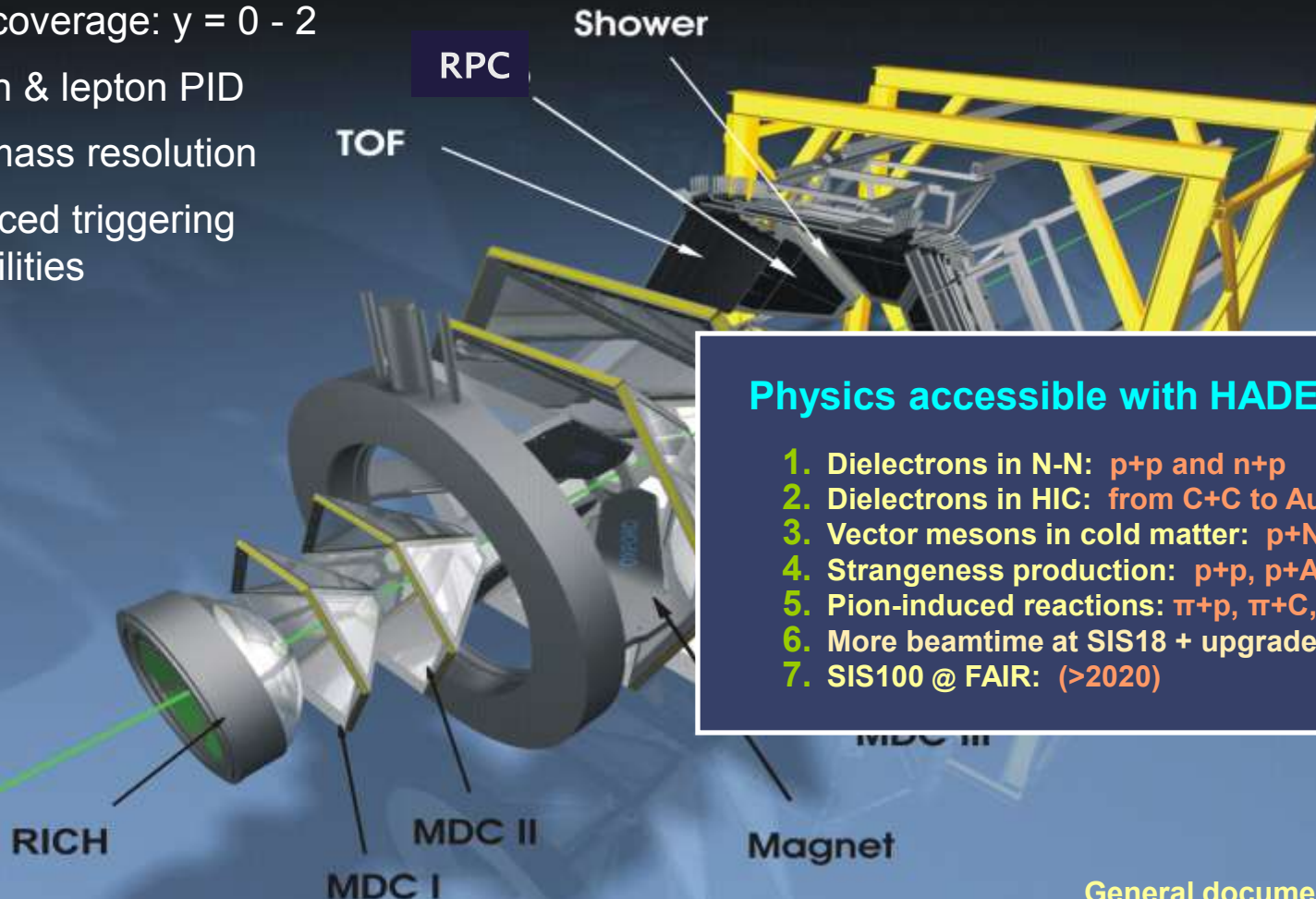
General documentation at:
<http://www-hades.gsi.de>

The HADES experiment at GSI



High Acceptance DiElectron Spectrometer

- azimuth. symmetry
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- hadron & lepton PID
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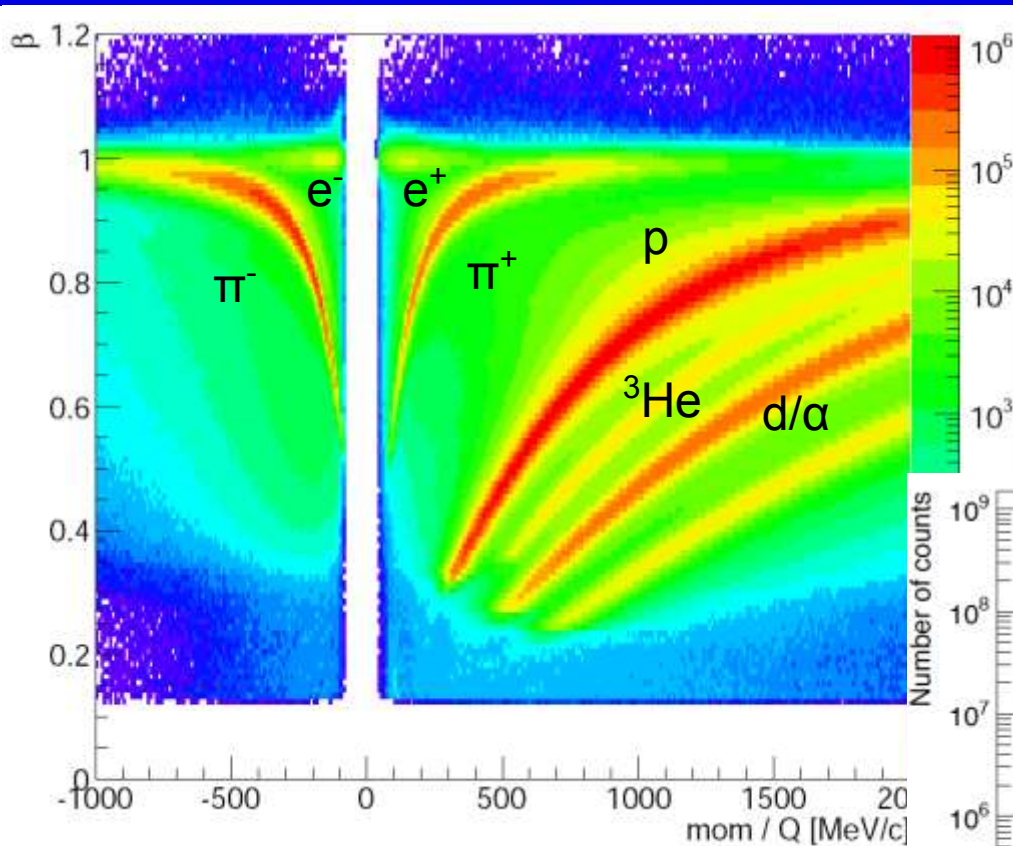


Physics accessible with HADES:

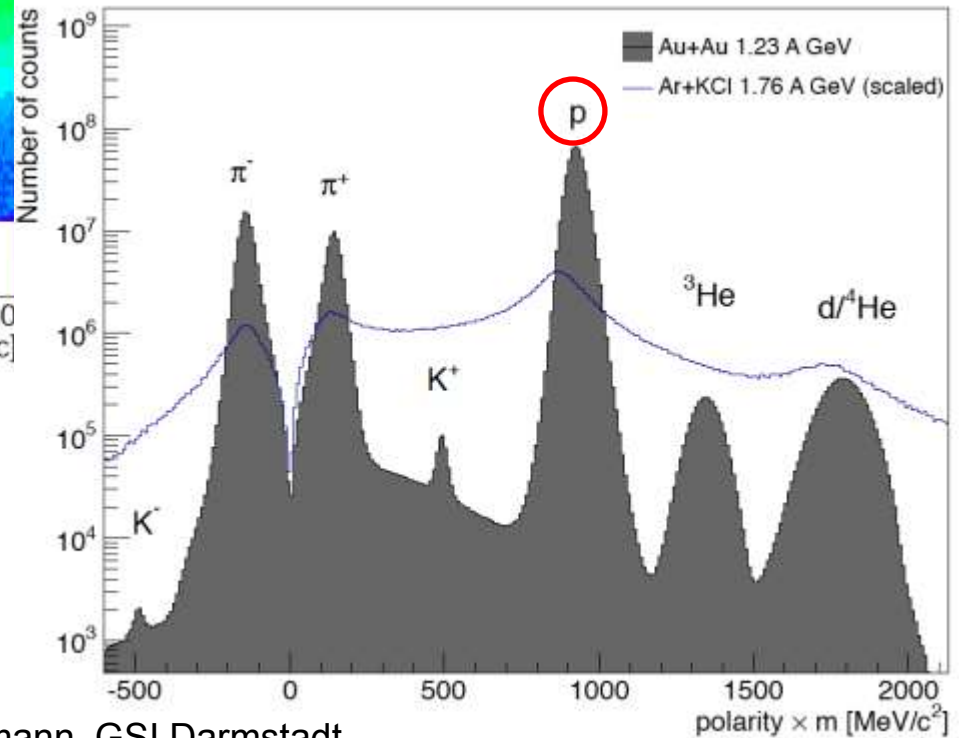
1. Dielectrons in N-N: $p+p$ and $n+p$
2. Dielectrons in HIC: from $C+C$ to $Au+Au$
3. Vector mesons in cold matter: $p+Nb$
4. Strangeness production: $p+p$, $p+A$, $A+Agg$
5. Pion-induced reactions: $\pi+p$, $\pi+C$, $\pi+W$
6. More beamtime at SIS18 + upgrade
7. SIS100 @ FAIR: (>2020)

General documentation at:
<http://www-hades.gsi.de>

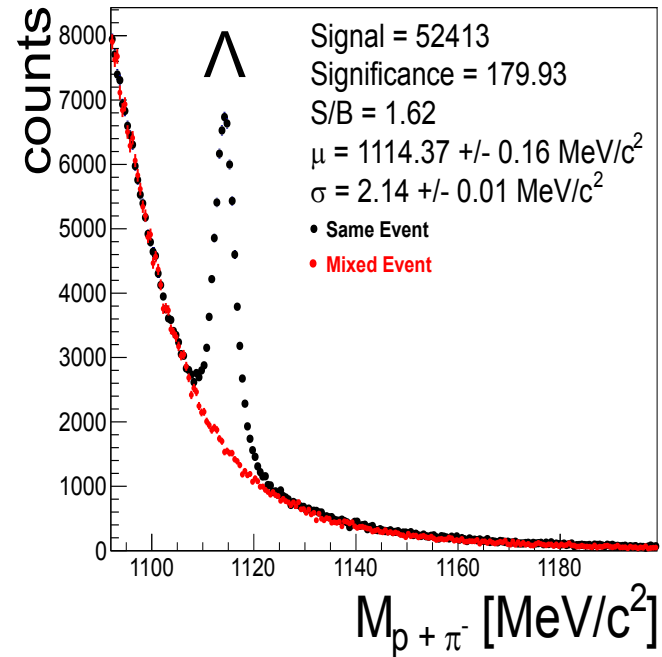
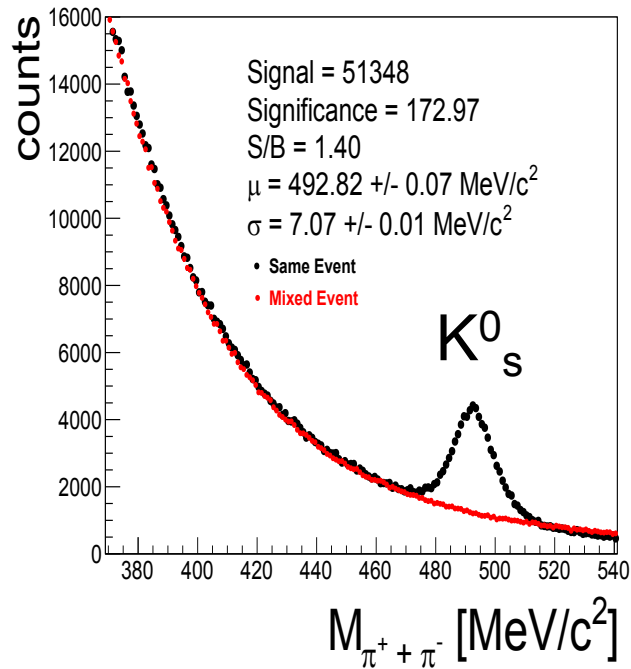
Performance: particle ID



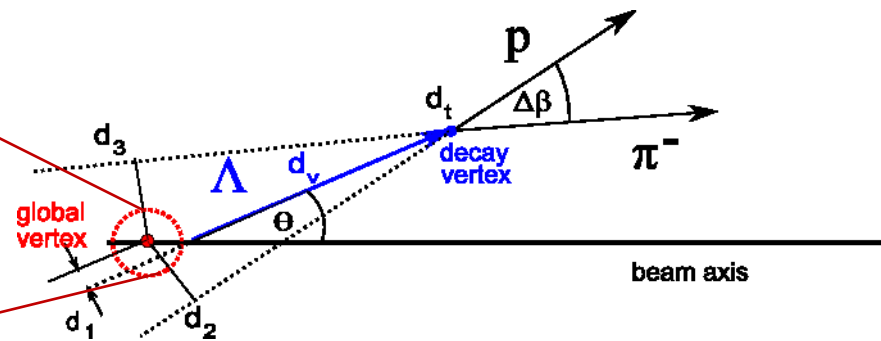
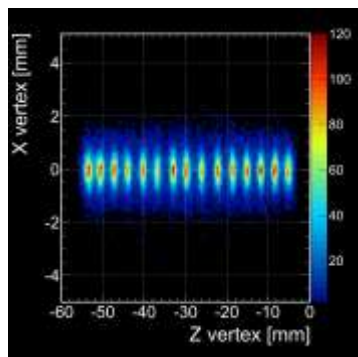
- good pid
- large stat
- good evt vertex



Performance: weak decays

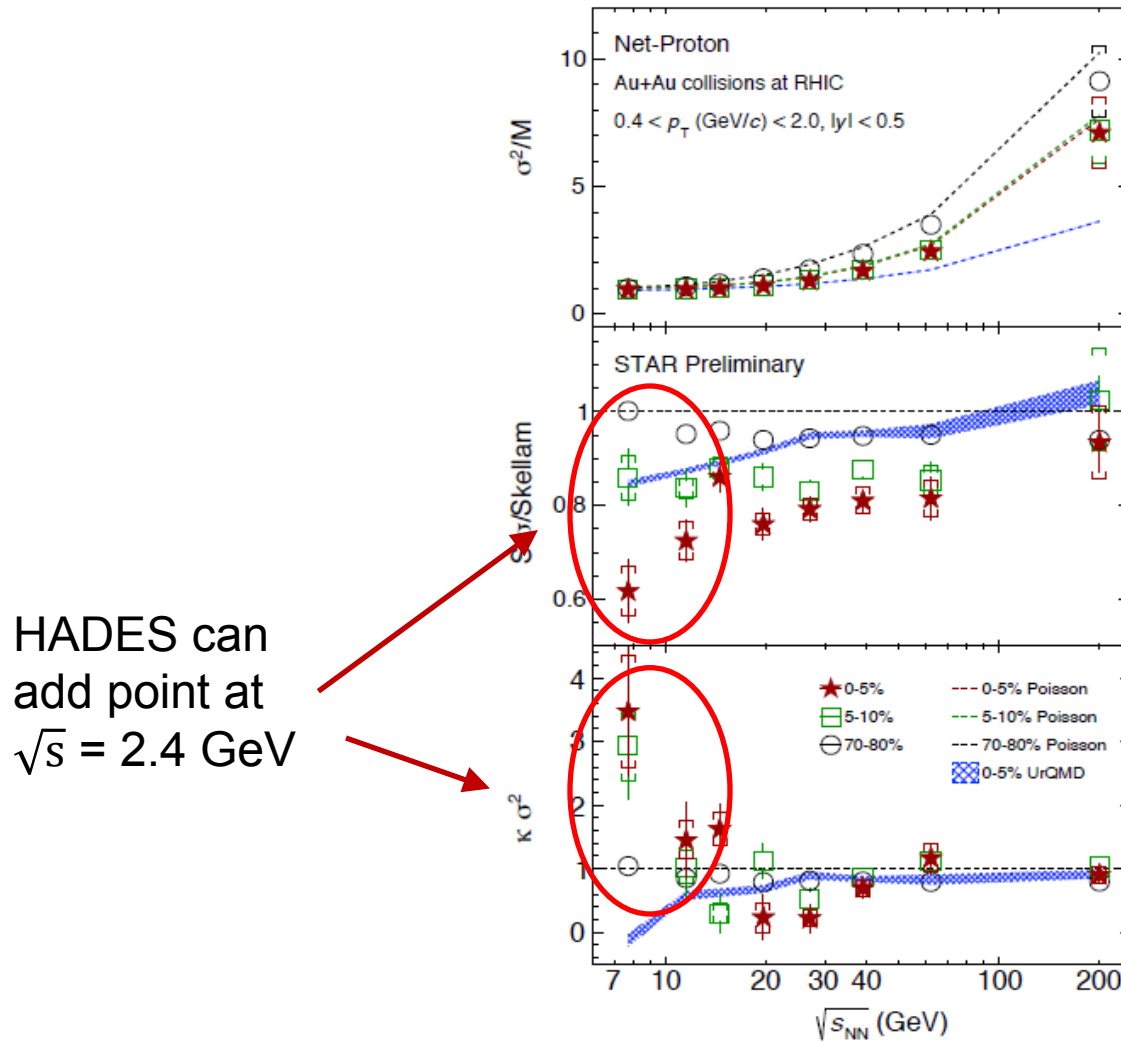


event vertex



HADES vs. RHIC BES

Collision Energy Dependence



Jochen Thäder

Our sandbox: the Poisson distribution

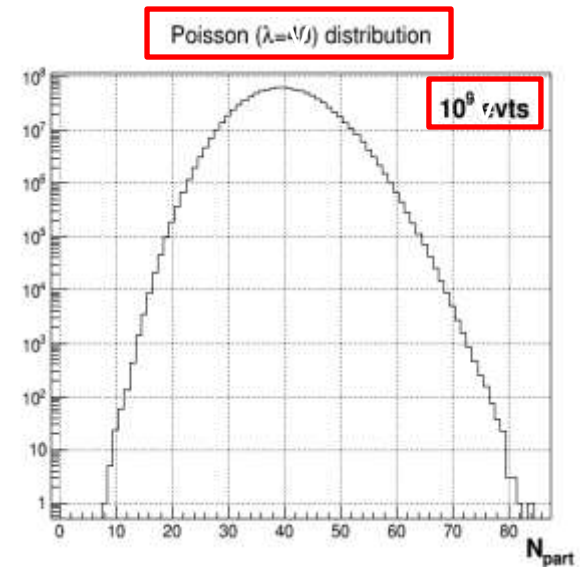
Reminder: for Poisson $P(\lambda)$ we have:

- mean $\mu = \lambda$
- width $\sigma = \sqrt{\lambda}$
- skewness $\text{Sk} = \frac{1}{\sqrt{\lambda}}$
- kurtosis $\kappa = 1/\lambda$

More generally: all cumulants $c_n = \lambda$

It follows that $\frac{c_n}{c_2} = 1$, in particular:

$$\rightarrow \omega = \frac{\sigma^2}{\mu} = 1, \quad \text{Sk} \times \sigma = 1, \quad \kappa \times \sigma^2 = 1$$



Mean	= 40	± 0.0002
Sigma	= 6.3245	± 0.0002
Skewness	= 0.1580	± 0.0002
Kurtosis	= 0.0251	± 0.0002

Omega	= 1.00000	± 0.00004
Skew * sig	= 0.9991	± 0.0005
Kurt * sig2	= 1.003	± 0.007
c5/c2	= 0.99	± 0.11

ok!

Data need efficiency corrections

$$\text{Efficiency} = \text{acc} \times \text{det. eff} \times \text{rec. eff}$$

1. Correct measured distributions (bayesian unfolding)

Garg et al., J. Phys. G: Nucl. Part. Phys. 40 (2013)

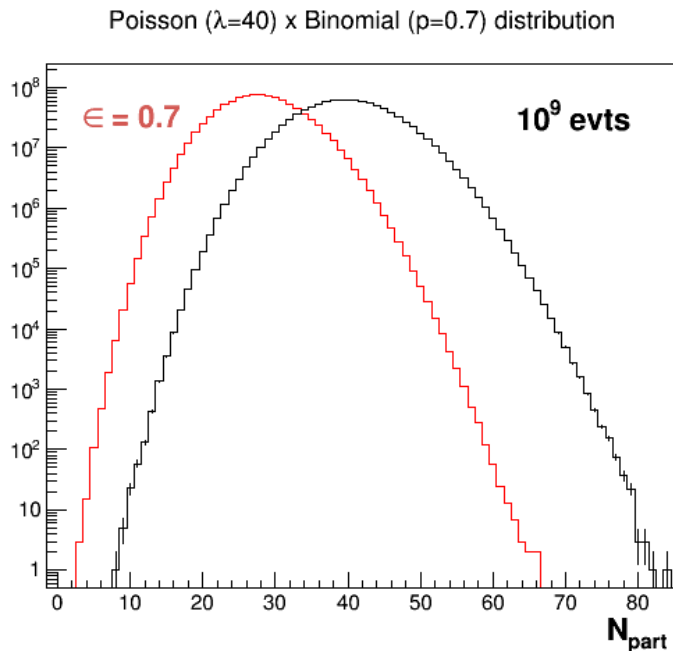
2. Correct the moments

A. Bzdak & V. Koch, PRC 86 (2012); X. Luo, arXiv:1410.3914

→ used by STAR

Experimental bias: Efficiency

With efficiency = 0.7



Mean = 28 ± 0.0002
 Sigma = 5.2915 ± 0.0001
 Skewness = 0.1889 ± 0.0001
 Kurtosis = 0.0358 ± 0.0002

Omega = 0.99998 ± 0.00004
 Skew * Sig = 0.9993 ± 0.0004
 Kurt * Sig² = 1.002 ± 0.005
 c5/c2 = 1.08 ± 0.07

Using the Bzdak-Koch correction procedure:

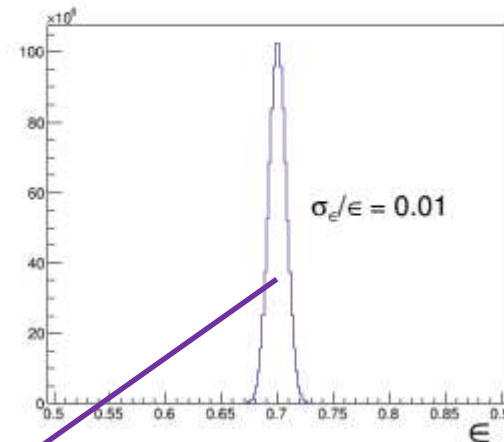
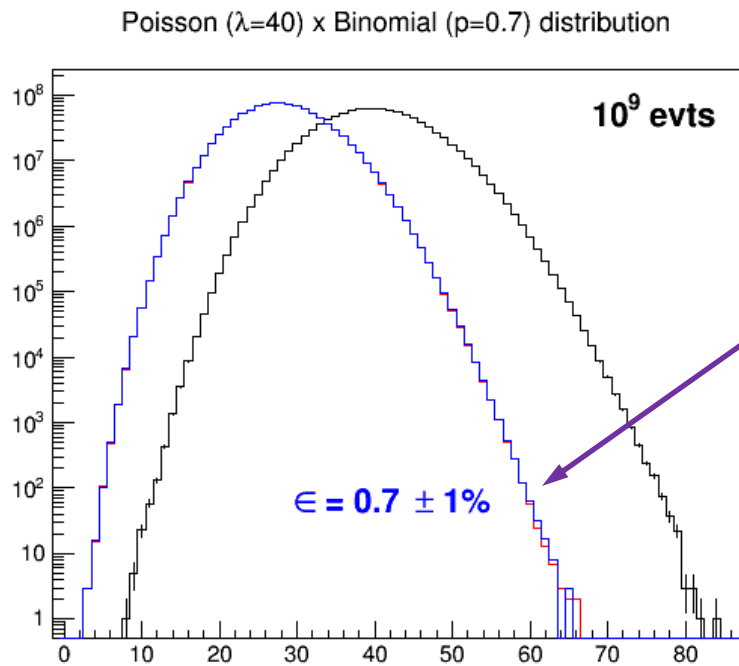
Mean = 40 ± 0.0002
 Sigma = 6.3248 ± 0.0002
 Skewness = 0.1585 ± 0.0001
 Kurtosis = 0.0255 ± 0.0003

Omega = 1.00009 ± 0.00006
 Skew * Sig = 1.0022 ± 0.0008
 Kurt * Sig² = 1.019 ± 0.014
 c5/c2 = 1.05 ± 0.25

ok, but
 errors
 increase!

Evt-by-Evt efficiency changes

Scenario 1: Apply **random changes** of ϵ in event generator, but correct with mean $\langle \epsilon \rangle = 0.7$



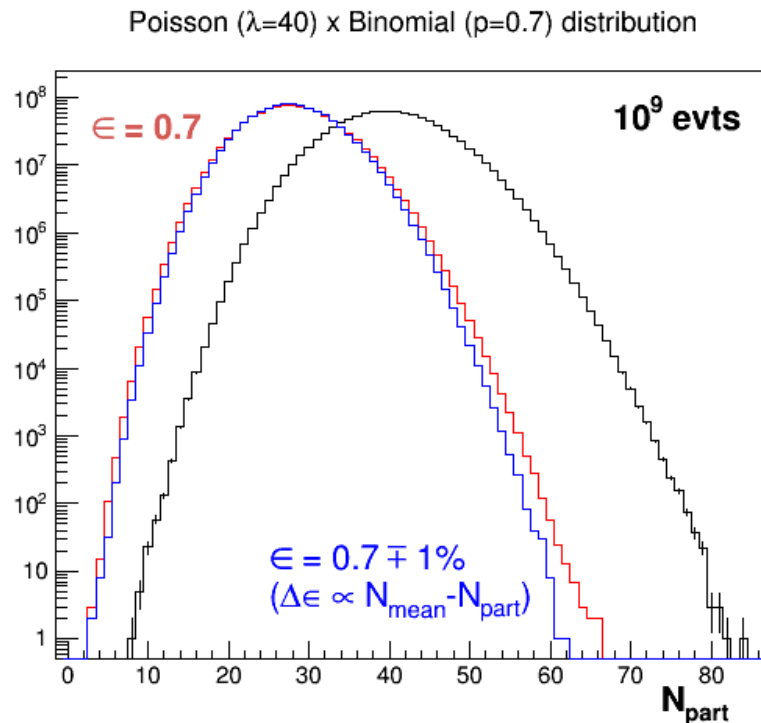
Mean = 40 \pm 0.0002
 Sigma = 6.3398 \pm 0.0002
 Skewness = 0.1592 \pm 0.0002
 Kurtosis = 0.0258 \pm 0.0002

Omega = 1.0050 \pm 0.00006
 Skew * sig = 1.0091 \pm 0.0008
 Kurt * sig2 = 1.036 \pm 0.014
 c5/c2 = 0.93 \pm 0.25

slight broadening!

Evt-by-Evt efficiency changes (II)

Scenario 2: Correlated changes of ϵ with track density: $\pm 1\%$ variation correct with mean $\langle \epsilon \rangle$



Mean = 39.93 \pm 0.0002
 Sigma = 5.8742 \pm 0.0002
 Skewness = 0.1223 \pm 0.0002
 Kurtosis = 0.0082 \pm 0.0004

Omega = 0.86424 \pm 0.00006
 Skew * Sig = 0.7185 \pm 0.0008
 Kurt * Sig² = 0.283 \pm 0.012
 c5/c2 = -0.08 \pm 0.21

→ strong effect !!!

Realistic simulations: UrQMD + Geant3 + event reconstruction

- UrQMD 1.23 GeV/u Au+Au events
 - Geant3 detector simulation
 - Full track reconstruction + particle ID
 - Proton cumulants + Bzdak-Koch correction
in $y = y_0 \pm 0.2$ & $p_t = 0.4 - 1.6$ GeV/c bin
- ➔ compare corrected c_n with UrQMD c_n

Scenario 1: assuming constant efficiencies

UrQMD evts:

Reconstruction corrected
with constant mean $\langle \epsilon \rangle$:

Centrality:

30-40%

Mean = 6.639 ± 0.002
Sigma = 2.892 ± 0.002
Skewness = 0.487 ± 0.002
Kurtosis = 0.245 ± 0.006

Mean = 6.702 ± 0.002
Sigma = 2.842 ± 0.003
Skewness = 0.448 ± 0.004
Kurtosis = 0.190 ± 0.013

Omega = 1.260 ± 0.002
Skew * sig = 1.407 ± 0.005
Kurt * sig2 = 2.051 ± 0.048

Omega = 1.210 ± 0.002
Skew * sig = 1.27 ± 0.01
Kurt * sig2 = 1.53 ± 0.10

→ wrong !

0 – 5%

Mean = 32.57 ± 0.01
Sigma = 5.622 ± 0.005
Skewness = 0.142 ± 0.002
Kurtosis = -0.014 ± 0.006

Mean = 32.86 ± 0.01
Sigma = 5.163 ± 0.005
Skewness = -0.105 ± 0.004
Kurtosis = 0.65 ± 0.03

Omega = 0.970 ± 0.002
Skew * sig = 0.793 ± 0.011
Kurt * sig2 = -0.45 ± 0.17

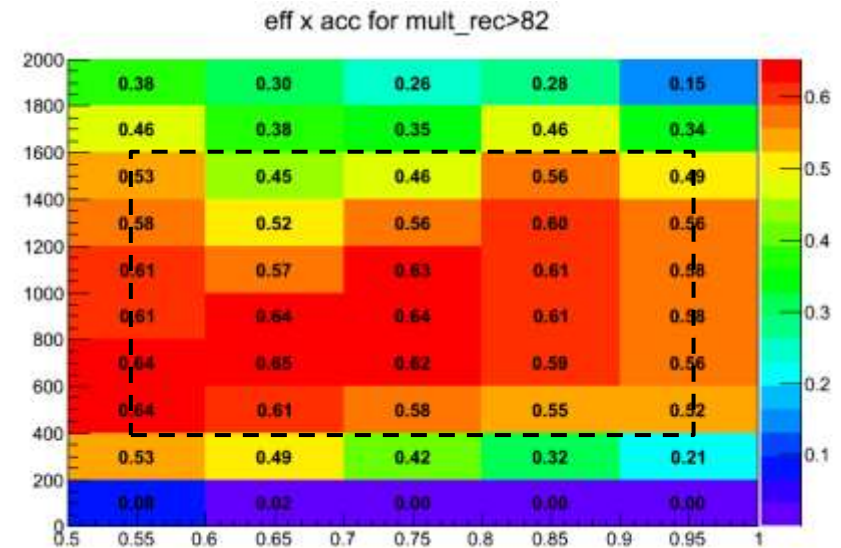
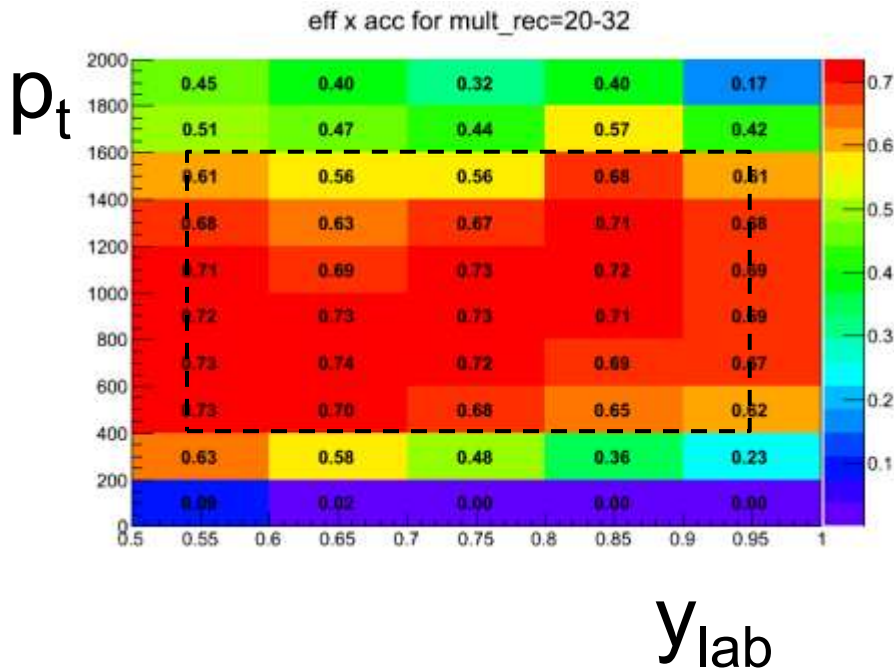
Omega = 0.811 ± 0.002
Skew * sig = -0.54 ± 0.02
Kurt * sig2 = 17.27 ± 0.66

→ badly wrong !

Hades proton efficiencies vs. p_t , y & centrality

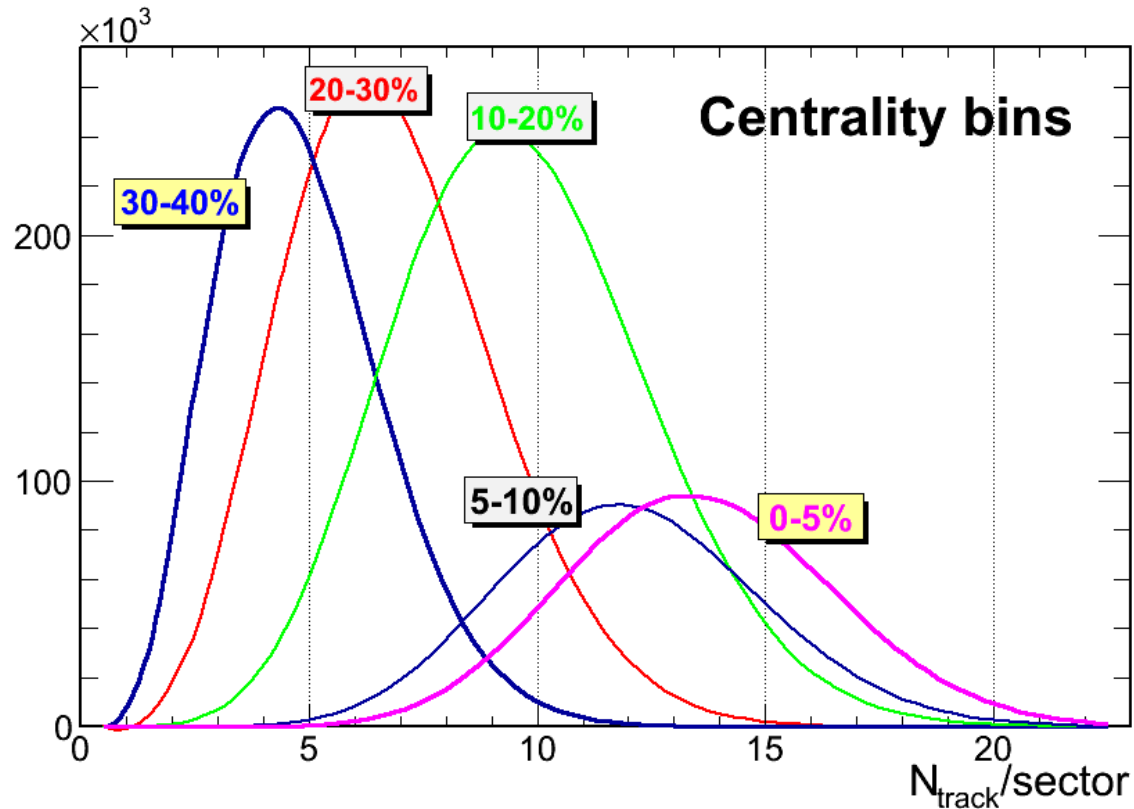
centrality = 30% - 40%

centrality = 0% - 5%



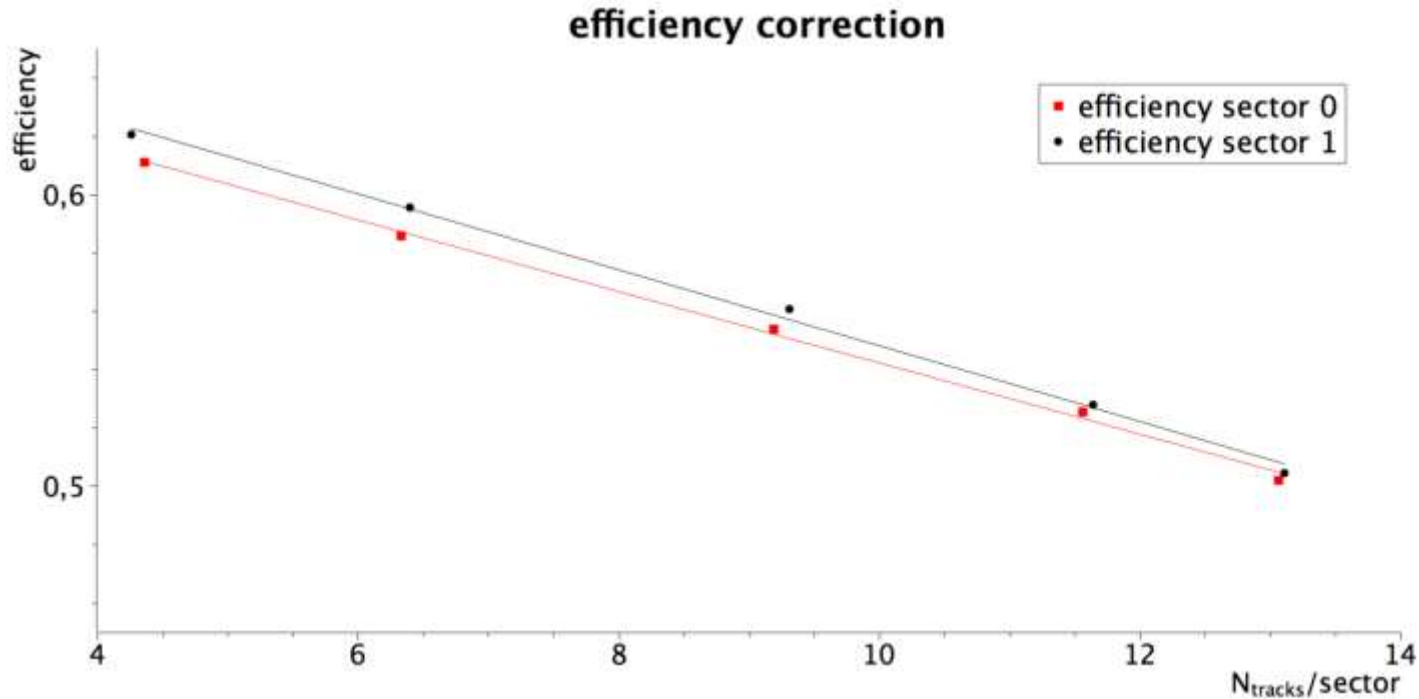
Centrality selection in Au+Au

UrQMD 1.23 GeV/u Au+Au evts into HADES:



→ large evt-to-evt fluctuations of the detector load, even in a given centrality bin

HADES proton efficiency vs. $N_{\text{track}}/\text{sector}$



- Efficiency decreases by 10-15% in most central events
- Use fitted dependence to correct on evt-by-evt basis ?

Generalized efficiency corrections

Efficiency depends on particle, centrality, pt & y...

→ need to correct differentially !

Can be done by using the „factorial moments“:

Bzdak & Koch, Phys. Rev. C 86 (2012)
Xiaofeng Liu, arXiv:1410.3914

$$(1) \quad F_{i,k}(N_p, N_{\bar{p}}) = \left\langle \frac{N_p!}{(N_p - i)!} \frac{N_{\bar{p}}!}{(N_{\bar{p}} - k)!} \right\rangle = \sum_{N_p=i}^{\infty} \sum_{N_{\bar{p}}=k}^{\infty} P(N_p, N_{\bar{p}}) \frac{N_p!}{(N_p - i)!} \frac{N_{\bar{p}}!}{(N_{\bar{p}} - k)!} \quad F_{i,k}(N_p, N_{\bar{p}}) = \frac{f_{i,k}(n_p, n_{\bar{p}})}{(\varepsilon_p)^i (\varepsilon_{\bar{p}})^k}$$

$$f_{i,k}(n_p, n_{\bar{p}}) = \left\langle \frac{n_p!}{(n_p - i)!} \frac{n_{\bar{p}}!}{(n_{\bar{p}} - k)!} \right\rangle = \sum_{n_p=i}^{\infty} \sum_{n_{\bar{p}}=k}^{\infty} p(n_p, n_{\bar{p}}) \frac{n_p!}{(n_p - i)!} \frac{n_{\bar{p}}!}{(n_{\bar{p}} - k)!}$$

$$(2) \quad A_{i,k}(x_1, \dots, x_i; \bar{x}_1, \dots, \bar{x}_k) = \langle N(x_1)[N(x_2) - \delta_{x_1, x_2}] \dots [N(x_i) - \delta_{x_1, x_i} - \dots - \delta_{x_{i-1}, x_i}] \bar{N}(\bar{x}_1)[\bar{N}(\bar{x}_2) - \delta_{\bar{x}_1, \bar{x}_2}] \dots [\bar{N}(\bar{x}_k) - \delta_{\bar{x}_1, \bar{x}_k} - \dots - \delta_{\bar{x}_{k-1}, \bar{x}_k}] \rangle$$

„local factorial moments“

$$a_{i,k}(x_1, \dots, x_i; \bar{x}_1, \dots, \bar{x}_k) = \langle n(x_1)[n(x_2) - \delta_{x_1, x_2}] \dots [n(x_i) - \delta_{x_1, x_i} - \dots - \delta_{x_{i-1}, x_i}] \bar{n}(\bar{x}_1)[\bar{n}(\bar{x}_2) - \delta_{\bar{x}_1, \bar{x}_2}] \dots [\bar{n}(\bar{x}_k) - \delta_{\bar{x}_1, \bar{x}_k} - \dots - \delta_{\bar{x}_{k-1}, \bar{x}_k}] \rangle$$

$$(3) \quad F_{i,k} = \sum_{x_1, \dots, x_i} \sum_{\bar{x}_1, \dots, \bar{x}_k} A_{i,k}(x_1, \dots, x_i; \bar{x}_1, \dots, \bar{x}_k)$$

$$f_{i,k} = \sum_{x_1, \dots, x_i} \sum_{\bar{x}_1, \dots, \bar{x}_k} a_{i,k}(x_1, \dots, x_i; \bar{x}_1, \dots, \bar{x}_k)$$

$$F_{i,k} = \sum_{x_1, \dots, x_i} \sum_{\bar{x}_1, \dots, \bar{x}_k} \frac{a_{i,k}(x_1, \dots, x_i; \bar{x}_1, \dots, \bar{x}_k)}{\epsilon(x_1) \dots \epsilon(x_i) \bar{\epsilon}(\bar{x}_1) \dots \bar{\epsilon}(\bar{x}_k)}$$

→ correct evt-by-evt
(assuming binomial !)

Scenario 2: efficiency depends on N_{track}

UrQMD evts:

Reconstruction corrected
with mean $\langle \epsilon \rangle$:

Centrality:

30-40%

Mean = 6.639 ± 0.002
 Sigma = 2.892 ± 0.002
 Skewness = 0.487 ± 0.002
 Kurtosis = 0.245 ± 0.006

Mean = 6.714 ± 0.003
 Sigma = 2.935 ± 0.003
 Skewness = 0.491 ± 0.003
 Kurtosis = 0.252 ± 0.012

Omega = 1.260 ± 0.001
 Skew * sig = 1.407 ± 0.005
 Kurt * sig2 = 2.051 ± 0.048

Omega = 1.283 ± 0.002
 Skew * sig = 1.440 ± 0.009
 Kurt * sig2 = 2.17 ± 0.10

→ ok !

0 – 5%

Mean = 32.57 ± 0.01
 Sigma = 5.622 ± 0.005
 Skewness = 0.142 ± 0.002
 Kurtosis = -0.014 ± 0.006

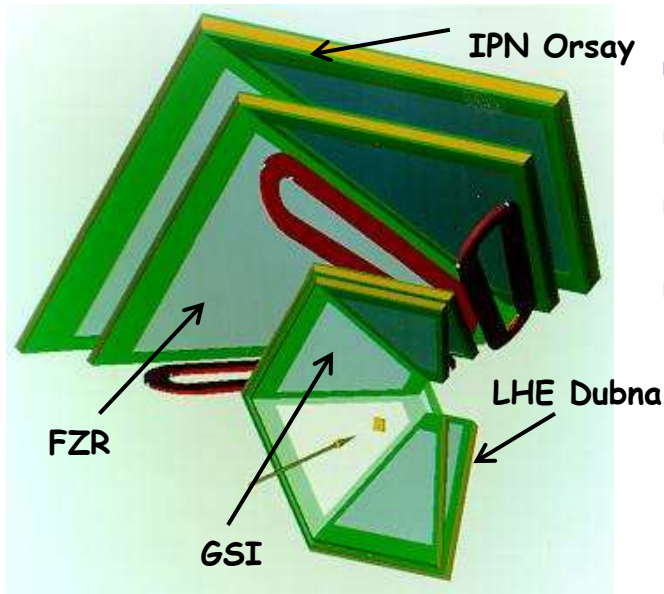
Mean = 32.53 ± 0.01
 Sigma = 5.743 ± 0.009
 Skewness = -0.004 ± 0.007
 Kurtosis = 0.390 ± 0.026

Omega = 0.970 ± 0.002
 Skew * sig = 0.793 ± 0.011
 Kurt * sig2 = -0.45 ± 0.17

Omega = 1.014 ± 0.003
 Skew * sig = -0.022 ± 0.043
 Kurt * sig2 = 12.86 ± 0.86

→ still off,
but better !

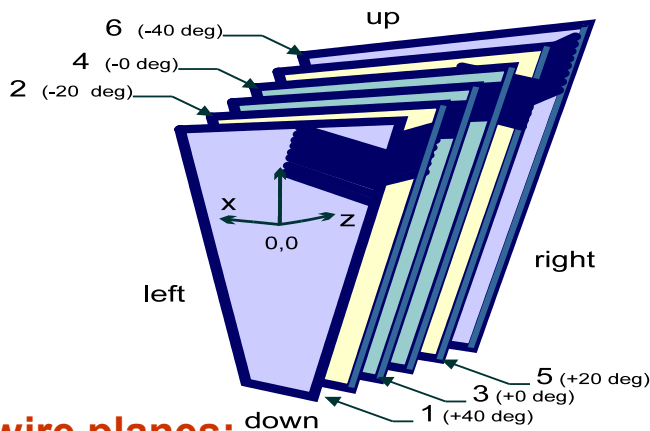
Tracking: the Multiwire Drift Chambers (MDC)



- 4 MDC/sector
- total 33 m² area, 27000 cells
- $\Delta y < 0.1$ mm resolution
- Ar-iC₄H₁₀ [60-40] gas and low-Z material



ORSAY plane

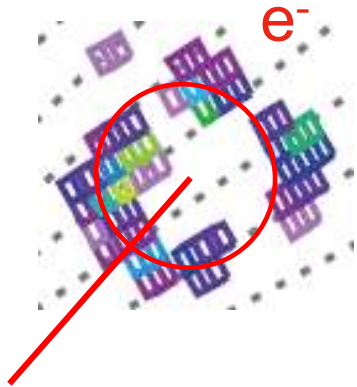


wire planes:
40°, -20°, 0°, 0°, 20°, -40°

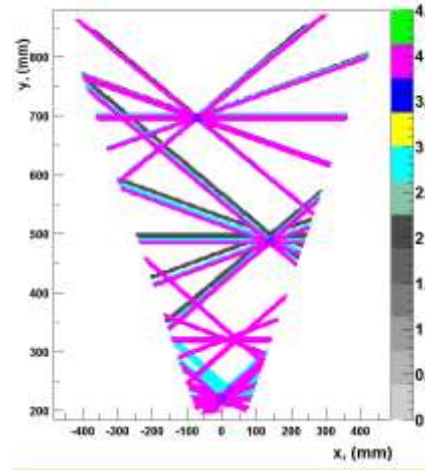
Layer	Width/m	Height	Area[m ²]
I	76,7	75,5	0,34
II	90,5	88,3	0,49
III	180,5	178,0	1,88
IV	222,4	219,9	2,83

Electron/positron identification

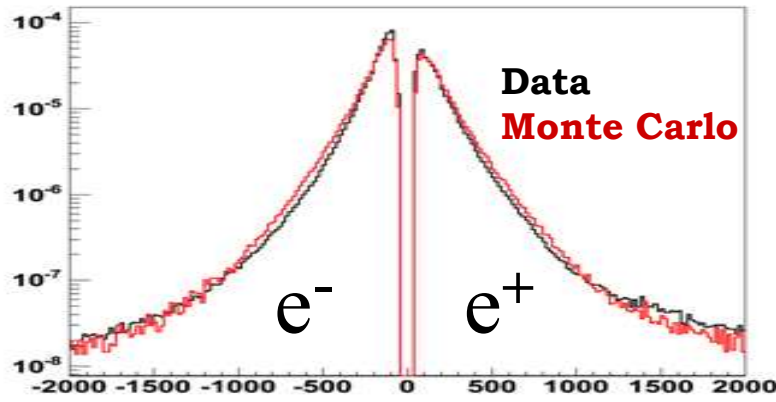
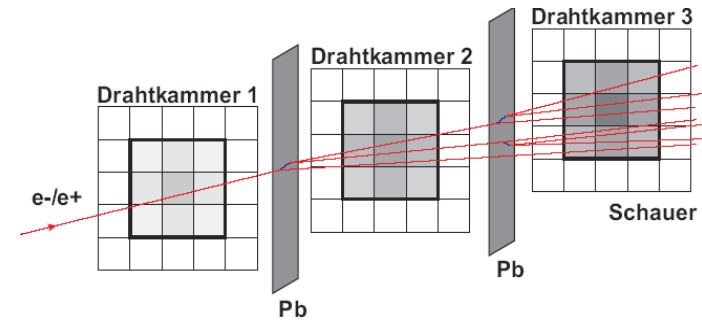
RICH pattern



MDC hit finder & hit/track matching

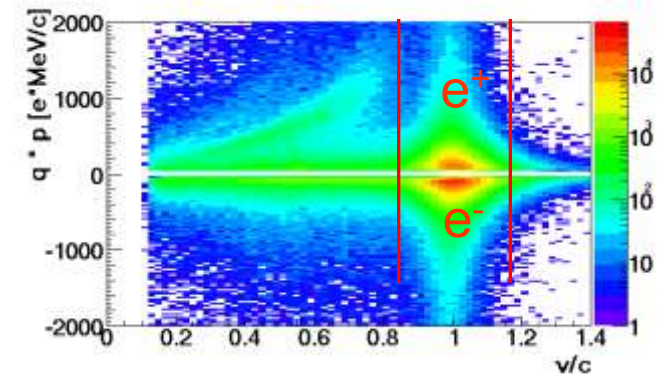


Pre-Shower condition



Momentum * charge [MeV/C]

velocity vs. momentum



Scenario 3: $\text{eff} = \text{eff}(N_{\text{track}}, N_{\text{wire}}/N_{\text{track}})$

UrQMD evts:

Reconstruction corrected
with mean $\langle \epsilon \rangle$:

Centrality:

30-40%

Mean = 6.639 ± 0.002
 Sigma = 2.892 ± 0.002
 Skewness = 0.487 ± 0.002
 Kurtosis = 0.245 ± 0.006

Mean = 6.967 ± 0.003
 Sigma = 2.966 ± 0.003
 Skewness = 0.478 ± 0.004
 Kurtosis = 0.244 ± 0.017

Omega = 1.260 ± 0.001
 Skew * sig = 1.407 ± 0.005
 Kurt * sig2 = 2.051 ± 0.048

Omega = 1.263 ± 0.003
 Skew * sig = 1.412 ± 0.011
 Kurt * sig2 = 2.15 ± 0.15

→ ok !

0 – 5%

Mean = 32.57 ± 0.01
 Sigma = 5.622 ± 0.005
 Skewness = 0.142 ± 0.002
 Kurtosis = -0.014 ± 0.006

Mean = 33.57 ± 0.01
 Sigma = 5.840 ± 0.010
 Skewness = 0.056 ± 0.010
 Kurtosis = 0.230 ± 0.029

Omega = 0.970 ± 0.002
 Skew * sig = 0.793 ± 0.011
 Kurt * sig2 = -0.45 ± 0.17

Omega = 1.016 ± 0.003
 Skew * sig = 0.328 ± 0.060
 Kurt * sig2 = 7.83 ± 1.00

→ even better !

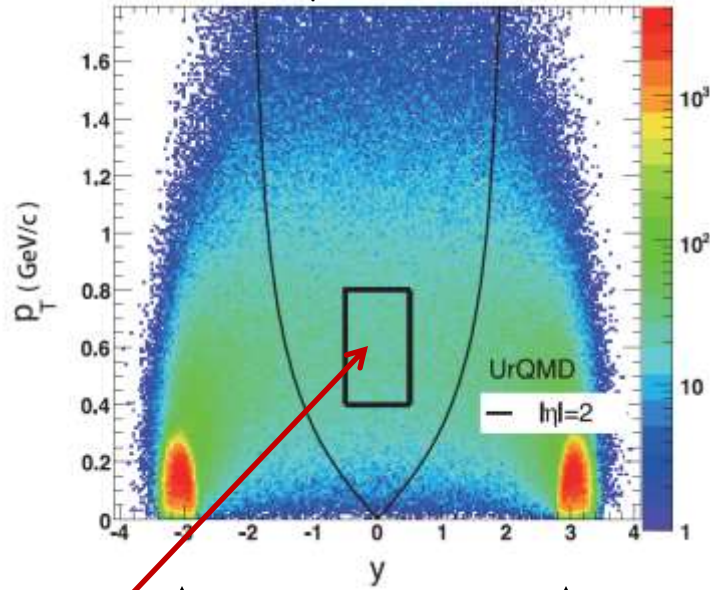
Things we are still investigating ...

- Centrality selection: META vs. FW → avoid autocorrelation
- Centrality bin width correction → remove volume fluctuations
- Event pile-up → avoid/remove contamination
- Track density dependence of efficiency correction
 - strong detector dependence
 - check independence assumption
- Analyze Au+Au data
- Role of fragments (d,t,He) → do they modify $\delta(\Delta N_p)$?

Back up slides

Which phase-space bite to use?

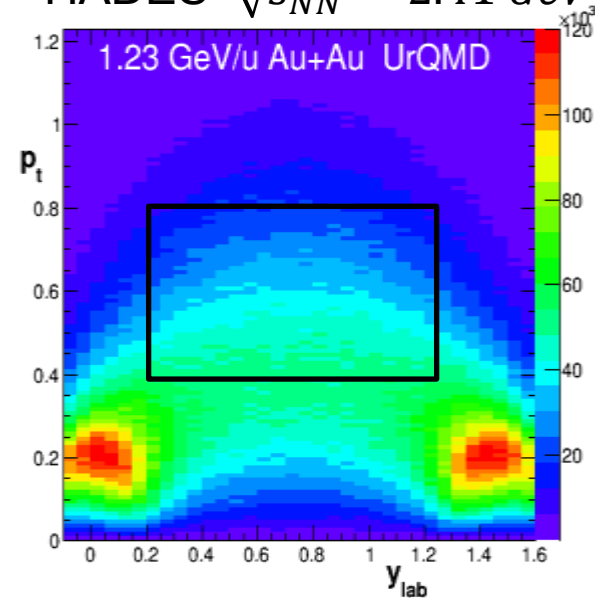
STAR $\sqrt{s_{NN}} = 19.6 \text{ GeV}$



rapidity gap = 6 units!

STAR phase space bin: $y = y_0 \pm 0.5$
 $p_t = 0.4 - 0.8 \text{ GeV}/c$
 $(p_t = 0.4 - 2.0 \text{ GeV}/c)$

HADES $\sqrt{s_{NN}} = 2.41 \text{ GeV}$



rapidity gap = 1.5 units!

➔ Need to reduce the y bite:
 $\pm 0.5 \rightarrow \leq \pm 0.25$

Towards higher-order cumulants?

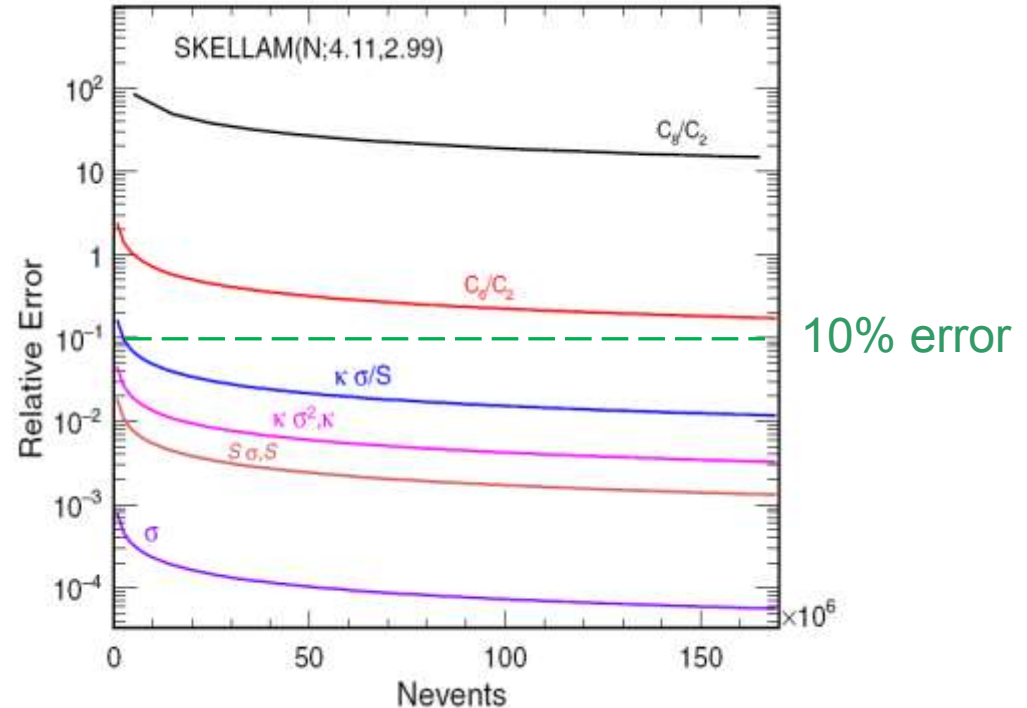
Sensitivity to critical fluctuations is expected to increase with order, but same is true for stat. flucs., etc.

$$\text{error}(\hat{S}\hat{\sigma}) \propto \frac{\sigma}{\sqrt{n}}$$

$$\text{error}(\hat{\kappa}\hat{\sigma}^2) \propto \frac{\sigma^2}{\sqrt{n}}$$

$$\text{error}\left(\frac{\hat{C}_6}{\hat{C}_2}\right) \propto \frac{\sigma^4}{\sqrt{n}}$$

$$\text{error}\left(\frac{\hat{C}_8}{\hat{C}_2}\right) \propto \frac{\sigma^6}{\sqrt{n}}$$



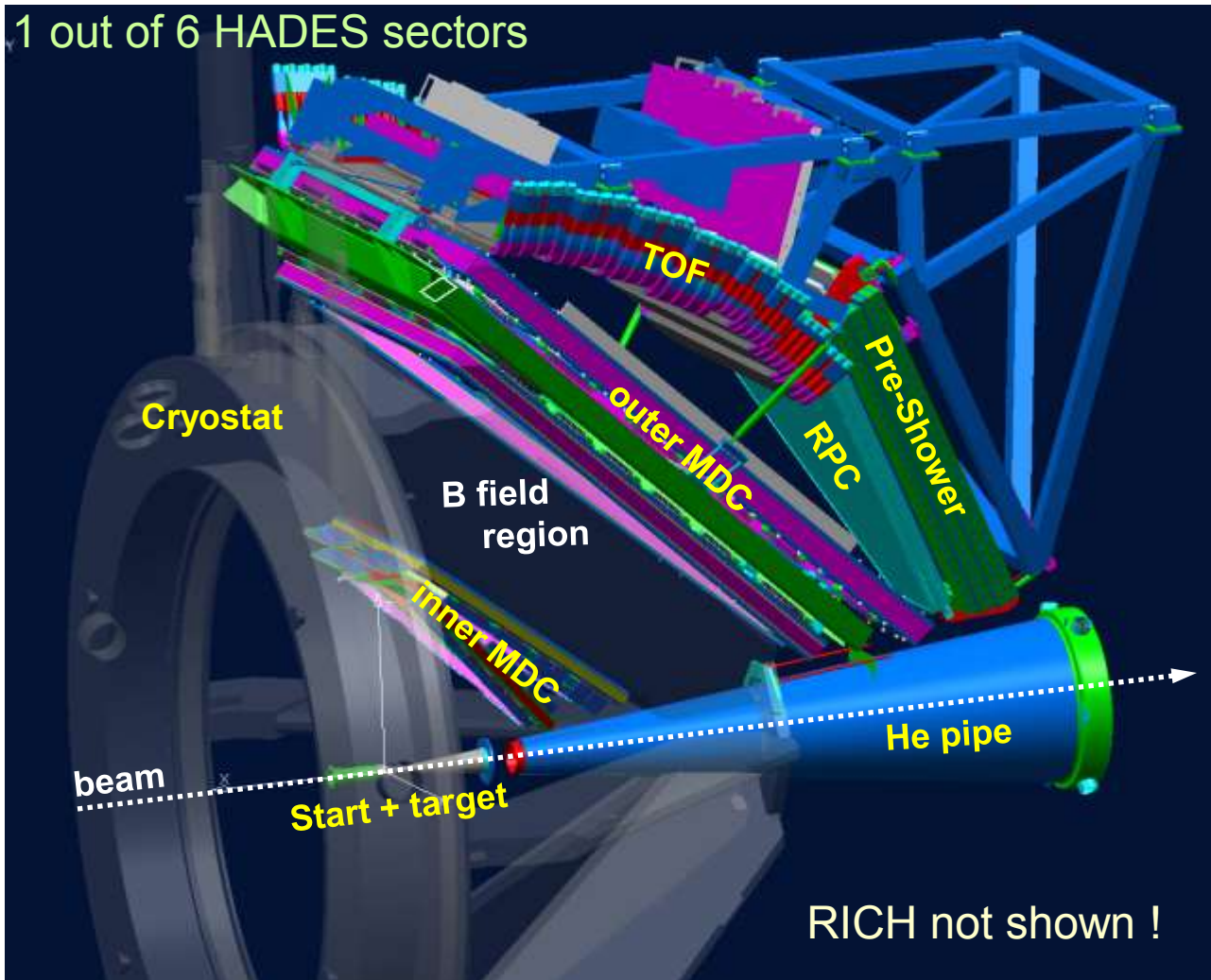
→ <10% stat. error needs $\approx 10^9$ evts for 6th order, and $\approx 10^{13}$ for 8th order !

(Note: corrections for limited eff & acc will further increase the error)

Technical layout of HADES

HADES

1 out of 6 HADES sectors



HADES + FW



inner MDC

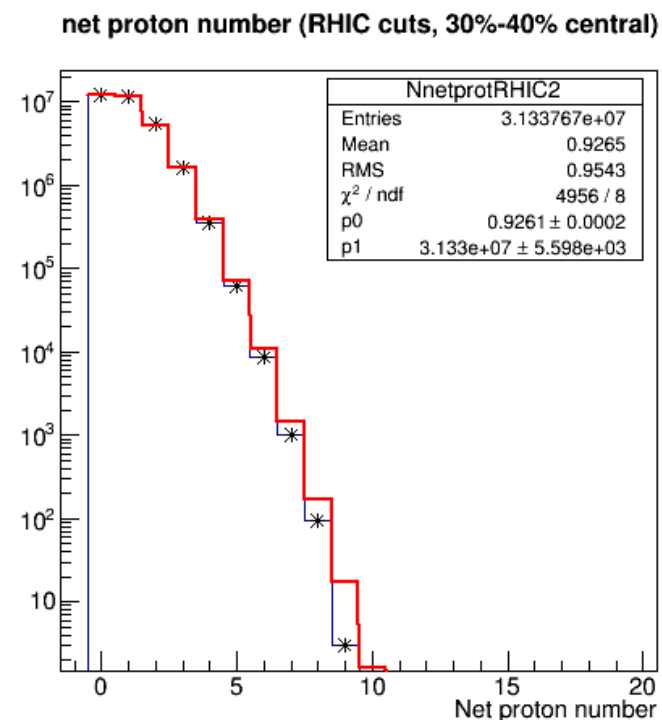
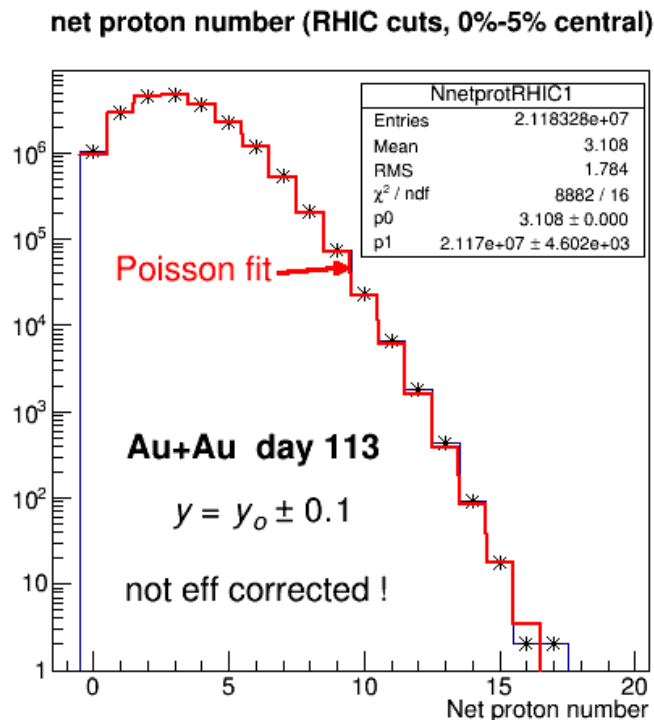


RICH readout



A look at HADES data: Au+Au 1.23 GeV/u

2 different centrality cuts, $y = y_0 \pm 0.1$, $p_t = 0.4 - 0.8 \text{ GeV}/c$



➔ Data look Poisson-like, but not quite ...

Characterizing a distribution by its moments

(or alternatively by its cumulants)

A distribution $f(x)$ is fully characterized by its (central) moments:

$$\mu_n = \text{E} [(X - \text{E}[X])^n] = \int_{-\infty}^{+\infty} (x - \mu)^n f(x) dx.$$

$n=0, 1, 2, 3, 4, \dots, \infty$

- $n=0$: normalization = μ_0
- $n=1$: **mean** = μ_1
- $n=2$: **variance** = μ_2 (or $\sigma = \sqrt{\mu_2}$) measures width
- $n=3$: **skewness** = $\frac{\mu_3}{\sigma^3}$ measures asymmetry
- $n=4$: **kurtosis** = $\frac{\mu_4}{\sigma^4} - 3$ measures pointedness/flatness
- $n>4$: ...

A few common examples

	Mean	Variance	Skewness	Kurtosis
Gauss	μ	σ^2	0	0
Binomial	np	$np(1-p)$	$\frac{1-2p}{\sqrt{np(1-p)}}$	$\frac{1-6p(1-p)}{np(1-p)}$
Poisson	μ	μ	$\mu^{-1/2}$	μ^{-1}
Skellam	$\mu_1 - \mu_2$	$\mu_1 + \mu_2$	$\frac{\mu_1 - \mu_2}{(\mu_1 + \mu_2)^{3/2}}$	$\frac{1}{\mu_1 + \mu_2}$

For easier comparison with Skellam (and to reduce the so-called volume effect), one often computes:

$$Sk * \sigma = \frac{\mu_1 - \mu_2}{\mu_1 + \mu_2} \text{ and } \kappa * \sigma^2 = 1$$

At low beam energies, $\mu_2 = N_{\bar{p}} = 0$ and $Sk * \sigma = \kappa * \sigma^2 = 1!$

Ultimate goal: Compare with STAR data

7 c.m. energies, 3 centrality selections, within a defined y - p_t bin:

STAR @ RHIC PRL 112, 032302 (2014)

