Robust Integration of Fast Flavor Conversions in Classical Neutrino Transport

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The quantum kinetic evolution of neutrinos in dense environments, such as the core-collapse supernovae or the neutron star mergers, can result in fast flavor conversion (FFC), presenting a significant challenge to achieving robust astrophysical modeling of these systems. Recent works that directly simulate the quantum kinetic transport of neutrinos in localized domains have suggested that the asymptotic outcome of FFCs can be modeled by simple analytical prescriptions when coarse grained over a size much larger than the FFC length scale. In this *Letter*, by leveraging such a scale separation, we incorporate the analytical prescriptions into global simulations that solve the classical neutrino transport equation including collisions and advection under spherical symmetry. We demonstrate that taking this approach allows to obtain results that *quantitatively* agree with those directly from the corresponding global quantum kinetic simulations and precisely capture the collisional feedback effect for cases where the FFC happens inside the neutrinosphere. Notably, the effective scheme does not require resolving the FFC time and length scales, hence only adds negligible computational overhead to classical transport. Our work highlights that efficient and robust integration of FFCs in classical neutrino transport used in astrophysical simulation can be feasible.

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Introduction—The robustness of astrophysical simulations for explosive events such as core-collapse supernovae (CCSNe) and neutron star mergers (NSMs) relies on the accurate numerical modeling of the nonequilibrium classical neutrino transport from the neutrino-trapped to free-streaming regime [1–8]. However, recent studies have revealed that the quantum mechanical nature of neutrinos, as manifested in their flavor oscillations, can lead to various collective modes of flavor conversions in regions where neutrinos decouple from matter [9–11]. These collective phenomena, originated from the nonlinear self-interaction of neutrino forward scattering, may significantly affect the theoretical modeling and the aftermath of CCSNe and NSMs, thus emphasizing the need for quantum kinetic transport of neutrinos [12–18].

In particular, the so-called fast flavor instability (FFI) [19–22] can exist at locations where the angular distribution of the neutrino electron lepton number (ELN) density contains a crossing, i.e., the ELN angular distribution function changes sign [23,24], and is commonly found in the interior of CCSNe and NSM remnants [25–37]. The presence of FFI can lead to the emergence of fast flavor conversion (FFC) and alter the neutrino flavor content within a timescale of subnanosecond and a length scale of subcentimeter, much shorter than the typical hydrodynamical or interaction time and length scales governed by classical processes.

The large separation of scales implies an exceedingly high computational cost for a direct implementation of neutrino quantum kinetic transport in simulations of CCSNe and NSMs. Intensive studies aimed to obtain solutions to this problem have been explored. A large number of recent works have numerically solved the neutrino quantum kinetic equation (ν QKE) in localized simulation domains to understand the outcome of FFCs and found that the ELN crossings are erased on a coarsegrained sense [38-54]. Attempts to model the coarsegrained FFC outcome based on these simulations using analytical formulas [39,49,51,55] or through the use of machine learning [56,57] have proven that accurate outcomes relevant to the modeling of neutrino transport can be obtained. Global transport simulations solving νQKE with different simplifications, e.g., using artificially quenched neutrino self-interaction or adopting spatial resolutions lower than the FFC length scales, have been performed [58-67] and provided improved understanding of the emergence and evolution of FFCs in static background. On the other hand, preliminary efforts to include the FFC

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outcome in parametric fashions in hydrodynamical simulations equipped with classical neutrino transport solvers were conducted in [68–72]. Meanwhile, new theoretical frameworks including the subgrid methods, "miscidynamics," and quasilinear approach have been proposed in [52,73–76].

Despite all these tremendous efforts, an important question that remains to be answered is whether or not the coarse-grained FFC outcome learned from local ν QKE simulations can be effectively included in classical transport used in astrophysical simulations based on the principle of separation of scales. In this *Letter*, we tackle this question by directly incorporating the coarse-grained FFC prescriptions into global simulations that solve the classical neutrino transport equation under spherical symmetry. For the first time, we show *robust* agreement between results from this approach and those directly solving the ν QKE in our earlier work [64], which reveals the viability of effective integration of FFC in astrophysical simulations.

Model—We use the extended version [64] of $COSE\nu$ [77] to numerically solve the ν QKE and the corresponding effective classical transport (ECT) under static and spherically symmetric background using discrete ordinates method in Minkowski spacetime. The code includes the effective Hamiltonian terms governing flavor oscillations of neutrinos and their interactions with matter and can reproduce the general properties of classical neutrino transport results obtained from SN simulations using the AGILE-BOLTZTRAN code [78–80]. All ν QKE and the ECT simulations are performed in a radial range from r =20 km to 80 km, spanning both the neutrino-trapped and free-streaming regimes for spherically symmetric CCSN background profiles based on a $25M_{\odot}$ progenitor as used in [64] with two-flavor scheme. We focus on Models II, III, and IV of [64,81]. These models take different electron fraction profiles, which result in different number density ratios of $n_{\bar{\nu}_e}/n_{\nu_e}$ and ELN angular distributions. In Models II and III, $n_{\nu_e} \gtrsim n_{\bar{\nu}_e}$ at all radii without FFCs, while $n_{\nu_e} \lesssim n_{\bar{\nu}_e}$ in Model IV, representing the ν_e and $\bar{\nu}_e$ dominating cases, respectively [82]. Moreover, the FFIs exist inside the neutrinospheres in Models III and IV, which leads to an intriguing collisional feedback effect that continuously alters the evolution of neutrino distributions after the prompt FFC.

To perform ν QKE simulations within a reasonable time frame, we take an attenuation function $a_{\nu\nu}(r)$ to quench the strength of the neutrino-neutrino forwarding scattering Hamiltonian, $\mathbf{H}_{vv}(v_r) = \sqrt{2}G_F \int dE' dv'_r (1 - v_r v'_r) [\varrho(E', v'_r) - \bar{\varrho}^*(E', v'_r)]$ with G_F the Fermi constant, E and v_r the energy and the radial velocity of neutrinos, and ϱ and $\bar{\varrho}$ the density matrices of neutrino and antineutrinos. Natural units with $\hbar = c = 1$ are adopted throughout this Letter. The diagonal elements of ϱ and $\bar{\varrho}$ are related to $n_{\nu_i}(r, t) = \int dE dv_r \bar{\varrho}_{ii}$ for i = e, x with xdenoting the heavy-lepton flavor. The attenuation function takes the form $a_{\nu\nu}(r) = a/[1 + e^{(30-r[\text{km}])/2.5}]$, such that $a_{\nu\nu}(r) \approx a$ at r > 35 km and $a_{\nu\nu}(r) > 0.1a$ at r > 25 km for the radial range where FFC occurs. We take two different attenuation factors $a = 4 \times 10^{-3}$ and $a = 10^{-3}$. Even with the quenched \mathbf{H}_{vv} , the ν QKE simulations need 50,000 (25,000) uniform radial grids for larger (smaller) value of *a*, labeled as the " ν QKE-H" (" ν QKE-L") models in the rest of the Letter, to resolve the attenuated FFC scale. For the time step size, we take $\Delta t = 1.6$ ns (3.2 ns) for the " ν QKE-H" (" ν QKE-L") models. Details of simulation parameters are given in [82].

For the ECT simulations, we assume that when the FFI exists at any radius, i.e., when there exists an ELN angular crossing, the FFC quickly relaxes the local neutrino flavor content to a quasistationary state when averaging over a length scale much larger than the FFC one, within a timescale much shorter than the collisional and the advection timescales. In the quasistationary state, the coarsegrained ELN angular distribution does not contain any crossings [39,41,49,51,75], and can be approximated by certain analytical prescriptions, where flavor equilibration takes place in a part of the angular domain [49,51], with the off-diagonal elements of ρ and $\bar{\rho}$ averaging to ~0. Based on the above assumptions rooted from the scale separation nature of the problem, we then take the following practical procedures to effectively incorporate the FFC in the classical transport.

First, the coarse-grained nature allows (and demands) us to take a radial grid size much larger than the size of \mathbf{H}_{nn} . Thus, in all our ECT models, we use 250 radial grids in the same simulation domain and a larger $\Delta t = 32$ ns. Second, we only evolve the diagonal elements ρ_{ii} and \bar{q}_{ii} by setting the coherent propagation Hamiltonian term, as well as the off-diagonal entries in the collision term, to zero. Third, at each time step, we check whether there exist angular crossings in the ELN function $G(v_r) =$ $\langle \varrho_{ee} \rangle_E - \langle \bar{\varrho}_{ee} \rangle_E - \langle \varrho_{xx} \rangle_E + \langle \bar{\varrho}_{xx} \rangle_E$, where $\langle \cdot \rangle_E$ denotes the energy-integrated quantities at all radial grids. If an ELN crossing is found at a given radius, we then assume that immediate flavor redistribution takes place due to FFC and apply the power-1/2 prescription scheme derived in [51] to calculate the redistributed diagonal elements in ρ and $\bar{\varrho}$. Specifically, under the two-flavor scenario, the energy-independent survival probability for (anti)neutrinos is given by

$$P(v_r, w) = \begin{cases} \frac{1}{2} & \text{for } v_<, \\ 1 - \frac{1}{2}h(|v_r - v_c|/w) & \text{for } v_>, \end{cases}$$
(1)

where the function $h(x) = (x^2 + 1)^{-1/2}$ and v_c is the crossing velocity at which $G(v_c) = 0$. In Eq. (1), $v_<(v_>)$ is defined as the v_r range over which the absolute value of the positive or negative ELN, $|I_+|$ or $|I_-|$, is smaller (larger). The quantities I_{\pm} are defined as

 $\int dv_r G(v_r)\Theta[\pm G(v_r)]$, with Θ denoting the Heaviside function. The parameter *w* can be calculated so that the conservation of ELN holds, i.e., $\int dv_r [2P(v_r, w) - 1]G(v_r) = I_+ + I_-$. Finally, with Eq. (1), the neutrino flavor content $\varrho_{ii}^{(f)}$ after the FFC are replaced by

$$\varrho_{ee} \rightarrow \varrho_{ee}^{(f)} = \varrho_{ee}P + \varrho_{xx}(1-P),$$

$$\varrho_{xx} \rightarrow \varrho_{xx}^{(f)} = \varrho_{ee}(1-P) + \varrho_{xx}P$$
(2)

for all energy grids at the radial grids where sizable ELN crossings with $(I_{<}/I_{>}) > 10^{-3}$ are found, where $I_{<} = \min(|I_{+}|, |I_{-}|)$ and $I_{>} = \max(|I_{+}|, |I_{-}|)$. A similar redistribution applies to antineutrinos as well. We also note that Eq. (1) only applies to cases with at most a single angular crossing at any given radius, which is true in all models examined in this work. Besides the procedures outlined above, all other settings in the ECT simulations are identical to those in the corresponding ν QKE simulations.

We run all the ν QKE and the ECT simulations up to 320 µs when the systems have settled into the asymptotic states. The large reduction of radial grid numbers and the increased Δt in the ECT simulations significantly reduce the computation time by $\geq O(10^3)$ from the ν QKE runs for all cases (see [82]). The ν QKE results shown in all figures are coarse-grain averaged over a size of 0.6 km.

Results—Figure 1 compares the asymptotic radial profiles of the number density ratio between cases with FFCs and those with no flavor conversion, denoted as $n_{\nu}^{\text{FFC}}/n_{\nu}^{\text{NFC}}$ for all models. The FFC leads to more enhanced n_{ν_x} in Model IV, followed by Models III and II in the ν QKE simulations. For ν_e and $\bar{\nu}_e$, since the FFC takes place at radii inside neutrinospheres in Models III and IV, collisions inside neutrinospheres can repopulate their number densities after the prompt phase of FFC and work along with the subsequent FFC to largely alter the ELN distribution [see Fig. 2(d) and the discussion below] [64]. Consequently, the amount of reduction in n_{ν_e} and $n_{\bar{\nu}_e}$ are less so that their number density ratios in Models III and IV become comparable to those in Model II. All changes of neutrino flavor content can alter their net heating rates and affect the fate of SN explosions [62,64,71,72].

Comparing the ECT results to the ν QKE ones, it clearly shows remarkably *quantitative* agreement in all three models. In particular, the differences in number density ratios between the ECT and ν QKE-H models are only up to $\lesssim 5\%$ in Model II and are even smaller than 2% for most radii in Models III and IV, where the collisional feedback effects are important. For the ν QKE-L models, slightly larger differences exist around $r \simeq 45$ km in Model II and $r \simeq 37$ km in Model IV.

The differences in Model II are primarily associated with the flavor overconversion of the forward propagating neutrinos during the FFC, which occurs before the ELN distribution settles down to the quasistationary state that can be approximated by Eq. (1). With the lower attenuation factor applied in the ν QKE-L model (more attenuated), the lengthened FFC timescale prevents FFC from reaching the quasistationary state before advection takes effect. Consequently, a larger amount of flavor overconversion is imprinted in the asymptotic state in Model II-vQKE-L than in Model II-vQKE-H. In Model IV, the negative ELN dominated by $\bar{\nu}_e$'s before FFCs changes to positive due to the collisional feedback. The strong attenuation of $H_{\nu\nu}$ in ν QKE-L model can cause the ELN to switch back to negative at $r \sim 30-37$ km, resulting in a sharp transition at $r \simeq 37$ km. However, in the ν QKE-H model where less



FIG. 1. Comparison of the radial profiles of neutrino number density ratios $n_{\nu}^{\text{FFC}}/n_{\nu}^{\text{NFC}}$ obtained in ν QKE simulations using different attenuation factors $a = 10^{-3}$ (yellow, labeled by " ν QKE-L") and $a = 4 \times 10^{-3}$ (blue, labeled by " ν QKE-H") with the corresponding classical transport simulation outcome taking into account the FFC effectively (red, labeled by "ECT") for ν_e , $\bar{\nu}_e$ [panels (a)–(c)] and ν_x [panels (d)–(f)] in Models II–IV, respectively.



FIG. 2. The neutrino angular distributions for ν_e , ν_x and their antineutrinos at $r \simeq 44$ km in Model II [panel (a)] and $r \simeq 41$ km in Model IV [panel (b)]. The corresponding ELN angular distributions are shown in panels (c) and (d). The yellow and blue curves are obtained by the ν QKE simulations and the red curves are from the corresponding ECT models. We only plot v_r from-0.5 to 1 in (a) and (c) for better illustration. The overlapping asymptotic curves of $\bar{\nu}_e$ and $\bar{\nu}_x$ are on the top of those of ν_e and ν_x for $v_r > 0.4$ in (b). The ELN distribution after the prompt FFC is shown by the dotted red curve in (d).

attenuation is applied, the transition becomes less sharp and the location is shifted inward to $r \simeq 33$ km. This indicates that the difference is related to the artificial attenuation and the modified ratios between the collisional and FFC rates. Overall, the better agreement between the ECT and ν QKE-H models than in between the ECT and ν QKE-L models suggests that in the ν QKE-L models the larger attenuation of $\mathbf{H}_{\nu\nu}$ may have quenched the FFC rate too much, particularly in Models II- ν QKE-L and IV- ν QKE-L.

To further illustrate the increasing agreement between the ECT and the ν QKE results with less attenuation, we compare the asymptotic neutrino angular distributions obtained at $r \simeq 44$ km in Model II in Fig. 2(a) and $r \simeq$ 41 km in Model IV in Fig. 2(b). In Model II-ECT, the reinforced flavor redistribution scheme leads to nearly equal amounts of electron and heavy lepton flavors in both neutrino and antineutrino sectors in the velocity range $v_r \gtrsim v_c$. In both Models II- ν QKE-L and II- ν QKE-H, slight flavor overconversion is obtained. In Model IV, although FFCs constrain the ELN distribution by prohibiting the reappearance of new crossing, collisions can still affect the angular distribution of each neutrino species and exert feedback. This leads to the gradual change of ELN



FIG. 3. The radial profiles of neutrino number density ratios $n_{\nu}^{\text{FFC}}/n_{\nu}^{\text{NFC}}$ [panel (a)] and the angular distributions at $r \simeq 41$ km [panel (b)] in Model III- ν QKE-H (blue), III-ECT (red), III-ECT- τ (cyan), and III-ECT-b (green). The ECT- τ model takes into account the FFC timescale. The ECT-b model uses a different flavor redistribution scheme from the ECT model. As the curves for Model III-ECT- τ completely overlap with those from the ECT model, they are shifted by -0.005 in (a) and -0.2 in (b), respectively.

distribution and v_c , so that flavor equilibration is reached at $v_r > v_c$ in the asymptotic state, in contrast to the range $v_r < v_c$ indicated by the red dotted curve in Fig. 2(d). Such an effect can be well reproduced in ECT models as illustrated by Figs. 2(b) and 2(d). Once again, the results in the ν QKE-H models clearly agree with those in the ECT models better than the ν QKE-L models.

Alternative effective schemes-Besides the fiducial effective model used above, we have also performed two additional ECT schemes. The first set (labeled by "ECT- τ ") is to involve an FFC timescale to relax the assumption of instantaneous flavor redistribution [52]. The characteristic FFC timescale may be estimated by $\tau = G_F^{-1} |I_+I_-|^{-1/2}$, which is proportional to the geometrical mean of the positive and negative ELNs [29]. At each time step, we compute τ at where the ELN crossings are found and update the neutrino and antineutrino flavor content by ρ + $(\varrho^{(f)} - \varrho) \min(\Delta t/\tau, 1)$ and $\bar{\varrho} + (\bar{\varrho}^{(f)} - \bar{\varrho}) \min(\Delta t/\tau, 1)$, respectively. When $\tau < \Delta t$, it reduces back to the instantaneous redistribution scheme. The results obtained with this method are nearly identical to those with instantaneous flavor redistribution in all models; see e.g., Fig. 3 for Model III. This comparison supports the validity of implementing instantaneous flavor redistribution in hydrodynamic simulations [68–72].

The second set of additional simulations is to use the "box" prescription [49] by setting $P(v_r) = 1 - I_</(2I_>)$ for $v_>$, labeled by "ECT-b," which offers easier implementation but creates discontinuities in neutrino angular distributions [51]. Interestingly, despite the presence of local discontinuities [see, e.g., Fig. 3(b)], the resulting neutrino densities do not differ substantially from those using the power-1/2 prescription, as shown in Fig. 3(a) for Model III. We have also examined that for other prescriptions, which model the FFC outcome by continuous functions in [51], the results are nearly as good as those with the power-1/2 scheme.

Summary and discussions-We have performed the first comparison of the global ν QKE simulations to the corresponding ECT simulations that effectively take into account the outcome of the FFC. Utilizing the scale separation between the FFC and the neutrino collisions as well as advection, we have implemented the analytical prescriptions proposed in [51] as the instantaneous FFC outcome over a radial size much larger than the FFC length scale into global simulation of classical neutrino transport that includes collisions and advection. The ECT scheme takes significantly reduced grid number and time step required by the ν QKE simulations to obtain results that quantitatively agree with those by solving the full νQKE with differences of less than a few percents. The models that we examined in this work cover different cases with number densities dominated by ν_e or $\bar{\nu}_e$ and scenarios where the FFIs exist inside and outside neutrinospheres. In particular, the evolution of neutrino flavor content driven by the collisional feedback when the FFIs are located inside the neutrinospheres can also be robustly modeled by the ECT method.

Moreover, we have found increasing agreement between the ν QKE results with less attenuation in the neutrinoneutrino forward scattering Hamiltonian with the results obtained by the ECT method. Taking into account the characteristic FFC timescale or using different analytical prescriptions to describe the redistribution of neutrinos by FFC also gives rise to similar results in the ECT models. Our results suggest that the flavor redistribution obtained on a coarse-grained level achieved in the FFC timescale in local FFC simulations with periodic boundary conditions can be used as a good proxy in global transport simulations that aim to include the impact of FFCs. Although the models examined are based on a particular SN simulation profile, we fully expect that the method is generally applicable to different SN backgrounds so long as the separation of scales holds at where FFC occurs. Also, due to the local nature of the ECT scheme, its implementation does not add significant computational burden to classical transport. Thus, this work serves as a proof-of-principle study for the upcoming efforts to accurately identify the FFIs and integrate the FFC outcome utilizing machine learning and neural networks into the CCSN and NSM simulations equipped with classical two-moment neutrino transport schemes.

Despite the remarkable success presented above, it remains desired to explore whether flavor swap through the electron lepton zero surface [63,85] or nearquasistationary flavor transformation due to the temporal evolution of the environment [73–75] can be treated effectively as well. If a non-negligible amount of flavor overconversion due to FFC may exist for certain ELN distributions that cannot be described by analytical prescriptions, it remains to be seen whether machine-learning based methods can accurately predict the related FFC outcome. All these as well as the generalization to multidimensional simulations and further development of ECT models for other types of collective oscillations are beyond the scope of this work but will be pursued in upcoming studies.

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